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Vitale Brovarone

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### 1 Ti-bearing minerals: from the ocean floor to subduction and back

2 Inês Pereira<sup>a\*</sup>, Emilie Bruand<sup>a,b</sup>, Christian Nicollet<sup>a</sup>, Kenneth T. Koga<sup>a</sup>, Alberto Vitale
3 Brovarone<sup>c,d</sup>

4 Affiliations

- 5 a Université Clermont Auvergne, CNRS, IRD, OPGC, Laboratoire Magmas et Volcans, F
- 6 63000 Clermont-Ferrand, France
- 7 b Geo-Ocean laboratory, Université Bretagne Occidentale, CNRS, Plouzané, France.
- 8 c Dipartimento di Scienze Biologiche, Geologiche e Ambientali (BiGeA), Alma Mater
- 9 Studiorum Università di Bologna, Piazza di Porta San Donato 1, 40126, Bologna, Italy
- 10 d Sorbonne Université, Muséum National d'Histoire Naturelle, UMR CNRS 7590, IRD,
- 11 Institut de Minéralogie, de Physique des Matériaux et de Cosmochimie, IMPMC, 4 Place
- 12 Jussieu, 75005, Paris, France
- \*corresponding author: <u>ines.pereira@dct.uc.pt;</u> present address: University of Coimbra,
  Department of Earth Sciences and Geosciences Center, Rua Sílvio Lima 3030-790 Coimbra,
  Portugal
- 16

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22 Abstract

Rutile, titanite, and ilmenite are the most common Ti-bearing minerals in metamorphic rocks. Experimental constraints have shown that titanite is stable at low-grade metamorphic conditions, rutile at HP, and ilmenite at HT-LP conditions. Yet, petrological evidence suggests that titanite can also be stable at LT-HP. This implies that both titanite and rutile can be used to develop proxies to track HP metamorphism, which can have interesting applications.

In this study, we have investigated the natural occurrence and chemistry of Ti-bearing minerals in gabbroic rocks that record different degrees of metamorphism, including LP amphibole-bearing gabbros from the ocean floor (Mid-Atlantic and Indian ridge IODP LEGs) and from an obducted ophiolite (Chenaillet) and HP Alpine metagabbros including blueschist and eclogite facies rocks from the Western Alps and Corsica. We have performed detailed petrography and Raman spectroscopy and analysed major and trace elements mineral chemistry using EPMA and LA-ICP-MS.

We found that rutile is stable at low pressure (< 2 kbar) in ocean-floor amphibole-36 bearing gabbros, lower than experimental constraints had previously suggested. Rutile is also 37 found in eclogitic metagabbros from the Western Alps and can be chemically distinguished 38 from LP rutile. Blueschist metagabbros from the Western Alps and eclogitic metabasalts from 39 Corsica have titanite stable instead of rutile. While the titanite to rutile transition is pressure 40 and temperature dependent, we demonstrate how small variations in bulk-rock Ti/Ca and 41 Ca/Al values within the NCKFMASHTO chemical system may shift their stabilities. High-42 pressure titanite from these metamafic rocks exhibits La depletion and low La/Sm<sub>N</sub> values in 43 44 comparison to titanite from amphibolite-facies mafic rocks. La/Sm<sub>N</sub> or Nb together with Yb and V can be used to distinguish HP titanite from titanite formed under other P-T settings. 45

46 These new systematics can be useful in studies using detrital Ti-bearing minerals to probe the

47 HP metamorphic record through time.

48 Keywords: rutile; titanite; P-T stability; trace element-based discrimination; metagabbro

### 49 **1. INTRODUCTION**

The concept of mineral equilibria has allowed petrologists to estimate pressure and 50 temperature conditions seen by metamorphic rocks (Hoschek, 1969; Burt and Ferry, 1982). 51 However, complete mineral equilibria of rocks are typically unattained during geological 52 processes, especially in coarse-grained rocks that often preserve primary textures, and/or peak 53 and retrograde metamorphic assemblages (non-equilibrium). Over the years, these different 54 snapshots of the rock's evolution have been exploited by petrologists, to better constrain its 55 pressure-temperature (P-T) history. Attribution of a mineral to a pre- or a given metamorphic 56 stage is paramount in reconstructing the P-T evolution of a rock, central to the geodynamic 57 interpretation of metamorphic petrology, and especially for tracing ancient subduction 58 histories (Rubatto, 2002; Engi, 2017; Yakymchuk, 2017). 59

The "classical" equilibrium thermobarometry approach requires a thorough chemical 60 characterisation of the different mineral phases that attained equilibrium in cation exchange 61 reactions. Alternative ways to this approach have arisen from the development of single grain 62 thermometers (e.g. Ti-in quartz, Zr-in-titanite, Zr-in-rutile), allowing to determine 63 temperature conditions from rocks with multiple heritages and/or as inclusions within another 64 mineral. So far, single grain methods are only available for temperature but not pressure, and 65 with known limitations related to uncertainties in elemental behaviour under different P-T 66 conditions and buffering assemblages (Kooijman et al., 2012; Cruz-Uribe et al., 2018; 67 68 Kendrick and Indares, 2018; Clark et al., 2019). To constrain pressures by single grains, barometry based on quartz Raman spectra have been developed over the last decade, mainly 69

on quartz in garnet but also on other host-inclusion pairs (*e.g.* Enami *et al.*, 2007; Ashley *et al.*, 2014; Gonzalez, 2019; Cisneros *et al.*, 2020; Gonzalez *et al.*, 2021). Yet, in the absence of
quartz inclusions, we need to rely on grain chemistries for an alternative pressure information
source.

Titanium-bearing minerals are common accessory minerals in mafic rocks and 74 potentially serve as single grain P-T indicators as referred above. The stability of these 75 76 minerals, titanite, ilmenite and rutile, are controlled by different factors: system composition, pressure and temperature (Force, 1991; Liou et al., 1998) and fluid composition and its 77 dissolved ions (Spear, 1981; Force, 1991). Rutile is commonly found at higher metamorphic 78 grades, at upper amphibolite and granulite facies (Goldsmith and Force, 1978; Hart et al., 79 2018), and at high pressures, such as in the eclogite facies or UHP (Liou et al., 1998; Gao et 80 al., 2004; Hart et al., 2016). Rutile is seldom found at lower grades (greenschist facies), and 81 usually associated with fluid alteration processes (Meinhold et al., 2008; Zeh et al., 2018). 82 Titanite, conversely, is mostly found at lower metamorphic grades (greenschist, low 83 amphibolite, blueschist and occasionally eclogite facies conditions; Force, 1991; Vitale 84 Brovarone et al., 2011). Commonly, with increasing metamorphic grade, at upper amphibolite 85 to granulite facies conditions, titanite becomes unstable and ilmenite or rutile replace it 86 (Spear, 1981; Force, 1991; Liou et al., 1998). Ilmenite is favoured by high Fe/Mg bulk-rock 87 compositions (Force, 1991) and medium fO<sub>2</sub> conditions (Spear, 1981; Guo et al., 2017). 88

Experimental constraints on the stability of these three Ti-phases for MORB 90 compositions are shown in Figure 1a. The lack of experimental data below 600 °C for mafic 91 compositions has been partly compensated by natural observations and by solution and 92 thermodynamic models of minerals and mineral equilibria (White *et al.*, 2014; Green *et al.*, 93 2016). Note that the low temperature boundary in Figure 1b is only weakly constrained. 94 Notably, titanite is predicted as more stable than rutile at HP and T <550 °C (Fig. 1b),

although the position of the *titanite-out* reaction changes considerably depending on the 95 96 choice of solution models (here using the most recent by Green et al., 2016 and White et al., 2014). At lower pressure, titanite and ilmenite are the common Ti-phases (Fig. 1). Yet, the 97 occurrence of rutile-titanite in oceanic gabbros has been reported (Batiza and Vanko, 1985; 98 Mével, 1987; Abe, 2011; Harigane et al., 2011; Rizzo et al., 2019). While titanite presence in 99 these gabbros has been generally linked to the cooling of the oceanic crust during ocean floor 100 metamorphism, the presence of rutile, however, is more puzzling, as it may form due to 101 magmatic or metamorphic processes. Altogether this implies that details on their stabilities at 102 different P-T and chemistry remain puzzling. Until now, despite some attempts to develop 103 trace element-based barometers applied to titanite and rutile (e.g. Al-F in titanite – Tropper et 104 al., 2002; Al in rutile - Escudero et al., 2012), no dependency has yet been found (Castelli 105 and Rubatto, 2002; Hoff and Watson, 2018). 106

In this study we focus on the fate of these Ti-phases from the magmatic stage of the 107 oceanic crust through its subduction. We have investigated the phase stability and trace 108 element systematics of rutile and titanite, two common Ti-bearing metamorphic minerals, 109 which are calibrated as geothermometers, to assess their potential as pressure discriminants. 110 Minor and trace element compositions of accessory minerals have been successfully used in 111 the past to discriminate rock types (e.g. Bruand et al., 2016, 2020; Guo et al., 2020; Scibiorski 112 and Cawood, 2022). By investigating the trace element compositions of these Ti-bearing 113 phases at different metamorphic grades, we explore systematics based on pressure variations. 114 Specifically, we have examined a suite of metamafic rocks with MORB-N affinity (gabbros 115 and basalts) from (i) both the modern and obducted oceanic crust, and (ii) from subducted 116 117 oceanic crust metamorphosed at blueschist and eclogite facies and exhumed back to upper crustal levels (Fig. 1a and 2). We ultimately investigate which mechanisms control their 118 stability and test if their chemistries change with metamorphic grade. The main purpose is to 119

provide a novel tool that enables the discrimination of Ti-phases either in or out ofpetrological context.

#### 122 2. SAMPLES AND METHODS

#### 123 **2.1. Sampling and Geological settings**

Samples come from two different settings: (i) high temperature, low pressure (HT-LP) 124 gabbros metamorphosed during slow seafloor spreading, and (ii) low to medium temperature, 125 high pressure (L/MT-HP) gabbros and basalts metamorphosed during convergent processes. 126 127 These samples were selected based on their N-MORB-affinity and due to the presence of Tiphases. HT-LP samples were collected from the Chenaillet Massif, Western Alps (3 samples) 128 and from three different IODP legs, 305 (site U1309D), 209 (site 1271B), both from the 129 Atlantic Ocean Ridge, and 118 (site 735B) in the Indian Ocean Ridge, all with HT 130 amphiboles. L/MT-HP samples were collected from the Western Alps (Queyras, Monviso) 131 and Corsica (San Petrone). More details are given in Figure 2 and Table 1 and the exact 132 sampling locations are given in the Supplementary Data: Electronic Appendix 1. 133

# 134 2.1.1 International Ocean Drilling Program samples: Atlantic and Indian oceans

Samples collected from all the following IODP sites and cores are affected by ocean floor 135 metamorphism during gabbro emplacement and cooling of the oceanic crust. Samples used in 136 this study correspond to gabbro that yield a significant oxide content (> 1 %), and 137 metamorphic amphiboles. Preserved metamorphic grades are variable, with some samples 138 preserving a HT metamorphic mineral assemblage with equivalent temperatures of granulite 139 facies whilst others have been re-equilibrated at amphibolite facies conditions. This allows an 140 assessment of the Ti-bearing phase budget present in the oceanic crust before it undergoes 141 subduction and consequent HP metamorphism. More details about the drilling sites and cores 142 can be found in Supplementary Data: Electronic Appendix 1. 143

#### 145

#### LEG 209, SITE 1271B (0 – 104 m, 15°20'N, 45°00'W)

Site 1271 is located along the track of Faranaut 15°N on the western flank of the Mid-Atlantic 146 rift valley. Roughly, 100 m of core were retrieved, corresponding to dunite, harzburgite, 147 troctolite, gabbro, minor gabbronorite, and a small but significant amount of chromitite cm-148 thick lenses (Shipboard Scientific Party, 2003). Interstitial gabbros in impregnated peridotites, 149 150 olivine gabbros, and troctolites from Site 1271 include a large proportion of high-temperature amphibole. Samples used in this study, 1316, 1318 and 1320, are all metamorphosed gabbros 151 and were retrieved from cores 3R (piece 19, 18 m depth), 5R (piece 3, 28 m depth), and 8R 152 (piece 5, 41 m depth), respectively, all within Unit I, which was classified as dunite/gabbro 153 (Shipboard Scientific Party, 2003). 154

155

5 LEG 305, SITE U1309D (0 – 1645 m, 30°10'N, 42°07'W)

Site U1309 is located on the central dome of Atlantis Massif, 14–15 km west of the median valley axis of the Mid-Atlantic Ridge (MAR), where the seafloor coincides with what is interpreted to be a gently sloping, corrugated detachment fault surface (Blackman, 2006). In this site, oxide gabbro (defined by the presence of >2% modal Fe-Ti oxide minerals) makes up about 7% of the rocks recovered. Sample 1319 is an amphibolitised oxide gabbro and corresponds to unit 353, piece 1, recovered at about 662 m depth (Structural Unit II; Blackman, 2006), defined as the Gabbronorite Zone 2 (Blackman, 2006).

163 LEG 118, SITE 735B (0 – 501 m; 32°43'S, 57°16'E)

Site 735 is located on a shallow platform, on the east rim of the Atlantis II Transform located in the magnetic anomaly equivalent to about 12 Ma (Robinson *et al.*, 1991). This platform, about 9 km long in a north-south direction and 4 km wide, is one of a series of uplifted blocks. A total of 435 m of olivine gabbro, olivine-bearing gabbro, two pyroxene gabbro,

iron-titanium oxide gabbro, troctolite, and microgabbro with rare basalt and trondjhemite was 168 169 recovered from Hole 735B. These rocks have undergone varying degrees of plastic and brittle deformation, and many have well-developed foliations. The sample used in this study (5450) 170 is a metamorphosed oxide gabbro and comes from unit VI, core 85R, section 4, corresponding 171 to piece 1E of the original log and to approximately 475m depth. The unit this sample comes 172 from corresponds to olivine-rich gabbro having frequent layers of troctolite (Robinson et al., 173 1991). Frequent intervals of gabbro are characterized by mylonitic to porphyroclastic textures, 174 the latter corresponding to the collected sample. 175

176 2.1.2. The Liguro-Piemontese domain in the Western Alps and Corsica

The Piemonte-Liguria domain in the Western Alps represents the westernmost part of 177 the Alpine arc and extends to Corsica in the south (Fig. 2 A, B). This domain is interpreted as 178 a remnant of the accretionary wedge, where oceanic sediments and relics of oceanic 179 lithosphere are accumulated (Marthaler and Stampfli, 1989; Deville et al., 1992; Agard et al., 180 2001). In the Piemonte-Liguria zone we studied two areas: (i) a west-east transect going from 181 the obducted ophiolite in the Chenaillet Massif, to the blueschist Queyras/Combin unit 182 towards the eclogitic Viso/Zermatt unit, and (ii) Alpine Corsica in the San Petrone Mount 183 area. Together, these samples allow a comparison between what constituted the pre-Alpine 184 oceanic lithosphere crust (preserved in the Chenaillet ophiolite as the least metamorphosed 185 past of the Western Alps) and oceanic lithosphere units that underwent subduction and 186 metamorphism at different depths during the Alpine orogeny. 187

*Chenaillet Massif.* The Chenaillet amphibole-bearing gabbros were long interpreted as corresponding to obducted amphibole-bearing gabbros recording ocean floor metamorphism (Mevel et al., 1978). They formed in the Piemonte-Liguria basin during the Jurassic, and were so far described as negligibly affected by Alpine metamorphism (Manatschal *et al.*, 2011 and

references therein). Recent work by Corno et al. (2023) describe for the first time the presence 192 193 HT assemblage along with remnant of Alpine HP assemblages (e.g. lawsonite, omphacite) in some of the gabbros and volcanics, questioning the interpretation of previous authors 194 implying the obduction of the Chenaillet massif. HP alpine metamorphism (presence of 195 lawsonite and omphacite) were not observed in the studied gabbros. Recent ages obtained 196 from magmatic zircon in the studied gabbros yield crystallisation ages of around  $161 \pm 0.8$ 197 Ma (Nicollet et al., 2022). In this locality, three amphibole-bearing gabbros have been 198 sampled (Table 1). 199

Queyras and Monviso Massifs. The blueschist-facies Queyras and eclogite-facies 200 Monviso metaophiolite units consist mainly of meter to kilometer-sized metagabbroic lenses 201 included in metasediments (Schistes Lustrés complex; SL) and/or serpentinite with 202 ophicalcite (Ballèvre et al., 1990). Magmatic ages of the Monviso eclogite protoliths have 203 been determined using coupled zircon U-Pb and trace element chemistry to  $163 \pm 2$  Ma 204 (Rubatto and Hermann, 2003). In the Western Alps, the subduction stage started during the 205 late Cretaceous at around 80 Ma (e.g. Rubatto et al., 2011) and lasted until at least 45 Ma 206 (Rubatto and Hermann, 2003). A west-east transect across the Queyras and Monviso units is 207 affected by a HP-LT metamorphic event, which increases towards the East from blueschist (8-208 14 kbar; 350-450°C) to eclogite facies conditions (up to 20 kbar; 450-500°C) (Blake and 209 Moore, 1995; Agard et al., 2001; Schwartz, 2002; Schwartz et al., 2013). The Liguro-210 Piemontese units are partially retrogressed into greenschist facies conditions along an 211 isothermal retrograde path (Caby, 1995), but fast exhumation, probably combined with 212 213 restricted fluid circulation, favoured the preservation of the peak metamorphism assemblages 214 (Rubatto and Hermann, 2001; Angiboust and Glodny, 2020).

Across the Queyras and Monviso, different localities where HP metagabbros occur have been sampled: (i) three lawsonite blueschist from Sommet Bucher (Q1), Vallon de Clausis (Q6) and Tour Réal (TR10); ii) three eclogite from Verne (VE29), Passo Gallarino
(PG) and Lago Superiore (LS). Sampling details can be found in Table 1 and full petrographic
descriptions in section 4.

Alpine Corsica. The schistes lustrés (SL) units in Corsica have undergone lawsonite 220 blueschist- to lawsonite eclogite-facies metamorphism during the Alpine orogeny and 221 suffered variable retrogression to the greenschist facies (Fournier et al., 1991; Vitale 222 223 Brovarone et al., 2011; Vitale Brovarone and Agard, 2013; Vitale Brovarone and Herwartz, 2013). Mafic eclogites found in the SL have N-MORB affinity (Saccani et al., 2008). Ages of 224 gabbros from the oceanic units have been constrained using Sm-Nd isochrons to  $162 \pm 10$  and 225  $159 \pm 15$  Ma (Rampone *et al.*, 2009). HP metamorphism affecting these units has been dated 226 using a Sm-Nd isochron yielding  $84 \pm 5$  Ma (Lahondère and Guerrot, 1997), but more recent 227 studies have clearly demonstrated that HP metamorphism in the SL took place during the Late 228 Eocene (37 – 34 Ma) (Vitale Brovarone and Herwartz, 2013). 229

Two samples were selected for this study that were collected from similar areas to the ones figuring in Vitale Brovarone et al. (2011) (samples OF3158 and OF3310) from the San Petrone Mount in Corsica. These correspond to coexisting lawsonite eclogite and blueschist eclogite rocks equilibrated at the same P-T conditions of  $520 \pm 20$  °C and  $23 \pm 1$  kbar (Vitale Brovarone et al., 2011).

235 2.2 Methodology

236 2.2.1 Whole-rock geochemistry

Eclogite VE29 and some selected blueschist samples (Supplementary Data: Electronic Appendix 1) were crushed using a jaw crusher and pulverised using the sample preparation facilities at the Laboratoire Magmas et Volcans (LMV), Clermont-Ferrand (France) to obtain major element compositions. Bulk rock major elements of these samples were analysed using a Jobin-Yvon ULTIMA Inductively Coupled Plasma-Atomic Emission Spectroscopy at the
LMV, following similar analytical and instrumental procedures to the ones reported in Barette
et al. (2017) and Guitreau et al. (2019). Bulk-rock major and minor element compositions of
IODP samples were compiled from the IODP data repository. Bulk rock available on Corsica
eclogites were compiled from Vitale-Brovarone et al. (2011). All whole-rock data can be
found in the Supplementary Data: Electronic Appendix 1.

247 2.2.2 Scanning Electron Microscopy

Most collected samples were fist-sized and therefore only minerals from polished thin section were analysed during this study. Ti-bearing minerals were identified and imaged to reveal internal zoning or inclusions using Scanning Electron Microscopy (SEM) sessions at the LMV using a JEOL JSM-5910 LV with a tungsten filament and 15 kV accelerating voltages for backscatter electron imaging (BSE).

#### 253 2.2.3. Raman Spectroscopy

To determine TiO<sub>2</sub> polymorphism we used a Renishaw InVia confocal Raman microspectrometer equipped with a 532 nm diode laser (200 mW output power) and a CCD detector and a Leica DM 2500 M optical microscope at the LMV. A 100× objective in high confocality setting (slit aperture of 20  $\mu$ m) was used and the range of power applied was always lower than 1% of the laser energy. Acquisition time was set at 10 s, and 5 accumulations.

260 2.2.4. Electron Probe Microanalyses

Silicate and Ti-mineral phase minerals were analysed for major, minor and occasionally trace elements with a CAMECA SX100 electron probe microanalyser (EPMA) at LMV. Variable analytical conditions were used to target different elements for the different minerals of interest. A detailed instrumental and analytical setup for each session can be found in the Supplementary Data: Electronic Appendix 1. For silicate mineral characterisation the following elements were determined: Si, Ti, Al, Mg, Fe, Mn, Ca, Na, K, and Cr. Data can be found in Supplementary Data: Electronic Appendix 2. For Ti-mineral phases, the following elements were part of different analytical sessions: Si, Ca, Al, P, Ti, Mn, Fe, Cr, F, Zr, Nb, La, Ta, Y, Sc, Sn, V. Khan titanite reference material (Heaman, 2009) was used for quality control monitoring during Ti-mineral phase analyses. Supplementary Data: Electronic Appendix 3.

272 2.2.5. Laser Ablation-Inductively Coupled Plasma Mass Spectrometry

Further trace elements chemistry of titanite, rutile and ilmenite was obtained using laser ablation-inductively coupled plasma mass spectrometry (LA-(HR)ICPMS) using the Resonetics Resolution M-50E 193 nm excimer laser coupled to the Thermo Element XR at LMV daily calibrated and optimised using NIST612.

The spot sizes used ranged between 9 and 27 µm with a repetition rate of 1 to 2 Hz and a 277 fluence of 2.5 to 3 J/cm<sup>2</sup>. Each analysis consisted of 20s background and between 50 and 90s 278 of signal acquisition. More details can be found in the appendix. We used a sample bracketing 279 method using GSE-1G and GSD-1G as primary standards and GSC-1G, GSD-1G or BHVO 280 in the middle of the sequence as secondary standards (Jochum et al., 2005). Khan (Heaman, 281 2009) was also used as a secondary reference material for additional quality control. Details 282 on the isotopes that were measured, data processing and quality control can be found in 283 284 Supplementary Data: Electronic Appendix 1, 3 and 4.

- 285 **3. RESULTS**
- 286 **3.1. Petrography**

287 *3.1.1. Modern oceanic crust amphibole-bearing gabbros* 

Amphibole-bearing gabbros. All samples (1316, 1318, 1320, 1319 and 5450) 288 289 correspond to coarse-grained gabbro, with primary clinopyroxene and plagioclase. Overall, plagioclase is the most abundant mineral phase (40-65%) but is either strongly or slightly 290 altered to secondary plagioclase (albite), quartz and variable chlorite/sericite (Fig.3-A). Relict 291 magmatic clinopyroxene (augite in composition) has reacted with plagioclase to produce 292 brown hornblende-edenite coronas around it (1320, 5450, 1316) together with albite and 293 epidote. In samples 1319, 5450 and 1316, brown amphiboles are then replaced or overgrown 294 by green hornblende and in 1319 also by actinolite + chlorite. Accessory minerals (< 3 % vol) 295 correspond to ilmenite, titanite, rutile and occasionally hematite, mostly associated to 296 amphibole. In all samples, preserved disequilibrium mineral assemblages are found. Due to 297 the variable presence and proportion of brown and green amphiboles in these samples, there is 298 a spectrum from sample 1320, which records granulite facies conditions with a brown 299 300 amphibole found as idioblastic crystals and in equilibrium with plagioclase, to samples 5450, 1318 and 1316 that have suffered partial retrogression to the amphibolite facies, and sample 301 1319 that has completely retrogressed to lower amphibolite/upper greenschist facies 302 conditions. Despite not constituting the most robust estimate, Ti-in-amphibole thermometry 303 (Otten, 1994) applied to two endmembers of these units, samples 1320 and 1319, yield 304 temperatures compatible with metamorphism under the granulite facies (840 to 740 °C) and 305 amphibolite facies (590 °C), respectively (Supplementary Data: Electronic Appendix 2), 306 providing a working frame for metamorphic conditions at which we find the Ti-phases. 307

308 *Ti-bearing mineral assemblage*. Titanite is the most ubiquitous Ti-bearing mineral in 309 all the metagabbro samples. It commonly forms coronas around an earlier Ti-bearing mineral 310 phase (Fig.3-A-F), often ilmenite (Fig.3-C), more rarely rutile (Fig.3-A, B) or around both 311 (Fig.3 D to F), but it can also occur as idioblastic crystals under amphibolite conditions 312 (Fig.3-G,H). In sample 1320, rutile occurs as grains typically between 50 to 100 μm in size,

associated with brown hornblende and within ilmenite-rutile-titanite aggregates (Fig.3-D). In 313 314 samples recording a lower temperature retrogression (sample 1316 and 1319), ilmenite is surrounded by coronas formed by titanite and rutile showing some host crystallography-315 controlled growth (Fig.3-D). In this case, titanite and rutile replace ilmenite during cooling to 316 amphibolite or greenschist facies conditions, illustrating its disequilibrium (Fig.3-E-F), In 317 summary, whilst in cases titanite and rutile form intergrowths in coronas around ilmenite, 318 with rutile showing some host crystallography-controlled growth (Fig.3-D), there are also 319 cases where rutile is found as coarser (20 to 50 µm), sub-spherical inclusions within ilmenite 320 (Fig.3-D-F) or, more rarely, as single grains, up to 100 µm large, directly overgrown by 321 titanite (Fig.3-A,B). While the core of some of these complex grains, formed by either 322 ilmenite or rutile, corresponds to at least granulite conditions, the titanite replacement (+ rutile 323 lamellae) or new crystals growth takes place at lower grade conditions, during cooling to 324 325 amphibolite or greenschist facies conditions. Ilmenite in these metagabbro rocks can either be of magmatic origin or recrystallised during ocean floor metamorphism. In this way, we have 326 three (but continuous) metamorphic stages: i) augite, plagioclase and brown hornblende + 327 ilmenite/rutile, followed by ii) green hornblende, plagioclase +/- rutile, and a iii) actinolite, 328 albite, chlorite + titanite (+/- rutile lamellae). 329

## 330 *3.1.2. Western Alps obducted metagabbros*

These samples were collected from highly heterogeneous outcrops in the Chenaillet Massif, where a coarse-grained texture is preserved, but occasionally amphiboles reach several centimetres, and deformation variably affects these units, from highly sheared to nondeformed.

335 *Amphibole-bearing gabbros.* Three samples from the Chenaillet were examined, two 336 corresponding to a gabbro composition (Ch1 and Ch3) and one interpreted as more dioritic

(ChDi). ChDi primary mineral assemblage has been replaced by a pale green amphibole, 337 338 albite, sericite, epidote and chlorite, and primary ilmenite is overgrown by titanite ± rutile. Fe-Ti oxides make up more than 3 wt% and have no preferred textural association to any major 339 rock forming mineral. The two samples of gabbroic composition are very similar, and the 340 predominant mineral assemblage corresponds to albite (secondary) and amphibole (Fig.4-A), 341 but while Ch1 has preserved relict of clinopyroxene (augite), in Ch3 all clinopyroxene have 342 been replaced by amphibole. There are three different Ca-amphiboles: brown edenite or 343 hornblende, brown-green hornblende and green actinolite (Fig. 4). All amphiboles have less 344 than 0.5 Ti apfu, and they range in  $Al^{iv} = 0.35 - 1.51$ , and  $Na_{M4} = 0.04 - 0.52$  (Supplementary 345 Data: Electronic Appendix 2). Brown amphibole is either arranged as overgrowths on 346 destabilised magmatic clinopyroxene (Fig.4-A) or as large (> 0.5 mm) xenoblastic crystals 347 (Fig.4-B). These represent the main high-temperature mineral assemblage successively 348 replaced during subsequent recrystallisation along their edges by a brown-green and then a 349 green amphibole arranged as a corona texture. The break down of the Ti-bearing amphiboles 350 caused the release of Ti and the growth of the titanite and rutile. Maximum TiO<sub>2</sub> 351 concentrations in Ch1 amphiboles reach 2.62 to 3 wt% while in Ch3 they reach 3.02 to 3.07 352 wt% (Supplementary Data: Electronic Appendix 2), yielding temperatures between 890 and 353 790 °C (Ch1) and 940-930 °C (Ch3; Supplementary Data: Electronic Appendix 1) (Otten, 354 355 1984). These conditions comfortably indicate temperatures typical of granulite facies metamorphism even if the amphibole thermometer may have poor accuracy. A corona of 356 brown-green amphibole develops around these brown amphiboles, which are either 357 hornblende (725 to 600 °C; 0.3 – 1.73 wt% TiO<sub>2</sub>) or actinolite (610-570 °C; 0.5 wt% TiO<sub>2</sub>), 358 359 and to which titanite grains are associated with (Fig.4-A-C). Similarly to oceanic gabbros, a 360 partial preservation of primary magmatic minerals towards granulite and amphibolite facies conditions was likely acquired during cooling of these units at the oceanic ridge. 361

*Ti-bearing mineral assemblage.* Titanite is the stable metamorphic Ti-phase in all the 362 363 amphibole-bearing gabbros, and ilmenite in the metadiorite. Titanite can be found in two types of associations: (i) commonly as small grains ( $< 20 \mu m$ ) at the transition between brown 364 and green hornblende (Fig.4-B-C); or (ii) more rarely as hypidioblastic grains (up to 80 µm in 365 size) associated with actinolite in later veins (Fig.4-G, H). Although the former is 366 homogeneous in BSE, the latter shows strong BSE zonation (Fig.4-H). Ilmenite is more 367 common in the ChDi than in the other metagabbros, where grains are often replaced by rutile-368 titanite association (Fig.4-D-F). In this case, rutile either occurs as coarse grains of up to 25 369 µm or as very fine grains, often showing preferred crystallographic orientation. Both coarse 370 and lamellae rutile are common, easily found in multiple sites in this sample. Rutile lamellae 371 develop at 60%/120% angles with ilmenite (Fig.4-E,F), and they seem to indicate topotactic 372 replacement of ilmenite, such as described by Force et al. (1996), during titanite growth, with 373 374 excess Ti forming rutile. No evident BSE zonation was identified.

375

376 3.1.3. Alpine blueschist and eclogite

Nine samples were collected from the Western Alps and Alpine Corsica localities. These unitsare usually enclosed in HP metapelites and marls.

379 Blueschist (LT, HP)

Metagabbros - QE-1, QE-6, TR10. These blueschist rocks all located in the Queyras valley preserve their original gabbroic texture, where both pyroxene and plagioclase sites can be recognised despite the development of Alpine foliation. The HP mineral assemblage is given by glaucophane, lawsonite, albite and titanite  $\pm$  epidote and secondary minerals present are actinolite  $\pm$  chlorite. Primary magmatic clinopyroxene has been nearly all replaced by glaucophane, although it is still recognisable (Fig. 5-A). The original magmatic plagioclase site is replaced by lawsonite and some epidote and albite aggregates (Fig. 5-A,B). TR10
sample is richer in blue-amphibole, with compositions ranging from ferroglaucophane to
riebeckite (Supplementary Data: Electronic Appendix 1 and 2), whereas blue amphibole in
QE-1 and QE-6 is glaucophane. These amphiboles are variably replaced by actinolite (Fig.5B).

Ti-bearing mineral assemblage. In these metagabbros, the stable Ti-bearing mineral is 391 titanite. Titanite is found as large and elongated stripes of xenoblastic crystals (100's µm) 392 between the clinopyroxene and plagioclase sites, and it also occurs as coronas, probably 393 formed at the expense of leached Ca and Ti from former pyroxene and Ca-rich plagioclase 394 (Fig. 5-A). Titanite also occurs together with glaucophane and lawsonite  $\pm$  albite. 395 Occasionally, it is possible to identify tiny rutile inclusions ( $< 10 \mu m$  wide) in the core of 396 larger (> 400 µm) titanite grains (Fig. 5-C-F), arranged as spotted texture within titanite. In 397 places where only smaller titanite grains are found, fully recrystallised at blueschist 398 conditions, rutile was absent. No evidence of rutile being in equilibrium with other blueschist 399 facies minerals was observed. Ilmenite is also absent and no visible BSE zoning in titanite 400 was identified. 401

402 Eclogite-facies rocks (M/LT, HP)

403 *Metagabbros* - *VE*/29, *LS and PG samples*. PG is a metagabbro that still preserves its primary 404 mineralogy, with about 50:50 proportion of clinopyroxene and plagioclase sites. A primary 405 clinopyroxene (augite) is found, with very light brown to greenish colour. The plagioclase site 406 is replaced by quartz, chlorite, zoisite and phengite. Rutile and titanite are accessories and 407 form elongated bands (Fig. 6-A). Amphibole occurs as very small ( $<20 \mu m$ ) idioblastic grains 408 associated to zoisite and epidote. This seems to occur as thin veins/films infiltrating the 409 metagabbro. Garnet is not observed. This sample have mostly preserved its magmatic texture and although no omphacite was clearly observed, this rock was sampled from the classical
locality in Passo Gallarino where other eclogites have been previously sampled. The exact
location is given in the Supplementary Data: Electronic Appendix 1.

413 VE-29 was collected in Verne locality and is a metagabbro with а granoporphyroblastic texture composed of omphacite, phengite, garnet, quartz, and rutile  $\pm$ 414 ilmenite ± epidote from the Viso unit (Groppo and Castelli, 2010; Fig. 6-B). Garnet 415 416 porphyroblasts (up to 500 µm) are almandine-rich and exhibit an increase of Fe and Mg contents towards the rim, similarly to Groppo and Castelli (2010). Garnet includes a 417 significant number of rutile inclusions, many of which are overgrown by a thin rim ( $<2 \mu m$ ) 418 of ilmenite. Paragonite and epidote are also frequently found as mineral inclusions. 419 Omphacite is occasionally associated with phengite. These can be altered to chlorite, and 420 some chlorite is also found in garnet fractures. 421

LS was collected in Lago Superiore area and corresponds to a strongly foliated eclogite with granoblastic texture, composed by omphacite, garnet and rutile (Fig. 6-C). Garnet grains are hypidioblastic but of small size (< 500  $\mu$ m) and frequently strongly zoned, with Ca-richer almandine cores and Mg-richer almandine rims (cores: Alm<sub>55</sub>Grs<sub>33</sub>Sps<sub>07</sub>Prp<sub>05</sub>; rims: Alm<sub>67</sub>Prp<sub>19</sub>Grs<sub>13</sub>Sps<sub>01</sub>; Angiboust et al., 2012). In similar eclogite retrieved from Lago Superiore (Angiboust et al., 2012), three omphacite generations, with variable compositions, could be identified.

429 *Ti-bearing mineral assemblage.* In all metagabbroic rocks (PG, VE29, LS), rutile is 430 the stable Ti-phase. Rutile grains can be large (200-300  $\mu$ m), sub-spherical and with brown-431 amber colour in plane polarised light. In PG, rutile can be found as coarse, sub-spherical 432 grains in aggregates or individual crystals dispersed in the matrix. Some of these are 433 overgrown by titanite (Fig. 6-A) that was probably formed during decompression from

eclogitic conditions (Schwartz et al., 2000). Rutile from VE-29 ranges in size from 20 to 200 434 435 µm, hypidioblastic individual grains that can be found included in garnet or dispersed in the matrix between omphacite grains. Rutile always displays Fe-exsolution along its cleavage. 436 Less frequently, ilmenite overgrowths are seen. Larger rutile grains which are mostly hosted 437 in the garnet porphyroblasts (Fig. 6-B) often preserve mineral inclusions, such as omphacite, 438 epidote and more rarely apatite. Rutile grains in LS are of variable size (> 300 to  $< 100 \mu m$ ) 439 and frequently occur as bands of rutile aggregates (Fig. 6-C), aligned along the main foliation, 440 together with garnet and omphacite. These occur as aggregates of 50-200 µm, idio- to 441 hypidioblastic grains, forming stripes/bands. These rutile grains have a brown-honey colour. 442 They also occur as inclusions within garnet, but more frequently found along grain 443 boundaries. Ilmenite exsolution bands or overgrowth around rutile are frequent, although 444 these tend to be thin (< 2  $\mu$ m width; Fig. 6-D). Rutile grains from these metagabbros show no 445 446 BSE zoning.

Metabasalts - C16 sample. This sample is equivalent to sample OF3310 in Vitale 447 Brovarone et al. (2011). It corresponds to a foliated lawsonite-bearing "blueschist" eclogite 448 (Fig. 6-E) with an associated omphacite- and garnet-bearing vein (C16v; Fig. 6-F). The 449 sample paragenesis is glaucophane, lawsonite, quartz, albite, titanite, and Fe-Ti oxides (Fig. 450 6-E). The vein, which is ca. 1 cm wide; contains omphacite, garnet, lawsonite, glaucophane, 451 albite and titanite (Fig. 6-F). Both parts reached the same P-T conditions, but paragenesis is 452 dependent on the bulk composition of the basalt and the vein (Vitale Brovarone et al., 2011). 453 P-T conditions are constrained by the lawsonite-blueschist eclogite from Vitale Brovarone et 454 al. (2011) yielding  $520 \pm 20$  °C and  $23 \pm 1$  kbar. *C147 sample*. This sample is equivalent to 455 456 OF3158 in Vitale Brovarone et al. (2011) and corresponds to a lawsonite-bearing eclogite (Fig. 6-G), with occasionally preserved magmatic textures. The HP paragenesis is garnet, 457 omphacite, lawsonite, glaucophane, titanite and late Fe-Ti oxides. Garnet can be idio- to 458

459 hypidioblastic, usually small in grain size (<500  $\mu$ m). P-T conditions are similar to those of 460 C16 (Table 1).

Ti-bearing mineral assemblage. In C16, only titanite is found as the main Ti-host 461 462 phase. Titanite grains are abundant, hypidioblastic, but small ( $<20 \mu m$ ), and become more frequent and larger in the vein that crosscuts the basalt (Fig. 6-F). In C147, titanite is the only 463 Ti-mineral phase, occurring as small grain size idioblastic grains (< 20  $\mu$ m) and abundant 464 465 (Fig. 6-H). Titanite grains show no BSE zoning. Hematite is found as a product of sulphide oxidation. Titanite grains in C147 are in clear equilibrium with the mineral assemblage at 466 lawsonite eclogite facies conditions. This contrasts with the Monviso eclogites, where rutile is 467 468 the HP stable Ti-phase and occasionally titanite is only found as a retrograde phase (PG).

### 469 **3.2.** TiO<sub>2</sub> polymorphism

470 IODP and Chenaillet samples yielding what was recognised petrographically as rutile were 471 subjected to an inspection using Raman spectroscopy to confirm the polymorphism. Grains 472 from 4 IODP samples (1316, 1319, 1320, and 5450) and from 3 Chenaillet samples (Ch1, Ch3 473 and ChDi) were analysed. The Raman spectrum for each of these measurements was 474 compared with spectra of rutile, anatase and brookite from the RRUF database (Figs. 7-9 in 475 Supplementary Data: Electronic Appendix 1). Without exception, for both coarser grains and 476 finer lamellae, rutile was always the TiO<sub>2</sub> mineral identified.

- 477 A summary of P-T conditions associated with stable Ti-phase in each sample, can be 478 found in Figure 7. The larger T range in the IODP and Chenaillet samples is related to the 479 uncertainty in the conditions of Ti-phase growth in those rocks.
- 480 **3.3. Ti-phase mineral chemistry**
- 481 3.3.1. *Titanite*
- 482 *Major and minor elements.*

Titanite from the Atlantic IODP samples (1316, 1319, 1320) shows lower  $X_{Al}$ , between 0.026 and 0.045, and F ( $F_{pfu} < 0.003$ ), with respect to those of Chenaillet titanite, that yield  $X_{Al}$ between 0.066 and 0.077, and  $F_{pfu}$  between 0.003 and 0.007. On the other hand, titanite grains retrieved from the Indian IODP (5450) sample exhibit higher values, with  $X_{Al}$  of about 0.106 and  $F_{pfu}$  of about 0.026. Titanite formed at higher pressure, from the blueschist and eclogite metagabbros, are homogeneous with  $X_{Al}$  between 0.022 and 0.033 and low F concentrations with  $F_{pfu} < 0.005$  (Fig. 8).

In a X<sub>Al</sub> vs F<sub>pfu</sub> diagram, published data of titanite chemistry from mafic eclogite, 490 carbonate-rich eclogite, from UHP, HT calcsilicate, upper amphibolite calcsilicate, high-491 temperature marbles, felsic and mafic amphibolite rocks are also reported. The upper 492 amphibolite calcsilicate rocks and the marbles present the highest  $X_{Al}$  and  $F_{pfu}$ . These are 493 followed by the felsic upper amphibolite units and UHP-upper amphibolite calcsilicate 494 titanite. Our data plot at slightly lower F pfu than the literature dataset reported in Figure 8. 495 LT-HP metamorphic titanite show the lowest X<sub>Al</sub> and can be separated from MT-LP 496 metamorphic titanite in this dataset based on X<sub>Al</sub> but not F pfu. It should be noted that titanite 497 found in the granulite oceanic samples show the lowest F values (Fig. 8). The lack of mineral 498 chemistry data from more mafic protolith titanite prevents a more robust comparison. 499

## 500 High Field Strength Elements, V and Sn.

Titanite grains from the IODP amphibole-bearing gabbros yield average Nb/Ta values around 16.8  $\pm$  9 (Fig. 9a). This contrasts to the extreme, and variable Nb/Ta ratios in titanite grains from the pre-Alpine Chenaillet ophiolite (Ch3; Nb/Ta 16.2 – 208). The Chenaillet ophiolite seems to have been strongly affected by an ocean-floor, late-stage fluid circulation (*e.g.* Nicollet et al., 2022). Some titanite found in thin veins show increased Nb/Ta values in comparison to other titanite grains found in amphibole coronas (average Nb/Ta 128 vs 47, respectively). However, even corona-related titanite in the Chenaillet yields higher Nb/Ta values than titanite in IODP amphibole-bearing gabbros (Fig. 9a). Unlike Nb/Ta, Zr/Hf values are less variable across all HT samples (Fig. 8a). For instance, Chenaillet titanite yields average Zr/Hf of ca.  $29 \pm 11$ , in good agreement with Zr/Hf from IODP titanite ( $28 \pm 15$ ). Finally, blueschist (QE-1, QE-6, TR10, C16) titanite grains yield average Nb/Ta 17.1  $\pm 13$  (6 - 67) and Zr/Hf 19.8  $\pm 9$  (9 - 45). Overall, the majority of titanite grains have average chondritic Nb/Ta but lower than chondritic Zr/Hf values (Fig. 9a).

Figure 9b depicts our titanite data in a V vs Sn diagram (Xiao et al., 2020). The 514 majority of titanite grains yields Sn content < 40 ppm, irrespective of metamorphic grade, and 515 only a few titanite crystals (4 grains) show higher Sn content (40-100 ppm). Vanadium 516 compositions of all titanite are more variable and can go as high as 2500 ppm (Fig. 9b). 517 Titanite grains from MT-LP conditions show a wide range of Sn compositions: i) IODP1320 518 sample yield the lowest Sn and V values (< 10 and < 500 ppm, respectively), while ii) 519 IODP1319 and Ch3 yield higher Sn content (10 - 100 ppm) but similar V chemistry. In the 520 Queyras blueschist, Sn is within uncertainty of granulite-facies titanite grains but have 521 variable and higher V contents (up to 1500-2500 ppm). This is clearly at odds from titanite 522 grains from Corsica (LT eclogite) that yield narrower V (< 500 ppm) and Sn (10-30 ppm) 523 compositions (Fig. 9b). 524

*Yttrium, Rare Earth Elements and Th/U ratios.* In the metagabbros, titanite grains
yield highly variable Y and REE compositions, from more enriched to depleted, from strongly
fractionated to unfractionated (Fig. 10).

Titanite can incorporate significant amounts of Y (Scibiorski et al., 2019). LP titanite yields two groups based on Y (Supplementary Data: Electronic Appendix 3), one yielding a narrow concentration range (sample 1320: 1 – 252 ppm; and from Chenaillet Massif Ch3: 5 – 531 222 ppm), while the other shows more elevated and variable Y contents (samples 1316 and
532 1319: 152 to 4095 ppm). HP titanite (QE-6, TR10, and C16) range between 70 and 500 ppm
533 Y content.

Titanite grains that form under LP conditions (Ch3 and IODP1319 amphibolites) show 534 a wide range of REE contents ( $\Sigma REE = 10 - 3350 \text{ ppm}$ ) and flat REE patterns (La/Yb<sub>N</sub> = 0.02) 535 -1.67)(Fig. 10). Eu anomalies on these units are highly variable (Eu/Eu\* = 0.22 -29.5). 536 Titanite grains from IODP1320 sample have a similar range of REE contents ( $\Sigma REE = 12 - 12$ ) 537 2006 ppm), but show different chondrite-normalised REE patterns (Fig. 10), with LREE 538 enriched patterns (La/Yb<sub>N</sub> = 11 - 113), (La/Sm<sub>N</sub> = 5 - 93) and clear positive Eu anomalies 539 (Eu/Eu\* = 4.7 - 314). Th/U ratios from all these LP rocks are commonly < 1.1 except for one 540 outlier. They show either small scatter within sample (Ch3 and 1319 Th/U =  $0.35 \pm 0.1$ ) or 541 some variation (sample 1320; Th/U = 0.08 - 1.06) 542

Titanite grains retrieved from HP metagabbros and metabasalts show low REE contents ( $\sum REE = 32 - 465$  ppm), of which  $\sum LREE$  represent 4 to 210 ppm. Blueschist rocks from Queyras (QE-1, QE-6, TR10) show consistent depleted titanite LREE pattern (La/Sm<sub>N</sub> = 0.02 - 0.7), similarly to eclogite rocks from Corsica (La/Sm<sub>N</sub> = 0.02 - 0.3). Eu anomalies are non-existent to only slightly positive (Eu/Eu\* = 0.84 - 1.8). Th/U ratios show small scatter within the HP samples, commonly < 0.4. These features are summarised in Table 3.

549 3.3.2. Rutile

550 Minor and trace element composition of rutile grains from two IODP gabbros, from eclogite 551 and as inclusions within titanite from blueschist rocks are reported in Figure 11. Chromium 552 and Nb contents are generally low (< 600 ppm), with only some rutile grains exhibiting more 553 elevated Nb (800 - 1500 ppm) and Cr (800 - 1400 ppm) (Fig. 11a; IODP gabbros). In this 554 way, almost all rutile plot in the *metamafic field* defined by Meinhold et al. (2008), and only a 555 few IODP grains plot within the *metafelsic field*. In the diagram Zr/Hf vs Nb/Ta (Fig. 11b), the majority of the analyses display Zr/Hf (20 - 40) and Nb/Ta (12 - 30) values in the *metamorphic field* defined by Pereira et al. (2019). A limited number of rutile grains display lower ratios, typical of rutile precipitated from hydrothermal-related fluids (Fig. 11b), that plot within the metamafic in Figure 11a.

A ternary diagram constructed using V+Cr, Sc and Nb as endmembers to explore the 560 variance in our dataset (as determined by a PCA analysis) is presented in Figure 11c. Rutile 561 grains from eclogite plot in a field closer to higher V and Cr, unlike rutile grains that formed 562 under LP conditions (IODP gabbros or inclusions in blueschist titanite; Fig. 11c). The latter 563 plots closer to Sc, while the former has variable Sc-Nb. In Figure 10d we present the sum of 564 the HFSE ratios against the Nb/V ratio. Rutile in eclogite display high  $\Sigma$ Zr/Hf, Nb/Ta (>25) 565 and low Nb/V (<0.3), while low pressure granulite to upper amphibolite rutile show higher 566 Nb/V (>0.3). Using previously defined fields shown in Figure 11b, rutile grains with low 567  $\Sigma$ Zr/Hf, Nb/Ta (<25) are likely related to fluid precipitation processes (Pereira et al., 2019). 568

As shown in section 3.1, ilmenite in amphibole-bearing gabbros (IODP) often display partial replacement by titanite and rutile. These are also the samples where rutile chemistry is less homogeneous. In these composite grains, rutile is found such as the grain depicted in Figure 12, either as very fine lamellae, frequently displaying 120%60° intersections, or as coarser grains, mostly frequently rimmed by titanite (example spot 4 and 9\* in Fig 12a), but also just by ilmenite (example spot 7 in Fig 12a).

A diagram of Cr against Zr values is shown in Figure 12b, where two distinct rutile groups are highlighted: (i) grains that are low in Cr (< 600 ppm) and low in Zr (< 300 ppm), and (ii) grains with Cr content above 600 ppm with higher Zr content (500-2500 ppm). The most enriched analysis (7 in Fig. 11) corresponds to a rutile grain surrounded by ilmenite as an inclusion (Fig 12a). The analyses with the lowest Zr, Cr, and V are fine-grained rutile, 581 Yet, the relation between the enrichment in Cr-Nb-Zr and the rutile textural position within

the composite grain (lamellae or inclusion) is not perfectly systematic.

583 3.3.3. *Ilmenite* 

The chemical composition of ilmenite was investigated using EPMA and LA-ICP-MS 584 focusing on two samples from the IODP drill cores, 1320 and 5450, which contained the 585 largest ilmenite grains. These ilmenite grains yield variable Fe<sub>2</sub>O<sub>3</sub> (4.7 and 17.6 wt%) and 586 MgO (12.6 and 4.7 wt%) compositions, similar MnO (0.22 and 0.25 wt%) and TiO<sub>2</sub> (3.3 and 587 4.9 wt%) and overall low Al<sub>2</sub>O<sub>3</sub> (120 to 365 ppm). In sample 1320 (higher bulk Al and Mg), 588 ilmenite incorporates more Al (341 ppm) and Mg (4.1 wt%) than in sample 5450 (Al 134 ppm 589 and Mg 0.5 wt%). A small variation in WR MnO abundance ( $\Delta \sim 0.03$  wt%) seems to reflect 590 highly on Mn incorporation into ilmenite (4 369 to 8 289 ppm; Supplementary data: 591 592 Electronic Appendix 2).

To investigate WR dependency of trace elements in ilmenite, we use Al against other 593 trace elements (Mn yields similar results; Fig 13). Vanadium and Zr inversely correlate with 594 Al, while Cr correlates positively. This seems to imply that at higher  $Al_2O_3$  compositions in a 595 gabbro, ilmenite incorporates less V and Zr (Fig. 13). Conversely, more Cr will be 596 597 incorporated by ilmenite in a more enriched Al<sub>2</sub>O<sub>3</sub> composition. Incorporation of Yb, Nb and Ta by ilmenite does not correlate with Al (nor Mn). Particularly for one sample (IODP1320; 598 Fig. 13) there is either clustering (Yb and Ta) or scatter (Nb) for similar Al values in ilmenite 599 and for the same bulk  $Al_2O_3$ . This can indicate the presence of two distinct ilmenite 600 populations. 601

#### 602 **4. DISCUSSION**

#### 603 **4.1 Ti-phase stability: from oceanisation to subduction**

#### 604 4.1.1 *Ti-minerals in ocean floor amphibole-bearing gabbros (LP-MHT)*

Our data show that ilmenite, titanite, and rutile are present at very low-pressure, and moderate 605 to high temperature (amphibolite to granulite) conditions in the oceanic crust. Their formation 606 history can be deduced from their petrographic textures and chemical signatures. Ilmenite can 607 be a product of magmatic and/or later metamorphic recrystallisation, as it commonly forms 608 the core of ilmenite-titanite composite grains. Fine rutile-titanite lamellae showing 120°/60° 609 angles indicate an exsolution of excess Ti during titanite overgrowth on ilmenite at lower 610 grade. Rutile found as coarser grains (10-40 µm; e.g. sample IODP1316) is more puzzling. At 611 times, coarse rutile is found with no ilmenite surrounded by a titanite corona (e.g. IODP1316), 612 while other rutile grains are found surrounded/intergrown by/with ilmenite and then by a 613 titanite corona (IODP1320; Fig. 3). The former is suggestive of an earlier origin of rutile in 614 the gabbro, either HT metamorphic or a late-stage magmatic crystallisation related to 615 fluctuations in the magma fO<sub>2</sub> (Toplis and Carroll, 1995) enhanced by fluid/melt percolation 616 (Tribuzio et al., 2000). While texturally these could be envisaged as magmatic, this would 617 require magmas with elevated TiO<sub>2</sub> (7-10 wt%; Ryerson and Watson, 1987), such as in more 618 fractionated magmas in the oceanic crust (Whitaker et al., 2007). The second is more difficult 619 to interpret, as rutile can be potentially coeval with ilmenite recrystallisation or interpreted as 620 an earlier phase. Rutile (lamellae, coarse grain) found in IODP samples, can also be 621 tentatively distinguished by their chemistry: one group exhibits slightly higher Nb-Zr-Cr 622 compositions, potentially formed at higher temperatures than the other group that yields lower 623 624 Nb-Zr-Cr, corresponding strictly to the fine lamellae formed during titanite-rutile replacement of ilmenite (Fig.12). This replacement process generally takes place at amphibolite-facies 625

Titanite in the IODP and Chenaillet metagabbros is texturally associated with the breakdown of brown to green amphibole, or occurs as replacement textures after ilmenite, under amphibolite facies conditions. A rare third type of titanite exhibiting strong chemical zoning (Fig. 4-H), formed at a later stage during fluid percolation associated with actinolite at greenschist facies cooling conditions, is well documented in Ch1 – Chenaillet ophiolite (Nicollet et al., 2022).

In summary, Ti-phases in oceanic gabbros evolve from magmatic ilmenite at HT conditions to low temperature titanite by hydrothermal metamorphism during cooling, while rutile potentially crystallise during early and late ocean floor metamorphism due to local or global Ti excess.

638 4.1.2 Ti-phase stabilities in the subducting slab

As the oceanic crust undergoes subduction, new minerals stabilise to higher pressures and 639 temperatures. Here, metagabbros from the Queyras-Monviso and Corsica preserve mineral 640 assemblages equilibrated at blueschist and eclogite facies conditions. Overall, titanite appears 641 stable at LT-HP, except in eclogite samples from Monviso. Typically, titanite grains are found 642 in equilibrium with lawsonite in the Queyras blueschist and Corsica eclogite. They have small 643 grain sizes and are idioblastic in the metabasalt (Fig. 6), while often xenoblastic (Fig. 5) or in 644 small aggregates (Fig. 5) in the metagabbros. Occasionally in the blueschist from Queyras 645 (metagabbros), titanite shows porous texture, with frequent but very small rutile grains. We 646 interpret these as inclusions, as they only occur in the cores of larger titanite and never where 647 titanite grains are fine-grained, and thus better equilibrated. These inclusions are often 648 lamellar, with the preserved  $120^{\circ}/60^{\circ}$  angles. These textures are very similar to those found in 649 the oceanic metagabbros. No ilmenite is found in these titanite grains. These textures are 650

suggestive of rutile being a relic phase (such as magmatic clinopyroxene in the same rocks)
preserved within the recrystallised titanite grains, as testimonies of the ocean-floor
metamorphic evolution.

Rutile is stable in Monviso eclogite and titanite is absent, except for retrograde titanite rimming rutile in the Passo Gallarino eclogite. At similar P-T conditions, titanite is found stable together with lawsonite and omphacite in Corsica samples and rutile is found stable with garnet, omphacite and zoisite in Monviso. A study has argued that titanite reacts to form rutile tied with lawsonite disappearance to form epidote (Frost *et al.*, 2001), which potentially could explain the differences between these eclogite samples.

Lawsonite and titanite stabilities. The Corsican lawsonite eclogite is one of the few 660 occurrences reported in the literature where titanite is found in equilibrium at eclogite facies 661 (e.g. Vitale-Brovarone et al., 2011; Sassi et al., 2000; Ng et al., 2016). Frost et al. (2001) 662 explains that the dehydration reaction of lawsonite producing zoisite and quartz leads to 663 664 titanite breakdown and rutile formation, as zoisite incorporates more Ca than lawsonite. In detail, this reaction is sensitive to water activity (Clarke et al., 2006), bulk composition (Davis 665 and Whitney, 2006, SiO<sub>2</sub>-Al<sub>2</sub>O<sub>3</sub>) and CO<sub>2</sub> in fluids (Castelli and Rubatto, 2002). Previous 666 authors showed that this reaction yields a positive Clapeyron slope, taking place at higher 667 pressure with increased temperature, until about 500 to 600 °C, and then it becomes mainly 668 temperature-dependent (Clarke et al., 2006; Whitney and Davis, 2006). Perhaps as a 669 consequence, there have been few reported occurrences of rutile stable with lawsonite under 670 blueschist or eclogite conditions (e.g. Bianco et al., 2019 in the Elba Island – Italy, Maruyama 671 672 et al., 1998 in the Franciscan Complex - USA, and Groppo and Castelli, 2010 in Monviso -673 W. Alps). This implies that while the net-reactions might influence what components are available to stabilise titanite or rutile, rutile can appear while lawsonite is yet stable, and 674

therefore the transition between HP titanite to rutile is an interplay with other factors such as major oxides,  $H_2O$ , and/or  $CO_2$ .

#### 677 4.1.3 The effects of MORB whole rock composition on titanite and rutile stabilities

It has been well established that for rocks with MORB composition, rutile is more stable at 678 high pressure conditions, while titanite and ilmenite at lower pressures, with ilmenite more 679 stable than titanite at higher temperatures (e.g. Liou et al., 1998). Yet, due to natural 680 exceptions to these phase relationships, in the following we explore other factors that play a 681 682 role in their stability, with a special focus on bulk rock chemistry. The stability of rutile, for example, is mostly influenced by the bulk TiO<sub>2</sub> concentration, whereas titanite is more 683 sensitive to variations in overall bulk geochemistry (Frost et al., 2001; Angiboust and Harlov, 684 2017). In Ca-rich protoliths, the following reactions maximise titanite stability: rutile + 685 grossular + diopside  $\leftrightarrow$  pyrope + titanite (1) and rutile + quartz + calcite  $\leftrightarrow$  titanite + CO<sub>2</sub>(2) 686 (Spear, 1993). The presence of carbonate minerals favour titanite (Cave et al., 2015). The 687 water activity at HP is also an important parameter, as H<sub>2</sub>O will favour hydrous minerals and 688 thus titanite as indicated by the following reaction: lawsonite + titanite  $\leftrightarrow$  zoisite + rutile + 689 quartz +  $H_2O$  (3) (Frost *et al.*, 2001). Indeed, the Corsican eclogite is richer in  $H_2O$  than the 690 Monviso eclogite (Supplementary Data: Appendix 1), where one yields titanite and the other 691 rutile. Yet, as also seen in reaction (2), CO<sub>2</sub>-rich fluids favour rutile instead of titanite. It has 692 also been shown that the F-Al coupled substitutions seems to stabilise titanite above its 693 commonly known stability conditions, more so in Ca-rich protoliths (marbles), at HT or HP 694 695 (Castelli and Rubatto, 2002; Tropper et al., 2002; Fig. 8). While titanite tends to disappear with increasing SiO<sub>2</sub> (Frost *et al.*, 2001), this is mostly linked to the concurring decrease of 696 CaO. Thus, it has long been established that the Ca/Al ratio is key in the stabilisation of 697 698 titanite (Frost et al., 2001). Reflecting on the considerations above, the presence or absence of Ti-phases are shown in a plot comparing Ti/Ca and Ca/Al bulk rock values, using data from 699

our LT-HP samples combined with Syros HP rocks, and average MORB published data (Fig. 700 701 14). Ti/Ca values are expected to influence rutile stability while Ca/Al should mostly influence titanite stability as it competes with plagioclase, lawsonite or epidote. With 702 increasing Ca/Al, even at elevated Ti/Ca, we expect titanite to be stable, leading to a field 703 where both rutile and titanite should be stable (Ca/Al values above ~ 0.85). It should be noted 704 that the compositional range of this set of rocks is limited: bulk SiO<sub>2</sub> (46-51 wt%) and CaO 705 (9-14 wt%), and TiO<sub>2</sub> ranges from 0.25 to 3.17 wt% (Supplementary Data: Electronic 706 Appendix 1 -Table 5) and that no calcite was observed or reported. 707

- 708 Several conclusions can be drawn from the figure:
- Blueschist and eclogite rocks with stable titanite yield variable Ca/Al (commonly higher than 0.75) while Ti/Ca is systematically below 0.15. On the contrary, blueschist and eclogite with stable rutile exhibit Ti/Ca values above 0.15. These data seem to support the major control of reaction (1) on rutile and titanite stability at LT-HP conditions. This is consistent with data on blueschist and eclogite from Syros (Marschall 2005) for which rutile-bearing HP units plot in the Ti/Ca > 0.15 field.
- Unaltered MORB values yield relatively low Ti/Ca values (0.05 0.15) and variable but 715 high Ca/Al that range from 0.8 to 1.2 (Fig. 14; Klein, 2004). These compositions at low 716 geothermal gradients favour titanite over rutile growth. However, a relative increase of Ti 717 718 and Fe in respect to the other major elements is expected for the altered oceanic crust metamorphosed by either hydrothermal or seawater fluid interactions (Staudigel and Hart, 719 1983). Al-saponite type alteration of Fe-bearing minerals (e.g., oxides and 720 721 clinopyroxenes) and albitisation of primary plagioclase can also lead to considerable Caloss (Alt et al., 1992), and thus alteration of MORB rocks are likely to increase Ti/Ca and 722 variably decrease Ca/Al values. 723

#### 724 **4.2.** Controls on Ti-bearing mineral chemistry

4.2.1. Ilmenite trace element compositions: bulk-rock effect

Our limited data show that ilmenite major and minor element compositions are strongly dependent on bulk-rock composition. The binary plots of Al against multiple trace element contents in ilmenite illustrate how Cr seems to be positively correlated with Al, which in turn correlates with the  $Al_2O_3$  content of the metagabbros (Fig. 13). Niobium and Ta showed no significant correlation with Al, indicating the presence of two chemically distinct ilmenite populations in each sample. This implies that other factors are more relevant in the intake of Nb and Ta by ilmenite, such as competition with rutile (*e.g.*, sample 1320).

4.2.2. Titanite trace element compositions: the role of fluids and HREE-bearing minerals

Titanite from LP settings. Titanite from our dataset show Nb/Ta values that are within 734 uncertainty of the chondritic value (Fig. 9a; McDonough and Sun, 1995), except a few titanite 735 from the Chenaillet Massif amphibolite metagabbro that yield Nb/Ta > 100 (Fig. 9a). Titanite 736 Zr/Hf values, on the other hand, are persistently subchondritic. The high-field strength 737 elements (HFSE) pairs, Zr and Hf, and Nb and Ta have a degree of geochemical affinity that 738 prevents their fractionation under most geological conditions. However, some fractionation 739 between each pair in metamorphic titanite has previously been reported (Chen and Zheng, 740 2015). Cherniak (2015) has constrained the diffusivities of these elements in titanite and her 741 data suggest that unlike in rutile, diffusive fractionation should not be a likely mechanism to 742 explain variable titanite Nb/Ta ratios. Therefore, fractionation between these elements in 743 titanite is promoted by other processes such as growth in the presence of external fluids or 744 745 after the break down of minerals with variable compatibility with these elements. In this way, we interpret these higher to extreme Nb/Ta values found in the Chenaillet Massif to be 746 derived from external fluids related with differentiation of the lower oceanic crust, during 747 748 cooling. Additionally, some of the Chenaillet Massif titanite grains (Ch3) can also yield

higher Sn (> 60 ppm; Fig. 9b). Elements like Sn and Mo are usually associated with titanite
growth from hydrothermal fluids (such as in Fu et al., 2016). This may support the existence
of externally derived late-stage fluids during titanite growth. These data are compatible with
the late-stage texture of titanite in these rocks, associated with brown to green amphibole
breakdown and alteration of former ilmenite grains.

754 Y-REE behaviour. Titanite mineral chemistry has been heavily addressed over the years, as its chemistry seems to be a faithful recorder of petrogenetic processes (Bruand et al., 755 2014, 2016; Scibiorski et al., 2019). Titanite incorporates a significant amount of REE and Y, 756 of which Y and HREE are also incorporated by major mineral garnet. Scibiorski et al. (2019) 757 have demonstrated how titanite growing in competition with garnet, in garnet amphibolites 758 and gneisses (both felsic and intermediate), is significantly affected in key HREE. In their 759 study, they demonstrated that by using Y vs Dy/Yb titanite grown in garnet-absent rocks can 760 be distinguished from garnet-bearing rocks. We have plotted titanite data from our study in 761 the same diagrams (Fig. 15). Except for the Corsica eclogite, all our titanite-bearing samples 762 are garnet-free. Commonly Y substitutes with HREE and thus there is a linear correlation 763 between HREE and Y values (Fig. 15a). Titanite grains retrieved from Corsica eclogite-facies 764 rocks have a different Y/HREE slope than titanite from the other units (Fig. 15a). While in the 765 original study by Scibiorski et al. (2019), Dy/Yb faithfully discriminated titanite of garnet-766 bearing from garnet-free domains, some of our titanite from garnet-absent units (e.g. Ch3) 767 yield high Dy/Yb (> 2.0; Fig. 15b). This suggests that i) there is competition with another 768 mineral for Yb, such as zircon and/or xenotime, ii) that titanite might incorporate more Dy 769 due to break down of a main phase that favours MREE such as zoisite (Frei et al., 2004) or iii) 770 771 that at HP-LT conditions, Yb partition coefficient in titanite changes.

Cr-Nb discrimination diagram. Rutile from our dataset (n=69) are derived from 773 metagabbros (mafic). We applied the Cr-Nb protolith discrimination diagram from Meinhold 774 et al. (2008) and while the majority of rutile plot in the metamafic field, there are 7 grains 775 classified as metafelsic-derived. These correspond to grains retrieved from the IODR 776 amphibole-bearing gabbros (samples 1316 and 1320). These results confirm that rutile 777 derived from high-grade conditions can be misidentified using this discrimination tool, such 778 as highlighted by Meinhold *et al.* (2008). This suggests that at high-grade conditions, Cr is not 779 always preferentially retained by rutile, as Cr is more diffusive than Nb at HT (Sasaki et al., 780 1985; Marschall et al., 2013) or as a new source of Nb becomes available (e.g. amphibole or 781 ilmenite breakdown). This seems to be the case, due to the incongruent replacement of 782 ilmenite in the ilmenite-titanite-rutile aggregates. It is not fully clear whether rutile and 783 ilmenite are initially stable together, as textural relationships seem to suggest, and thus 784 competing for Nb. No evidence of HT amphibole breakdown was identified in sample 1320, 785 and this can explain Cr being sequestered by either HT amphibole or by primary 786 clinopyroxene (such as in Tribuzio et al., 2019). 787

High Field Strength Elemental ratios. Zr/Hf and Nb/Ta values from rutile plot within 788 known compositions of other metamorphic rutile, except for 4 grains that have  $\Sigma Nb/Ta$ , Zr/Hf 789 < 25, which plot within the fluid-related rutile field of Pereira et al. (2019). These grains 790 come from sample IODP1319, a LP amphibolite, where rutile occurs as spherical inclusions 791 always surrounded by ilmenite. As discussed in the previous section, these rutile grains seem 792 793 to reflect a higher temperature growth, in respect with the fine rutile lamellae, as they occur as inclusions in ilmenite. Yet, these grains yield the lowest values for several trace elements 794 within the IODP dataset (Cr < 11, Zr < 204, Nb < 450 ppm). This might imply that these 795

797 Zr/Hf values (7.22 - 13.66) are consistent with rutile formed at LT (Ewing *et al.*, 2014).

#### **4.3.** Are titanite and rutile chemistries pressure and temperature sensitive?

799 *4.3.1. Titanite* 

We filtered our titanite chemical data using Principal Component Analysis. This allowed to 800 select certain trace elements that better explain the variance in the dataset, and plot them in a 801 Yb, V+Cr and Nb+Ta ternary diagram (Fig. 16). Titanite grains from our LP oceanic 802 amphibole-bearing gabbros plot closer to the Nb+Ta endmember, while titanite belonging to 803 the HP assemblage from HP metagabbros plot between the Yb and the V+Cr endmembers 804 (Fig. 16a). Similarly, La/Sm, V and Yb allow to distinguish LP from HP titanite, as HP 805 titanite are strongly depleted in LREE relatively to the MREE (Fig. 16c). To test these 806 systematics, we compiled titanite trace element data from igneous rocks, gneisses (ortho-807 mafic and intermediate, and calc-rich), and HP-LT rocks from the literature (Fig. 16b, d). 808

Nb+Ta vs Yb. Igneous-related titanite (either more mafic or more felsic) is clearly 809 distinguishable from other titanite and plot closer to the Nb+Ta apex in the ternary diagram 810 811 (Fig. 16b). HT metamorphic titanite grains show similar behaviour, despite some outliers in the garnet amphibolite that tend towards the Cr+V endmember. Blueschist-derived titanite 812 from the literature plot mostly in the centre of the diagram and tend to the Yb endmember (Nb 813 = 10 to 30 ppm; Yb = 15 to 40 ppm) and are easily distinguished from the remaining titanite 814 (igneous and HT). Yet, a few eclogite titanite grains from our study and other blueschist 815 titanite (New Caledonia) plot closer to the Nb+Ta endmember (Fig. 16b). This implies that 816 817 another mechanism is accountable for those titanite compositions. Eclogite from Corsica are often garnet bearing (C147), and therefore those titanite were in competition for HREE, 818 explaining their lower Yb. The New Caledonia titanite form in garnet-free, metamafic and 819 metapelitic blueschist (Spandler et al., 2003) and thus another explanation is required for their 820

Yb depletion when compared with the other blueschist-related titanite grains (8 to 30 ppm vs 821 822 20 to 50 ppm, respectively). There may be: i) competition for Yb with other minerals (such as apatite and lawsonite; Spandler et al., 2003), scavenging it before titanite growth, or ii) a 823 relatively low Yb bulk composition (Spandler et al., 2003). In summary, although a ternary 824 diagram such as presented in Fig.16b discriminate a large part of our LT-HP titanite dataset 825 from other titanite, this diagram fails to discriminate all LT-HP reported in this figure and 826 therefore a few other competing factors can influence these trace-elements distribution into 827 828 titanite.

V content in HP titanite. Some titanite grains in our dataset show significant V 829 contents (> 500 ppm), markedly in blueschist titanite from the Queyras. Minerals like 830 pargasite, phengite, omphacite or epidote are major carriers of V. Indeed, titanite growth after 831 biotite or hornblende breakdown has been shown to be more enriched in V (Xiao et al., 2020). 832 Blueschist from Queyras are predominantly lawsonite and glaucophane bearing, minerals that 833 do not particularly favour V incorporation (Spandler et al., 2003). This might explain why V 834 is more incorporated by titanite in these samples. Conversely, in Corsica samples, titanite is 835 equilibrated with omphacite who disputes more with titanite for V (Hermann, 2002; Spandler 836 et al., 2003). This seems to suggest that in the absence of omphacite or phengite, HP titanite 837 grains can yield higher V contents, but the role of  $fO_2$  should also be considered as V is redox 838 sensitive. 839

*REE contents, and the La/Sm vs Yb vs V ternary.* It must be noted that on average, *REE contents of igneous titanite are high, particularly for the LREE (Xiao et al., 2014;*Bruand *et al., 2020;* Scibiorski and Cawood, 2022) and therefore these can be used as a first
approach to distinguish igneous from metamorphic titanite. Titanite from our LP amphibolite
facies samples go from depleted to mildly enriched REE, with no significant REE
fractionation patterns. On the other hand, LT-HP blueschist and eclogite titanite grains show
strong depletion in REE and more particularly in LREE, which results in La/Sm<sub>N</sub> 846 fractionation and in negative slopes of the LREE pattern (Fig. 10). Lanthanum is mostly taken 847 by lawsonite and epidote in mafic blueschist, and by apatite in more pelitic compositions 848 (Spandler et al., 2003; Martin et al., 2014). However, epidote does not intake Sm (Spandler et 849 al., 2003). This implies that epidote can lead to variable La and Sm availability and thus 850 fractionation of these REE. We have thus used La/Sm<sub>N</sub> against Yb (HREE) to investigate their 851 potential as pressure discriminants applied to titanite from our study (Fig. 16c). Alpine HP 852 units plot at lower La/Sm<sub>N</sub> values and spread in between Yb and V endmembers. This is in 853 sharp contrast with titanite grains from igneous and lower amphibolite conditions that tend to 854 display no to small La/Sm<sub>N</sub> fractionation and therefore plot close to the La/Sm<sub>N</sub> apex (Fig. 855 16d). By applying a multiple-step discrimination approach, resourcing to the two ternary 856 diagrams (Fig. 16b and d), different titanite types formed at variable settings and grades can 857 858 be distinguished.

859 *4.3.2. Rutile* 

Previously, we showed that the variability in our rutile dataset can be best described using 860 V+Cr, Sc, and Nb endmembers in a ternary diagram (Fig. 11c). Based on this diagram, rutile 861 from eclogite can be distinguished from amphibolitic-granulitic rutile. Available rutile data 862 from the literature are also reported in two diagrams (Fig. 17). Despite the limited amount of 863 mafic rutile data available from the literature reporting Sc, data from the literature agree well 864 with the defined fields from this study. HP rutile from Syros and from a talc eclogite from 865 Monviso plot together with eclogite rutile from our study in the HP field, whilst rutile from a 866 mafic granulite plot together with LP rutile from our study in the newly defined LP fields of 867 both diagrams (Fig. 17). Some rutile grains from the talc eclogite reported by Angiboust et al. 868 869 (2014) are partially overlapping with some HT rutile in the 0.3-0.5 Nb/V range (Fig. 17b). Yet, the overall reproducibility of this diagram seems to sustain. 870

While only mafic-derived rutile were used here to test trace-element pressure 871 872 dependency, there are however subtle compositional variations within gabbros and basalts from the ocean floor that can lead to variable rutile compositions. Due to the lack of available 873 bulk trace element compositions for the set of rocks used in this study, we have, alternatively, 874 tested the influence of composition by forward modelling trace element re-distribution in 875 rutile using RStudio<sup>®</sup>. We have resourced to a compilation of trace element abundances from 876 MORB and ophiolite gabbros (including some in our collection), and selected the elements of 877 interest (V, Sc, Nb). We used parageneses and a range of mineral proportions like those found 878 in our samples for both eclogitic (HP) and granulitic (HT) conditions, and published 879 partitioning coefficients for all the mineral pairs considered. Details on the modelling 880 parameters, inputs and outputs (ternary diagrams) are presented in the Supplementary Data: 881 Electronic Appendix 5 and the script is available as Appendix 6. Even though generous 882 uncertainties are incorporated into the forward model, as a reflection of large uncertainties in 883 the mineral partitioning coefficients and phase proportions, rutile compositions of eclogite 884 and granulite are always clearly separated, showing that this dichotomy is not primarily a 885 function of bulk-rock geochemistry. The eclogite to granulite split occurs along the line 886 parallel to the Sc-Nb join, as seen with our natural rutile data (Fig. 11c), indicating significant 887 differences in the fractionation of V. Given that rutile-pyroxene partition coefficients differ 888 only within a factor of two or so (Holycross and Cottrell, 2022), it appears that V distribution 889 is controlled by major mineral assemblages which are clearly different between eclogite and 890 granulite. It suggests that variations towards Nb are mainly controlled by the fraction of rutile 891 in the rock, and the smaller the rutile fraction is, Nb in rutile increases. 892

What can explain V and Sc pressure dependency in rutile? Metamorphic rutile formed
from mafic protoliths is commonly enriched in V (> 1000 ppm; Miller *et al.*, 2007; Gao *et al.*,
2010; Hart *et al.*, 2016). In our dataset, HP rutile plot systematically closer to the Cr+V

endmember when compared to the HT rutile. Vanadium is generally compatible with rutile, 896 although at times it behaves as incompatible  $(D_V^{\text{Rutile/melt}} 0.1 - 18)$ , Klemme et al., 2005; 897  $D_V^{\text{Rutile/melt}} 47 - 124$ , Foley et al., 2000). It has been demonstrated that V partitioning to rutile 898 increases with increasing fO<sub>2</sub> (Holycross and Cottrell, 2020). This implies that V will be 899 easily incorporated by rutile under most oxidising geological conditions. At higher 900 temperature conditions, V might be preferentially incorporated into other phases (e.g. 901 ilmenite, clinopyroxene, amphibole; Nehring et al., 2010), and, as a consequence, rutile will 902 contain less V and proportionally more Nb, for which it is the main carrier (Zack et al., 2002). 903 This seems to be the case of HT rutile grains that are plotted in Figure 17. The modelling also 904 highlighted that V incorporation by rutile can also be controlled by bulk composition. Thus, V 905 in rutile is controlled both by the element concentration,  $fO_2$  and the paragenesis. As a result, 906 if the compositional system is homogeneous (MORB), the critical factors are oxygen fugacity 907 908 and minerals in competition with rutile.

Although rutile is commonly poor in Sc, Sc can behave similarly to Sb. Sc is often 909 retained by amphiboles (e.g. glaucophane, pargasite), garnet or clinopyroxene in metabasites 910 (Hermann, 2002; Spandler et al., 2003). Some Sc can be partitioned into rutile due to 911 amphibole or primary Ti-rich clinopyroxene breakdown. This could account for some variable 912 Sc found in rutile from the IODP and as inclusions in the Queyras titanite grains. As the 913 modelling highlighted, the interplay of partitioning coefficients between the different 914 clinopyroxenes (omphacite and augite) and hornblende exerts a strong control on Sc contents 915 in rutile. 916

# **4.4.** Applications and future directions using Ti-bearing minerals for LT-HP hunting As we probe the exposed and preserved metamorphic rock record through time, retrieving HP metamorphic conditions before the Neoproterozoic has been challenging (e.g., Brown and Johnson, 2018; Palin *et al.*, 2020). There seems to exist a preferential preservation of HT

orogenic cores in ancient mountain belts (e.g. Hercynian Massif Central). The scarcity or 921 absence of HP metamorphic rocks in ancient metamorphic belts can potentially be explained 922 by a low preservation potential and, therefore, detrital grains formed after weathering of those 923 units may be their only survivors. Rutile and titanite are easily found as detrital grains in 924 stream sediments and in sedimentary rocks (e.g. Pereira et al., 2020; Zhou et al., 2020) but P-925 dependent proxies are required to discriminate HP from LP metamorphic rutile and titanite. In 926 this contribution we have presented data of mafic rocks that underwent metamorphism both in 927 the ocean floor, at low pressure, and during subduction at higher pressures. Our data suggest 928 that Nb/V and Sc-Nb-V+Cr in rutile and V-Yb-La/Sm in titanite are pressure dependent (Figs. 929 16-17). To apply these petro-geochemical tools in sediments, 930 we propose a multi-step discrimination approach, explained in the following. 931

For titanite, a first approach is to examine BSE images to select igneous from metamorphic titanite based on BSE zoning (Bruand et al., 2014 for magmatic; Garber et al., 2017 for metamorphic). Grains showing no zoning or metamorphic zoning can be selected and their trace element data plotted on the ternary diagrams proposed here (Fig. 16-b,d). This discrimination, however, leads to some HP titanite potentially being discriminated as LP.

For rutile, data should be plotted in diagrams that discriminate metamorphic derived from other rutile types (Agangi *et al.*, 2019; Pereira *et al.*, 2019). A sub-selection of metamorphic-only rutile can then be plotted using the Cr-Nb discriminating diagram (Meinhold et al., 2008 and others) to pool metamafic-derived rutile grains. This subset can then be projected in the Sc-Nb-V+Cr and Nb/V vs  $\Sigma$ Zr/Hf, Nb/Ta diagrams (Fig. 17). Further work is required to assess suitability for rutile grains derived from metapelitic protoliths.

#### 943 **5.** Conclusions

In this study we have compared metamafic rocks with MORB affinities of modern and prealpine oceanic crust (IODP and Chenaillet massif) to metagabbros and metabasalts that have
undergone HP during subduction to the blueschist and eclogite facies conditions (Queyras,
Monviso, Corsica).

Despite rutile being mostly recognised as a stable phase at eclogite and 948 upper amphibolite/granulite conditions due to Barrovian metamorphism, we found that amphibole-949 950 bearing gabbros in the oceanic crust can stabilise rutile at HT and very LP (ca. 2 kbar) (Fig. 18). Pressure and temperature are among the most critical parameters affecting the stability of 951 these two minerals. Yet, we show that for rocks that experienced similar P-T conditions, 952 subtle variations in the Ti/Ca and Ca/Al values seem to affect the titanite to rutile formation, 953 even in rocks with similar bulk chemistry and grain sizes. These ratios, which may shift 954 locally during mineral replacement reactions, can be the main controlling factor when both 955 rutile and titanite are found stable at either LP or HP, especially in only partially equilibrated 956 metagabbros. 957

Trace-element based discrimination tools are complex, and the chemistry of mineral 958 phases are dependent on several factors including the bulk rock composition, local equilibria, 959 960 partitioning coefficients between minerals and rocks and co-existing mineral phases during Ti-phase growth. Our results highlight that titanite trace element compositions in different 961 MORB affinity metamafic rocks show strong variability. Titanite from our study shows that 962 blueschist and eclogite yield a marked La depletion, and thus low La/Sm<sub>N</sub> values in 963 964 comparison to amphibolite grade titanite. La/Sm<sub>N</sub> or Nb+Ta together with Yb and V can be used to discriminate HP titanite. Rutile grains from the metagabbros studied here display 965 966 distinctive chemistries that are pressure indicative; rutile from HT-LP yield V/Nb values above 0.5 while V/Nb in rutile from eclogite facies are below 0.3. In the same way, V+Cr, Sc, 967

and Nb compositions can be used, following the workflow described in the discussion, todistinguish HP from LP metamafic rutile.

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978 Data Availability Statement

979 The data underlying this article are available in the article and in its online supplementary980 material.

Supplementary Materials are available in the electronic version online. Appendix 1 comprises a full method description, sampling details, amphibole mineral chemistry detailed description and phase diagram modelling details. Appendix 2 includes silicate mineral chemistry and amphibole thermometry. Appendix 3 includes the full titanite, rutile and ilmenite chemistry dataset.

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Grade		I (.	LP-H' IODI	T ?)		1 (W	LP-H 7. AL	T PS)	]	HP-LT			HP-MT									
Samples	1316	1318	1320	1319	5450	CH1	CH3	CHDI	QE01	QE06	TR10	C16 Cv16 C147		0 C16 Cv16 C		C16 Cv16 C147		C16 Cv16 C147		LS	VE29	
Locality	LEG 209 3		305	118	C	henai Massi	llet if	Quey		s Co		Corsica		Passo Gallarino	Lago Superiore	Verne Monviso						
T (°C)	740 - 840			-	700-900		330 -	- 400	380 - 470	5	500 - 550		450; 525	550 - 600	500-550	R						
P (kbar)	< 2				< 2		11 -	- 15	12 - 18		21 - 2	4	12; 25-29	19; 25 - 27	21-27							
Ref.	This study			This study; G			I	Ξ	А		F		С; В	C; B	D							

A. Debret et al., 2016; B – Angiboust et al., 2012; C. Schwartz et al., 2000; D. Groppo and Castelli, 2010; E. Agard et al., 2001; F. Vitale-Brovarone et al. 2011; G. Mevél et al., 1978.

#### 1293 and HP-MT. The peak assemblage represents the highest metamorphic grade each rock attained.

				LP –	H/M1						]	HP – L	/MT				
Samples/	13	13	13	13	54	CIII	CII2	CH	OF 01		TR	С	C v	С	DC	TC	VE
Minerals	16	18	19	20	50	Спі	Спз	DI	QE-01	QE-00	10	16	16	147	PG	LS	29
Augite	m	m	m		m	m				m							
Plagioclase	m	m	m	m	m												
Hornblende <sup>a</sup>	х	х		х	х	Х	х										
Heden-																	
Diopside			х													<b>N</b>	
Garnet													х	x		X	х
Omphacite													х	x	x	x	х
Hornblende <sup>b</sup>	х	х	х	х	х	х	х		х	Х			$\mathcal{A}$		)		
Actinolite						ret	ret	х	х					)			
Glaucophane <sup>c</sup>									х	Х	х	x	x	х	х		
Epidote					х				х	х	х	x		х	х	х	х
Lawsonite									х	х	x	$\checkmark$	х	Х			
Albite		ret	ret	ret		ret	ret	х	ret	ret	ret	ret					
Quartz	х									X	$\mathbf{Y}$	х		Х	х	х	х
white mica										x	x				х	х	х
Ilmenite		m/x	m/x	m/x	m/x	х	Х	х			х				х	х	х
Rutile	x i		x i	x i	x i	x i	x i	x i	i	li	i				х	х	х
Titanite	х	х	x ov	x ov	х	х	х	x	x	х	х	х	х	х	ret		
Chlorite	ret				ret	ret					х		ret				
Carbonate	ret		ret		ret				7	ret	ret						

Note that some minerals indicated on the list are one possible end-member. a. another equivalent brown amphibole; b. another equivalent green amphibole; c. another equivalent blue Na, Na-Ca amphibole. In the table, m - magmatic, x - metamorphic and i - inclusion, ov - overgrowth, ret - late retrograde phase

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	∑REE	∑LREE	La/SmN	Eu/Eu*	Th/U
Amphibolite	10 - 3358	2.6 - 2400	0.02 - 1.2	0.2 - 29	0.01 - 0.48
-					
Granulite*	13 - 2000	11 - 1678	5 - 93	4.7 - 314	0.08 - 1.06
Blueschist	32 - 370	3.8 - 59	0.02 - 0.7	0.9 - 3.2	0.04 - 0.2
Eclogite	168 - 465	20 - 208	0.02 - 0.3	0.84 - 1.8	0.004 - 2.3

REE sums are in ppm. Data was normalized to Sun and McDonough (1995) Chondrite values. Amphibolite corresponds to a combination of Ch3 and IODP1319, granulite to IODP1320, blueschist to a combination of QE1, QE6, TR10, and eclogite to C147, C16v and C16 data. \*titanite in the granulite sample is most probably formed at lower temperatures, but the sample rock forming minerals did not retrogress to those conditions.



1320

1327

1321 Figure 1 Experimental vs pseudosection modelling diagrams for a MORB composition and focusing on Ti-

- 1322 phases. A. P-T diagram highlighting the fields of known stability of rutile, titanite and ilmenite for MORB
- 1323 compositions (Liou et al., 1998). In white a field of no experimental data available for mafic compositions. Grey
- 1324 lines delimit different metamorphic facies. Two dotted lines separate the fields of low (< 440 °C/GPa) from
- 1325 moderate and higher geothermal gradient (> 770 °C/GPa) (Palin et al., 2020). Boxes correspond to
- 1326 approximate peak P-T conditions from samples collected for this study, highlighting cooling of amphibole-

bearing gabbros from the oceanic crust, and burial history of blueschist and eclogite rocks from the Alpine

- 1328 chain. B. MORB phase diagram modelling from Green et al. (2016) showing titanite-out reaction (in green),
- solid blue line the same as in a) and a dotted white line at 600 °C, below which there are no data constraints on
- **1330** *Ti-phase stability for mafic compositions (taken from inset a).*
- Figure 2. Sampling locations in the Western Alps and Corsica (A, B) and from IODP drilling sites (C). Western Alps maps were modified after Angiboust et al., (2012), Debret et al. (2016), Locatelli et al., (2019), and Vitale-Brovarone et al., (2011), and IODP drilling site information and map were modified from data and maps available at iodp.org.
- Figure 3. Microphotographs of samples from IODP using PPL (A, C, G, H), reflected light (D- E) and BSE imaging (B, F). A-B, occurrence of sub-spherical rutile grains rimmed by titanite enclosed by altered plagioclase and brown to green amphibole; C, occurrence of large ilmenite grains rimmed by titanite; E-F ilmenite-rutile-titanite grains, with titanite replacing former ilmenite usually found at the cores; F is a zoom in of E, illustrating the occurrence of rutile rimmed by ilmenite and only then by titanite; G-H, occurrence of idioblastic titanite associated with altered plagioclase and brownish-green amphibole.
- Figure 4. Microphotographs of samples from the Chenaillet Massif using PPL (A-D, G), reflected light (E, F) and BSE imaging (H). A and B, general textural and mineralogical relationships; C, of the late-stage titanite and actinolite shear zone in Ch1; D-F. detail of ilmenite-rutile-titanite occurrences in ChDi; G-H ilmenitetitanite growth in the shear zone, and strong BSE zoning in these titanite grains.
- Figure 5. Microphotographs of blueschist samples from the W. Alps in the Queyras using PPL (A, B, C),
  reflected light (D) and BSE imaging (E, F). A- plagioclase domains with lawsonite, albite and epidote, and relic
  magmatic clinopyroxene replaced by glaucophane and epidote and titanite. B- overgrowth of actinolite on

- glaucophane in plagioclase sites, with titanite. C-F, titanite grains with very fine-grained lamellae rutile
  inclusions in all the different investigated samples.
- 1350 Figure 6. Microphotographs of eclogite samples from the W. Alps in the Passo Gallarino-Monviso-Lago
- 1351 Superiore (A-D) and Corsica (E-H) using PPL (B, C, E-G), reflected light (A) and BSE imaging (D). A in
- 1352 Passo Gallarino a thin rim of titanite is found around rutile; B- omphacite, garnet, and phengite associated to
- 1353 rutile; C Foliated eclogite with omphacite, garnet, and rutile ribbons; D a detail of a rutile ribbon from C
- 1354 with ilmenite/Fe-Ti oxide exsolution; E- blueschist eclogite, with albite, glaucophane crossed-cut by a vein
- 1355 showing omphacite, garnet, lawsonite and titanite; and G-H- eclogitic metabasalt with garnet, omphacite,
- 1356 *lawsonite, glaucophane and titanite.*
- Figure 7. P-T diagram illustrating the conditions of either rutile or titanite stability found in the samples studied
  here. P-T constraints are from references found in table 1.
- Figure 8. X<sub>Al</sub> vs F<sub>applu</sub> in titanite binary diagram. Different symbols and colours represent average titanite values from different samples. Green symbols are data from this study. Diamonds are for titanite from IODP samples, squares for the obducted Chenaillet metagabbro, circles for blueschist and triangles for eclogite titanite grains from this study. All other coloured symbols correspond to titanite data from other metamorphic grades and bulkrock compositions that are plotted for comparison. Literature data from Franz and Spear, 1985, Markl and Piazolo, 1999, Ye et al., 2002, Castelli and Rubatto, 2002, Garber et al., 2017.
- Figure 9. Trace element diagrams for titanite, A) Nb/Ta vs Zr/Hf. Star refers to Chondrite compositions (Sun
  and McDonough, 1989); B) Sn vs V binary diagram. Diamonds correspond to IODP, squares to the Chenaillet
  Massif amphibolite metagabbro, circles to blueschist and triangles to eclogite rock titanite.
- 1368 Figure 10. REE and Th/U values of titanite from HT and HP conditions.
- Figure 71. Discrimination diagrams for rutile. A) The Cr-Nb discrimination modified from Meinhold et al.
  (2008), B) The Zr/Hf vs Nb/Ta modified from Pereira et al. (2019); C) based on PCA, the three main vectors V,
  Cr, Sc and Nb used in a ternary diagram; D) Binary diagram of the sum of HFSE ratios vs Nb/V. HT IODP are
  presented with two colours, the brighter red corresponding to sample 1320 and darker red to 1319 and 1316.
  Some labelled datapoints correspond to LA-ICP-MS acquired data used in Figure 11.
- Figure 82. Rutile occurrence in amphibole-bearing gabbros on the oceanic crust (example from IODP1320); a)
  Reflected light microphotograph in grayscale of an ilmenite grain being replaced by titanite and rutile. Rutile

- 1376 occurs as exsolution lamellae and coarse grains; b) elemental concentrations of Zr and Cr in rutile from
- 1377 *different textural domains labelled in part a*). *The symbol* \* *correspond to a laser spot measurement.*
- **1378** Figure 93. Binary trace element diagrams for ilmenite grains from samples IODP 1320 and 5450.
- 1379 Figure 10. Composition in Ti/Ca and Ca/Al of several metagabbros, blueschist and eclogite rocks from different
- 1380 settings (IODP, Syros, Western Alps, Corsica) grouped by titanite, rutile or titanite-rutile stable (symbols).
- 1381 Range and average of unaltered MORB composition is given by the red circle and the red star respectively
- 1382 (Klein, 2004). Altered vectors correspond to altered MORB compositions from Zhang and Smith-Duque (2014).
- 1383 Corsica data are from Vitale-Brovarone et al. (2011) and Syros samples from Marschall (2005). Data used here
- 1384 *can be found in the Supplementary Data: Electronic Appendix 1.*
- 1385 Figure 11. Trace element diagrams for titanite modified from Scibiorski et al. (2019). A. **EHREE vs Y; B. Y vs**
- 1386 Dy/Yb; In grey, labels adapted from Scibiorski et al. (2019) and in pink based on data from this study. Pink-
- 1387 coloured circles correspond to titanite grains in a garnet-bearing assemblage.
- Figure 12. Ternary diagrams for titanite compositions from multiple geodynamic environments. Compiled
  literature data are from a. Bruand et al., 2014, b. Garber et al., 2017, c. Holder and Hacker, 2019, d. Marsh and
  Smye, 2017, e. selection of Scibiorski et al., 2019, f. Spandler et al., 2003, g. Ng et al. 2016, h. El Korh et al.,
  2009.
- Figure 13. Rutile discrimination diagrams for metamafic, gabbro in composition, rocks .a) is a V+Cr-Sc-Nb ternary and b) a V/Nb vs sum of HFSE ratios. Grey bar is highlighting some overlap between HP and LP signatures and arrow indicate the field towards HP rutile. a. relic in bs correspond to HT-LP rutile grains found as inclusions in blueschist titanite. Published data are from metamafic rocks b. and c. from blueschist and eclogite from Syros (Hart et al., 2016), talc and lawsonite eclogite from Monviso (Angiboust et al., 2014) and d. LP granulite (Cruciani et al., 2020).
- 1398 Figure 14. Ti-mineral phases in metagabbros and metabasalts associated to divergent and convergent settings1399 and some of their main trace element systematics.
- 1400
- 1401
- 1402





## **Fig. 3**.



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### **Fig. 4**.



1414 **Fig. 5.** 





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Fig. 12. 1480



Fig. 13. 1504

1505


1508 Fig. 14.



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**Fig. 16.** 



**Fig. 18.** 

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