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1 **Ti-bearing minerals: from the ocean floor to subduction and back**

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22 Abstract

23 Rutile, titanite, and ilmenite are the most common Ti-bearing minerals in metamorphic rocks.
24 Experimental constraints have shown that titanite is stable at low-grade metamorphic
25 conditions, rutile at HP, and ilmenite at HT-LP conditions. Yet, petrological evidence
26 suggests that titanite can also be stable at LT-HP. This implies that both titanite and rutile can
27 be used to develop proxies to track HP metamorphism, which can have interesting
28 applications.

29 In this study, we have investigated the natural occurrence and chemistry of Ti-bearing
30 minerals in gabbroic rocks that record different degrees of metamorphism, including LP
31 amphibole-bearing gabbros from the ocean floor (Mid-Atlantic and Indian ridge IODP LEGs)
32 and from an obducted ophiolite (Chenaillet) and HP Alpine metagabbros including blueschist
33 and eclogite facies rocks from the Western Alps and Corsica. We have performed detailed
34 petrography and Raman spectroscopy and analysed major and trace elements mineral
35 chemistry using EPMA and LA-ICP-MS.

36 We found that rutile is stable at low pressure (< 2 kbar) in ocean-floor amphibole-
37 bearing gabbros, lower than experimental constraints had previously suggested. Rutile is also
38 found in eclogitic metagabbros from the Western Alps and can be chemically distinguished
39 from LP rutile. Blueschist metagabbros from the Western Alps and eclogitic metabasalts from
40 Corsica have titanite stable instead of rutile. While the titanite to rutile transition is pressure
41 and temperature dependent, we demonstrate how small variations in bulk-rock Ti/Ca and
42 Ca/Al values within the NCKFMASHTO chemical system may shift their stabilities. High-
43 pressure titanite from these metamafic rocks exhibits La depletion and low La/Sm_N values in
44 comparison to titanite from amphibolite-facies mafic rocks. La/Sm_N or Nb together with Yb
45 and V can be used to distinguish HP titanite from titanite formed under other P-T settings.

46 These new systematics can be useful in studies using detrital Ti-bearing minerals to probe the
47 HP metamorphic record through time.

48 Keywords: rutile; titanite; P-T stability; trace element-based discrimination; metagabbro

49 1. INTRODUCTION

50 The concept of mineral equilibria has allowed petrologists to estimate pressure and
51 temperature conditions seen by metamorphic rocks (Hoschek, 1969; Burt and Ferry, 1982).
52 However, complete mineral equilibria of rocks are typically unattained during geological
53 processes, especially in coarse-grained rocks that often preserve primary textures, and/or peak
54 and retrograde metamorphic assemblages (non-equilibrium). Over the years, these different
55 snapshots of the rock's evolution have been exploited by petrologists, to better constrain its
56 pressure-temperature (P-T) history. Attribution of a mineral to a pre- or a given metamorphic
57 stage is paramount in reconstructing the P-T evolution of a rock, central to the geodynamic
58 interpretation of metamorphic petrology, and especially for tracing ancient subduction
59 histories (Rubatto, 2002; Engi, 2017; Yakymchuk, 2017).

60 The “classical” equilibrium thermobarometry approach requires a thorough chemical
61 characterisation of the different mineral phases that attained equilibrium in cation exchange
62 reactions. Alternative ways to this approach have arisen from the development of single grain
63 thermometers (*e.g.* Ti-in quartz, Zr-in-titanite, Zr-in-rutile), allowing to determine
64 temperature conditions from rocks with multiple heritages and/or as inclusions within another
65 mineral. So far, single grain methods are only available for temperature but not pressure, and
66 with known limitations related to uncertainties in elemental behaviour under different P-T
67 conditions and buffering assemblages (Kooijman *et al.*, 2012; Cruz-Uribe *et al.*, 2018;
68 Kendrick and Indares, 2018; Clark *et al.*, 2019). To constrain pressures by single grains,
69 barometry based on quartz Raman spectra have been developed over the last decade, mainly

70 on quartz in garnet but also on other host-inclusion pairs (e.g. Enami *et al.*, 2007; Ashley *et*
71 *al.*, 2014; Gonzalez, 2019; Cisneros *et al.*, 2020; Gonzalez *et al.*, 2021). Yet, in the absence of
72 quartz inclusions, we need to rely on grain chemistries for an alternative pressure information
73 source.

74 Titanium-bearing minerals are common accessory minerals in mafic rocks and
75 potentially serve as single grain P-T indicators as referred above. The stability of these
76 minerals, titanite, ilmenite and rutile, are controlled by different factors: system composition,
77 pressure and temperature (Force, 1991; Liou *et al.*, 1998) and fluid composition and its
78 dissolved ions (Spear, 1981; Force, 1991). Rutile is commonly found at higher metamorphic
79 grades, at upper amphibolite and granulite facies (Goldsmith and Force, 1978; Hart *et al.*,
80 2018), and at high pressures, such as in the eclogite facies or UHP (Liou *et al.*, 1998; Gao *et*
81 *al.*, 2004; Hart *et al.*, 2016). Rutile is seldom found at lower grades (greenschist facies), and
82 usually associated with fluid alteration processes (Meinhold *et al.*, 2008; Zeh *et al.*, 2018).
83 Titanite, conversely, is mostly found at lower metamorphic grades (greenschist, low
84 amphibolite, blueschist and occasionally eclogite facies conditions; Force, 1991; Vitale
85 Brovarone *et al.*, 2011). Commonly, with increasing metamorphic grade, at upper amphibolite
86 to granulite facies conditions, titanite becomes unstable and ilmenite or rutile replace it
87 (Spear, 1981; Force, 1991; Liou *et al.*, 1998). Ilmenite is favoured by high Fe/Mg bulk-rock
88 compositions (Force, 1991) and medium fO_2 conditions (Spear, 1981; Guo *et al.*, 2017).

89 Experimental constraints on the stability of these three Ti-phases for MORB
90 compositions are shown in Figure 1a. The lack of experimental data below 600 °C for mafic
91 compositions has been partly compensated by natural observations and by solution and
92 thermodynamic models of minerals and mineral equilibria (White *et al.*, 2014; Green *et al.*,
93 2016). Note that the low temperature boundary in Figure 1b is only weakly constrained.
94 Notably, titanite is predicted as more stable than rutile at HP and T <550 °C (Fig. 1b),

95 although the position of the *titanite-out* reaction changes considerably depending on the
96 choice of solution models (here using the most recent by Green et al., 2016 and White *et al.*,
97 2014). At lower pressure, titanite and ilmenite are the common Ti-phases (Fig. 1). Yet, the
98 occurrence of rutile-titanite in oceanic gabbros has been reported (Batiza and Vanko, 1985;
99 Mével, 1987; Abe, 2011; Harigane *et al.*, 2011; Rizzo *et al.*, 2019). While titanite presence in
100 these gabbros has been generally linked to the cooling of the oceanic crust during ocean floor
101 metamorphism, the presence of rutile, however, is more puzzling, as it may form due to
102 magmatic or metamorphic processes. Altogether this implies that details on their stabilities at
103 different P-T and chemistry remain puzzling. Until now, despite some attempts to develop
104 trace element-based barometers applied to titanite and rutile (*e.g.* Al-F in titanite – Tropper *et*
105 *al.*, 2002; Al in rutile – Escudero *et al.*, 2012), no dependency has yet been found (Castelli
106 and Rubatto, 2002; Hoff and Watson, 2018).

107 In this study we focus on the fate of these Ti-phases from the magmatic stage of the
108 oceanic crust through its subduction. We have investigated the phase stability and trace
109 element systematics of rutile and titanite, two common Ti-bearing metamorphic minerals,
110 which are calibrated as geothermometers, to assess their potential as pressure discriminants.
111 Minor and trace element compositions of accessory minerals have been successfully used in
112 the past to discriminate rock types (*e.g.* Bruand *et al.*, 2016, 2020; Guo *et al.*, 2020; Scibiorski
113 and Cawood, 2022). By investigating the trace element compositions of these Ti-bearing
114 phases at different metamorphic grades, we explore systematics based on pressure variations.
115 Specifically, we have examined a suite of metamafic rocks with MORB-N affinity (gabbros
116 and basalts) from (i) both the modern and obducted oceanic crust, and (ii) from subducted
117 oceanic crust metamorphosed at blueschist and eclogite facies and exhumed back to upper
118 crustal levels (Fig. 1a and 2). We ultimately investigate which mechanisms control their
119 stability and test if their chemistries change with metamorphic grade. The main purpose is to

120 provide a novel tool that enables the discrimination of Ti-phases either in or out of
121 petrological context.

122 **2. SAMPLES AND METHODS**

123 **2.1. Sampling and Geological settings**

124 Samples come from two different settings: (i) high temperature, low pressure (HT-LP)
125 gabbros metamorphosed during slow seafloor spreading, and (ii) low to medium temperature,
126 high pressure (L/MT-HP) gabbros and basalts metamorphosed during convergent processes.
127 These samples were selected based on their N-MORB-affinity and due to the presence of Ti-
128 phases. HT-LP samples were collected from the Chenaillet Massif, Western Alps (3 samples)
129 and from three different IODP legs, 305 (site U1309D), 209 (site 1271B), both from the
130 Atlantic Ocean Ridge, and 118 (site 735B) in the Indian Ocean Ridge, all with HT
131 amphiboles. L/MT-HP samples were collected from the Western Alps (Queyras, Monviso)
132 and Corsica (San Petrone). More details are given in Figure 2 and Table 1 and the exact
133 sampling locations are given in the Supplementary Data: Electronic Appendix 1.

134 *2.1.1 International Ocean Drilling Program samples: Atlantic and Indian oceans*

135 Samples collected from all the following IODP sites and cores are affected by ocean floor
136 metamorphism during gabbro emplacement and cooling of the oceanic crust. Samples used in
137 this study correspond to gabbro that yield a significant oxide content ($> 1\%$), and
138 metamorphic amphiboles. Preserved metamorphic grades are variable, with some samples
139 preserving a HT metamorphic mineral assemblage with equivalent temperatures of granulite
140 facies whilst others have been re-equilibrated at amphibolite facies conditions. This allows an
141 assessment of the Ti-bearing phase budget present in the oceanic crust before it undergoes
142 subduction and consequent HP metamorphism. More details about the drilling sites and cores
143 can be found in Supplementary Data: Electronic Appendix 1.

144

145 LEG 209, SITE 1271B (0 – 104 m, 15°20'N, 45°00'W)

146 Site 1271 is located along the track of Faranaut 15°N on the western flank of the Mid-Atlantic
147 rift valley. Roughly, 100 m of core were retrieved, corresponding to dunite, harzburgite,
148 troctolite, gabbro, minor gabbronorite, and a small but significant amount of chromitite cm-
149 thick lenses (Shipboard Scientific Party, 2003). Interstitial gabbros in impregnated peridotites,
150 olivine gabbros, and troctolites from Site 1271 include a large proportion of high-temperature
151 amphibole. Samples used in this study, 1316, 1318 and 1320, are all metamorphosed gabbros
152 and were retrieved from cores 3R (piece 19, 18 m depth), 5R (piece 3, 28 m depth), and 8R
153 (piece 5, 41 m depth), respectively, all within Unit I, which was classified as dunite/gabbro
154 (Shipboard Scientific Party, 2003).

155 LEG 305, SITE U1309D (0 – 1645 m, 30°10'N, 42°07'W)

156 Site U1309 is located on the central dome of Atlantis Massif, 14–15 km west of the median
157 valley axis of the Mid-Atlantic Ridge (MAR), where the seafloor coincides with what is
158 interpreted to be a gently sloping, corrugated detachment fault surface (Blackman, 2006). In
159 this site, oxide gabbro (defined by the presence of >2% modal Fe-Ti oxide minerals) makes
160 up about 7% of the rocks recovered. Sample 1319 is an amphibolitised oxide gabbro and
161 corresponds to unit 353, piece 1, recovered at about 662 m depth (Structural Unit II;
162 Blackman, 2006), defined as the Gabbronorite Zone 2 (Blackman, 2006).

163 LEG 118, SITE 735B (0 – 501 m; 32°43'S, 57°16'E)

164 Site 735 is located on a shallow platform, on the east rim of the Atlantis II Transform located
165 in the magnetic anomaly equivalent to about 12 Ma (Robinson *et al.*, 1991). This platform,
166 about 9 km long in a north-south direction and 4 km wide, is one of a series of uplifted
167 blocks. A total of 435 m of olivine gabbro, olivine-bearing gabbro, two pyroxene gabbro,

168 iron-titanium oxide gabbro, troctolite, and microgabbro with rare basalt and trondjemite was
169 recovered from Hole 735B. These rocks have undergone varying degrees of plastic and brittle
170 deformation, and many have well-developed foliations. The sample used in this study (5450)
171 is a metamorphosed oxide gabbro and comes from unit VI, core 85R, section 4, corresponding
172 to piece 1E of the original log and to approximately 475m depth. The unit this sample comes
173 from corresponds to olivine-rich gabbro having frequent layers of troctolite (Robinson *et al.*,
174 1991). Frequent intervals of gabbro are characterized by mylonitic to porphyroclastic textures,
175 the latter corresponding to the collected sample.

176 2.1.2. *The Liguro-Piemontese domain in the Western Alps and Corsica*

177 The Piemonte-Liguria domain in the Western Alps represents the westernmost part of
178 the Alpine arc and extends to Corsica in the south (Fig. 2 A, B). This domain is interpreted as
179 a remnant of the accretionary wedge, where oceanic sediments and relics of oceanic
180 lithosphere are accumulated (Marthaler and Stampfli, 1989; Deville *et al.*, 1992; Agard *et al.*,
181 2001). In the Piemonte-Liguria zone we studied two areas: (i) a west-east transect going from
182 the obducted ophiolite in the Chenaillet Massif, to the blueschist Queyras/Combin unit
183 towards the eclogitic Viso/Zermatt unit, and (ii) Alpine Corsica in the San Petrone Mount
184 area. Together, these samples allow a comparison between what constituted the pre-Alpine
185 oceanic lithosphere crust (preserved in the Chenaillet ophiolite as the least metamorphosed
186 part of the Western Alps) and oceanic lithosphere units that underwent subduction and
187 metamorphism at different depths during the Alpine orogeny.

188 *Chenaillet Massif.* The Chenaillet amphibole-bearing gabbros were long interpreted as
189 corresponding to obducted amphibole-bearing gabbros recording ocean floor metamorphism
190 (Mevel *et al.*, 1978). They formed in the Piemonte-Liguria basin during the Jurassic, and were
191 so far described as negligibly affected by Alpine metamorphism (Manatschal *et al.*, 2011 and

192 references therein). Recent work by Corno *et al.* (2023) describe for the first time the presence
193 HT assemblage along with remnant of Alpine HP assemblages (e.g. lawsonite, omphacite) in
194 some of the gabbros and volcanics, questioning the interpretation of previous authors
195 implying the obduction of the Chenaillet massif. HP alpine metamorphism (presence of
196 lawsonite and omphacite) were not observed in the studied gabbros. Recent ages obtained
197 from magmatic zircon in the studied gabbros yield crystallisation ages of around 161 ± 0.8
198 Ma (Nicollet *et al.*, 2022). In this locality, three amphibole-bearing gabbros have been
199 sampled (Table 1).

200 *Queyras and Monviso Massifs.* The blueschist-facies Queyras and eclogite-facies
201 Monviso metaophiolite units consist mainly of meter to kilometer-sized metagabbroic lenses
202 included in metasediments (Schistes Lustrés complex; SL) and/or serpentinite with
203 omphacite (Ballèvre *et al.*, 1990). Magmatic ages of the Monviso eclogite protoliths have
204 been determined using coupled zircon U-Pb and trace element chemistry to 163 ± 2 Ma
205 (Rubatto and Hermann, 2003). In the Western Alps, the subduction stage started during the
206 late Cretaceous at around 80 Ma (e.g. Rubatto *et al.*, 2011) and lasted until at least 45 Ma
207 (Rubatto and Hermann, 2003). A west-east transect across the Queyras and Monviso units is
208 affected by a HP-LT metamorphic event, which increases towards the East from blueschist (8-
209 14 kbar; 350-450°C) to eclogite facies conditions (up to 20 kbar; 450-500°C) (Blake and
210 Moore, 1995; Agard *et al.*, 2001; Schwartz, 2002; Schwartz *et al.*, 2013). The Liguro-
211 Piemontese units are partially retrogressed into greenschist facies conditions along an
212 isothermal retrograde path (Caby, 1995), but fast exhumation, probably combined with
213 restricted fluid circulation, favoured the preservation of the peak metamorphism assemblages
214 (Rubatto and Hermann, 2001; Angiboust and Glodny, 2020).

215 Across the Queyras and Monviso, different localities where HP metagabbros occur
216 have been sampled: (i) three lawsonite blueschist from Sommet Bucher (Q1), Vallon de

217 Clausis (Q6) and Tour Réal (TR10); ii) three eclogite from Verne (VE29), Passo Gallarino
218 (PG) and Lago Superiore (LS). Sampling details can be found in Table 1 and full petrographic
219 descriptions in section 4.

220 *Alpine Corsica.* The schistes lustrés (SL) units in Corsica have undergone lawsonite
221 blueschist- to lawsonite eclogite-facies metamorphism during the Alpine orogeny and
222 suffered variable retrogression to the greenschist facies (Fournier *et al.*, 1991; Vitale
223 Brovarone *et al.*, 2011; Vitale Brovarone and Agard, 2013; Vitale Brovarone and Herwartz,
224 2013). Mafic eclogites found in the SL have N-MORB affinity (Saccani *et al.*, 2008). Ages of
225 gabbros from the oceanic units have been constrained using Sm-Nd isochrons to 162 ± 10 and
226 159 ± 15 Ma (Rampone *et al.*, 2009). HP metamorphism affecting these units has been dated
227 using a Sm-Nd isochron yielding 84 ± 5 Ma (Lahondère and Guerrot, 1997), but more recent
228 studies have clearly demonstrated that HP metamorphism in the SL took place during the Late
229 Eocene (37 – 34 Ma) (Vitale Brovarone and Herwartz, 2013).

230 Two samples were selected for this study that were collected from similar areas to the
231 ones figuring in Vitale Brovarone *et al.* (2011) (samples OF3158 and OF3310) from the San
232 Petrone Mount in Corsica. These correspond to coexisting lawsonite eclogite and blueschist
233 eclogite rocks equilibrated at the same P-T conditions of 520 ± 20 °C and 23 ± 1 kbar (Vitale
234 Brovarone *et al.*, 2011).

235 **2.2 Methodology**

236 *2.2.1 Whole-rock geochemistry*

237 Eclogite VE29 and some selected blueschist samples (Supplementary Data: Electronic
238 Appendix 1) were crushed using a jaw crusher and pulverised using the sample preparation
239 facilities at the Laboratoire Magmas et Volcans (LMV), Clermont-Ferrand (France) to obtain
240 major element compositions. Bulk rock major elements of these samples were analysed using

241 a Jobin-Yvon ULTIMA Inductively Coupled Plasma-Atomic Emission Spectroscopy at the
242 LMV, following similar analytical and instrumental procedures to the ones reported in Barette
243 et al. (2017) and Guitreau et al. (2019). Bulk-rock major and minor element compositions of
244 IODP samples were compiled from the IODP data repository. Bulk rock available on Corsica
245 eclogites were compiled from Vitale-Brovarone et al. (2011). All whole-rock data can be
246 found in the Supplementary Data: Electronic Appendix 1.

247 *2.2.2 Scanning Electron Microscopy*

248 Most collected samples were fist-sized and therefore only minerals from polished thin section
249 were analysed during this study. Ti-bearing minerals were identified and imaged to reveal
250 internal zoning or inclusions using Scanning Electron Microscopy (SEM) sessions at the
251 LMV using a JEOL JSM-5910 LV with a tungsten filament and 15 kV accelerating voltages
252 for backscatter electron imaging (BSE).

253 *2.2.3. Raman Spectroscopy*

254 To determine TiO₂ polymorphism we used a Renishaw InVia confocal Raman micro-
255 spectrometer equipped with a 532 nm diode laser (200 mW output power) and a CCD
256 detector and a Leica DM 2500 M optical microscope at the LMV. A 100× objective in high
257 confocality setting (slit aperture of 20 µm) was used and the range of power applied was
258 always lower than 1% of the laser energy. Acquisition time was set at 10 s, and 5
259 accumulations.

260 *2.2.4. Electron Probe Microanalyses*

261 Silicate and Ti-mineral phase minerals were analysed for major, minor and occasionally trace
262 elements with a CAMECA SX100 electron probe microanalyser (EPMA) at LMV. Variable
263 analytical conditions were used to target different elements for the different minerals of
264 interest. A detailed instrumental and analytical setup for each session can be found in the

265 Supplementary Data: Electronic Appendix 1. For silicate mineral characterisation the
266 following elements were determined: Si, Ti, Al, Mg, Fe, Mn, Ca, Na, K, and Cr. Data can be
267 found in Supplementary Data: Electronic Appendix 2. For Ti-mineral phases, the following
268 elements were part of different analytical sessions: Si, Ca, Al, P, Ti, Mn, Fe, Cr, F, Zr, Nb,
269 La, Ta, Y, Sc, Sn, V. Khan titanite reference material (Heaman, 2009) was used for quality
270 control monitoring during Ti-mineral phase analyses. Supplementary Data: Electronic
271 Appendix 3.

272 2.2.5. Laser Ablation-Inductively Coupled Plasma Mass Spectrometry

273 Further trace elements chemistry of titanite, rutile and ilmenite was obtained using laser
274 ablation-inductively coupled plasma mass spectrometry (LA-(HR)ICPMS) using the
275 Resonetics Resolution M-50E 193 nm excimer laser coupled to the Thermo Element XR at
276 LMV daily calibrated and optimised using NIST612.

277 The spot sizes used ranged between 9 and 27 μm with a repetition rate of 1 to 2 Hz and a
278 fluence of 2.5 to 3 J/cm^2 . Each analysis consisted of 20s background and between 50 and 90s
279 of signal acquisition. More details can be found in the appendix. We used a sample bracketing
280 method using GSE-1G and GSD-1G as primary standards and GSC-1G, GSD-1G or BHVO
281 in the middle of the sequence as secondary standards (Jochum et al., 2005). Khan (Heaman,
282 2009) was also used as a secondary reference material for additional quality control. Details
283 on the isotopes that were measured, data processing and quality control can be found in
284 Supplementary Data: Electronic Appendix 1, 3 and 4.

285 3. RESULTS

286 3.1. Petrography

287 3.1.1. Modern oceanic crust amphibole-bearing gabbros

288 *Amphibole-bearing gabbros.* All samples (1316, 1318, 1320, 1319 and 5450)
289 correspond to coarse-grained gabbro, with primary clinopyroxene and plagioclase. Overall,
290 plagioclase is the most abundant mineral phase (40-65%) but is either strongly or slightly
291 altered to secondary plagioclase (albite), quartz and variable chlorite/sericite (Fig.3-A). Relict
292 magmatic clinopyroxene (augite in composition) has reacted with plagioclase to produce
293 brown hornblende-edenite coronas around it (1320, 5450, 1316) together with albite and
294 epidote. In samples 1319, 5450 and 1316, brown amphiboles are then replaced or overgrown
295 by green hornblende and in 1319 also by actinolite + chlorite. Accessory minerals (< 3 % vol)
296 correspond to ilmenite, titanite, rutile and occasionally hematite, mostly associated to
297 amphibole. In all samples, preserved disequilibrium mineral assemblages are found. Due to
298 the variable presence and proportion of brown and green amphiboles in these samples, there is
299 a spectrum from sample 1320, which records granulite facies conditions with a brown
300 amphibole found as idioblastic crystals and in equilibrium with plagioclase, to samples 5450,
301 1318 and 1316 that have suffered partial retrogression to the amphibolite facies, and sample
302 1319 that has completely retrogressed to lower amphibolite/upper greenschist facies
303 conditions. Despite not constituting the most robust estimate, Ti-in-amphibole thermometry
304 (Otten, 1994) applied to two endmembers of these units, samples 1320 and 1319, yield
305 temperatures compatible with metamorphism under the granulite facies (840 to 740 °C) and
306 amphibolite facies (590 °C), respectively (Supplementary Data: Electronic Appendix 2),
307 providing a working frame for metamorphic conditions at which we find the Ti-phases.

308 *Ti-bearing mineral assemblage.* Titanite is the most ubiquitous Ti-bearing mineral in
309 all the metagabbro samples. It commonly forms coronas around an earlier Ti-bearing mineral
310 phase (Fig.3-A-F), often ilmenite (Fig.3-C), more rarely rutile (Fig.3-A, B) or around both
311 (Fig.3 D to F), but it can also occur as idioblastic crystals under amphibolite conditions
312 (Fig.3-G,H). In sample 1320, rutile occurs as grains typically between 50 to 100 µm in size,

313 associated with brown hornblende and within ilmenite-rutile-titanite aggregates (Fig.3-D). In
314 samples recording a lower temperature retrogression (sample 1316 and 1319), ilmenite is
315 surrounded by coronas formed by titanite and rutile showing some host crystallography-
316 controlled growth (Fig.3-D). In this case, titanite and rutile replace ilmenite during cooling to
317 amphibolite or greenschist facies conditions, illustrating its disequilibrium (Fig.3-E-F). In
318 summary, whilst in cases titanite and rutile form intergrowths in coronas around ilmenite,
319 with rutile showing some host crystallography-controlled growth (Fig.3-D), there are also
320 cases where rutile is found as coarser (20 to 50 μm), sub-spherical inclusions within ilmenite
321 (Fig.3-D-F) or, more rarely, as single grains, up to 100 μm large, directly overgrown by
322 titanite (Fig.3-A,B). While the core of some of these complex grains, formed by either
323 ilmenite or rutile, corresponds to at least granulite conditions, the titanite replacement (+ rutile
324 lamellae) or new crystals growth takes place at lower grade conditions, during cooling to
325 amphibolite or greenschist facies conditions. Ilmenite in these metagabbro rocks can either be
326 of magmatic origin or recrystallised during ocean floor metamorphism. In this way, we have
327 three (but continuous) metamorphic stages: i) augite, plagioclase and brown hornblende +
328 ilmenite/rutile, followed by ii) green hornblende, plagioclase +/- rutile, and a iii) actinolite,
329 albite, chlorite + titanite (+/- rutile lamellae).

330 3.1.2. Western Alps obducted metagabbros

331 These samples were collected from highly heterogeneous outcrops in the Chenaillet Massif,
332 where a coarse-grained texture is preserved, but occasionally amphiboles reach several
333 centimetres, and deformation variably affects these units, from highly sheared to non-
334 deformed.

335 *Amphibole-bearing gabbros.* Three samples from the Chenaillet were examined, two
336 corresponding to a gabbro composition (Ch1 and Ch3) and one interpreted as more dioritic

337 (ChDi). ChDi primary mineral assemblage has been replaced by a pale green amphibole,
338 albite, sericite, epidote and chlorite, and primary ilmenite is overgrown by titanite \pm rutile. Fe-
339 Ti oxides make up more than 3 wt% and have no preferred textural association to any major
340 rock forming mineral. The two samples of gabbroic composition are very similar, and the
341 predominant mineral assemblage corresponds to albite (secondary) and amphibole (Fig.4-A),
342 but while Ch1 has preserved relict of clinopyroxene (augite), in Ch3 all clinopyroxene have
343 been replaced by amphibole. There are three different Ca-amphiboles: brown edenite or
344 hornblende, brown-green hornblende and green actinolite (Fig. 4). All amphiboles have less
345 than 0.5 Ti apfu, and they range in $Al^{iv} = 0.35 - 1.51$, and $Na_{M4} = 0.04 - 0.52$ (Supplementary
346 Data: Electronic Appendix 2). Brown amphibole is either arranged as overgrowths on
347 destabilised magmatic clinopyroxene (Fig.4-A) or as large (> 0.5 mm) xenoblastic crystals
348 (Fig.4-B). These represent the main high-temperature mineral assemblage successively
349 replaced during subsequent recrystallisation along their edges by a brown-green and then a
350 green amphibole arranged as a corona texture. The break down of the Ti-bearing amphiboles
351 caused the release of Ti and the growth of the titanite and rutile. Maximum TiO_2
352 concentrations in Ch1 amphiboles reach 2.62 to 3 wt% while in Ch3 they reach 3.02 to 3.07
353 wt% (Supplementary Data: Electronic Appendix 2), yielding temperatures between 890 and
354 790 °C (Ch1) and 940-930 °C (Ch3; Supplementary Data: Electronic Appendix 1) (Ottens,
355 1984). These conditions comfortably indicate temperatures typical of granulite facies
356 metamorphism even if the amphibole thermometer may have poor accuracy. A corona of
357 brown-green amphibole develops around these brown amphiboles, which are either
358 hornblende (725 to 600 °C; 0.3 – 1.73 wt% TiO_2) or actinolite (610-570 °C; 0.5 wt% TiO_2),
359 and to which titanite grains are associated with (Fig.4-A-C). Similarly to oceanic gabbros, a
360 partial preservation of primary magmatic minerals towards granulite and amphibolite facies
361 conditions was likely acquired during cooling of these units at the oceanic ridge.

362 *Ti-bearing mineral assemblage.* Titanite is the stable metamorphic Ti-phase in all the
363 amphibole-bearing gabbros, and ilmenite in the metadiorite. Titanite can be found in two
364 types of associations: (i) commonly as small grains (< 20 µm) at the transition between brown
365 and green hornblende (Fig.4-B-C); or (ii) more rarely as hypidioblastic grains (up to 80 µm in
366 size) associated with actinolite in later veins (Fig.4-G, H). Although the former is
367 homogeneous in BSE, the latter shows strong BSE zonation (Fig.4-H). Ilmenite is more
368 common in the ChDi than in the other metagabbros, where grains are often replaced by rutile-
369 titanite association (Fig.4-D-F). In this case, rutile either occurs as coarse grains of up to 25
370 µm or as very fine grains, often showing preferred crystallographic orientation. Both coarse
371 and lamellae rutile are common, easily found in multiple sites in this sample. Rutile lamellae
372 develop at 60°/120° angles with ilmenite (Fig.4-E,F), and they seem to indicate topotactic
373 replacement of ilmenite, such as described by Force et al. (1996), during titanite growth, with
374 excess Ti forming rutile. No evident BSE zonation was identified.

375

376 3.1.3. *Alpine blueschist and eclogite*

377 Nine samples were collected from the Western Alps and Alpine Corsica localities. These units
378 are usually enclosed in HP metapelites and marls.

379 *Blueschist (LT, HP)*

380 Metagabbros - QE-1, QE-6, TR10. These blueschist rocks all located in the Queyras
381 valley preserve their original gabbroic texture, where both pyroxene and plagioclase sites can
382 be recognised despite the development of Alpine foliation. The HP mineral assemblage is
383 given by glaucophane, lawsonite, albite and titanite ± epidote and secondary minerals present
384 are actinolite ± chlorite. Primary magmatic clinopyroxene has been nearly all replaced by
385 glaucophane, although it is still recognisable (Fig. 5-A). The original magmatic plagioclase

386 site is replaced by lawsonite and some epidote and albite aggregates (Fig. 5-A,B). TR10
387 sample is richer in blue-amphibole, with compositions ranging from ferroglaucophane to
388 riebeckite (Supplementary Data: Electronic Appendix 1 and 2), whereas blue amphibole in
389 QE-1 and QE-6 is glaucophane. These amphiboles are variably replaced by actinolite (Fig.5-
390 B).

391 *Ti-bearing mineral assemblage.* In these metagabbros, the stable Ti-bearing mineral is
392 titanite. Titanite is found as large and elongated stripes of xenoblastic crystals (100's μm)
393 between the clinopyroxene and plagioclase sites, and it also occurs as coronas, probably
394 formed at the expense of leached Ca and Ti from former pyroxene and Ca-rich plagioclase
395 (Fig. 5-A). Titanite also occurs together with glaucophane and lawsonite \pm albite.
396 Occasionally, it is possible to identify tiny rutile inclusions ($< 10 \mu\text{m}$ wide) in the core of
397 larger ($> 400 \mu\text{m}$) titanite grains (Fig. 5-C-F), arranged as spotted texture within titanite. In
398 places where only smaller titanite grains are found, fully recrystallised at blueschist
399 conditions, rutile was absent. No evidence of rutile being in equilibrium with other blueschist
400 facies minerals was observed. Ilmenite is also absent and no visible BSE zoning in titanite
401 was identified.

402 *Eclogite-facies rocks (M/LT, HP)*

403 *Metagabbros - VE-29, LS and PG samples.* PG is a metagabbro that still preserves its primary
404 mineralogy, with about 50:50 proportion of clinopyroxene and plagioclase sites. A primary
405 clinopyroxene (augite) is found, with very light brown to greenish colour. The plagioclase site
406 is replaced by quartz, chlorite, zoisite and phengite. Rutile and titanite are accessories and
407 form elongated bands (Fig. 6-A). Amphibole occurs as very small ($<20 \mu\text{m}$) idioblastic grains
408 associated to zoisite and epidote. This seems to occur as thin veins/films infiltrating the
409 metagabbro. Garnet is not observed. This sample have mostly preserved its magmatic texture

410 and although no omphacite was clearly observed, this rock was sampled from the classical
411 locality in Passo Gallarino where other eclogites have been previously sampled. The exact
412 location is given in the Supplementary Data: Electronic Appendix 1.

413 VE-29 was collected in Verne locality and is a metagabbro with a
414 granoporphyroblastic texture composed of omphacite, phengite, garnet, quartz, and rutile \pm
415 ilmenite \pm epidote from the Viso unit (Groppo and Castelli, 2010; Fig. 6-B). Garnet
416 porphyroblasts (up to 500 μm) are almandine-rich and exhibit an increase of Fe and Mg
417 contents towards the rim, similarly to Groppo and Castelli (2010). Garnet includes a
418 significant number of rutile inclusions, many of which are overgrown by a thin rim (<2 μm)
419 of ilmenite. Paragonite and epidote are also frequently found as mineral inclusions.
420 Omphacite is occasionally associated with phengite. These can be altered to chlorite, and
421 some chlorite is also found in garnet fractures.

422 LS was collected in Lago Superiore area and corresponds to a strongly foliated
423 eclogite with granoblastic texture, composed by omphacite, garnet and rutile (Fig. 6-C).
424 Garnet grains are hypidioblastic but of small size (< 500 μm) and frequently strongly zoned,
425 with Ca-richer almandine cores and Mg-richer almandine rims (cores: $\text{Alm}_{55}\text{Grs}_{33}\text{Sps}_{07}\text{Prp}_{05}$;
426 rims: $\text{Alm}_{67}\text{Prp}_{19}\text{Grs}_{13}\text{Sps}_{01}$; Angiboust et al., 2012). In similar eclogite retrieved from Lago
427 Superiore (Angiboust et al., 2012), three omphacite generations, with variable compositions,
428 could be identified.

429 *Ti-bearing mineral assemblage.* In all metagabbroic rocks (PG, VE29, LS), rutile is
430 the stable Ti-phase. Rutile grains can be large (200-300 μm), sub-spherical and with brown-
431 amber colour in plane polarised light. In PG, rutile can be found as coarse, sub-spherical
432 grains in aggregates or individual crystals dispersed in the matrix. Some of these are
433 overgrown by titanite (Fig. 6-A) that was probably formed during decompression from

434 eclogitic conditions (Schwartz et al., 2000). Rutile from VE-29 ranges in size from 20 to 200
435 μm , hypidioblastic individual grains that can be found included in garnet or dispersed in the
436 matrix between omphacite grains. Rutile always displays Fe-exsolution along its cleavage.
437 Less frequently, ilmenite overgrowths are seen. Larger rutile grains which are mostly hosted
438 in the garnet porphyroblasts (Fig. 6-B) often preserve mineral inclusions, such as omphacite,
439 epidote and more rarely apatite. Rutile grains in LS are of variable size (> 300 to $< 100 \mu\text{m}$)
440 and frequently occur as bands of rutile aggregates (Fig. 6-C), aligned along the main foliation,
441 together with garnet and omphacite. These occur as aggregates of $50\text{-}200 \mu\text{m}$, idio- to
442 hypidioblastic grains, forming stripes/bands. These rutile grains have a brown-honey colour.
443 They also occur as inclusions within garnet, but more frequently found along grain
444 boundaries. Ilmenite exsolution bands or overgrowth around rutile are frequent, although
445 these tend to be thin ($< 2 \mu\text{m}$ width; Fig. 6-D). Rutile grains from these metagabbros show no
446 BSE zoning.

447 *Metabasalts - C16 sample.* This sample is equivalent to sample OF3310 in Vitale
448 Brovarone et al. (2011). It corresponds to a foliated lawsonite-bearing “blueschist” eclogite
449 (Fig. 6-E) with an associated omphacite- and garnet-bearing vein (C16v; Fig. 6-F). The
450 sample paragenesis is glaucophane, lawsonite, quartz, albite, titanite, and Fe-Ti oxides (Fig.
451 6-E). The vein, which is ca. 1 cm wide; contains omphacite, garnet, lawsonite, glaucophane,
452 albite and titanite (Fig. 6-F). Both parts reached the same P-T conditions, but paragenesis is
453 dependent on the bulk composition of the basalt and the vein (Vitale Brovarone et al., 2011).
454 P-T conditions are constrained by the lawsonite-blueschist eclogite from Vitale Brovarone et
455 al. (2011) yielding $520 \pm 20 \text{ }^\circ\text{C}$ and $23 \pm 1 \text{ kbar}$. *C147 sample.* This sample is equivalent to
456 OF3158 in Vitale Brovarone et al. (2011) and corresponds to a lawsonite-bearing eclogite
457 (Fig. 6-G), with occasionally preserved magmatic textures. The HP paragenesis is garnet,
458 omphacite, lawsonite, glaucophane, titanite and late Fe-Ti oxides. Garnet can be idio- to

459 hypidioblastic, usually small in grain size (<500 μm). P-T conditions are similar to those of
460 C16 (Table 1).

461 *Ti-bearing mineral assemblage.* In C16, only titanite is found as the main Ti-host
462 phase. Titanite grains are abundant, hypidioblastic, but small (<20 μm), and become more
463 frequent and larger in the vein that crosscuts the basalt (Fig. 6-F). In C147, titanite is the only
464 Ti-mineral phase, occurring as small grain size idioblastic grains (< 20 μm) and abundant
465 (Fig. 6-H). Titanite grains show no BSE zoning. Hematite is found as a product of sulphide
466 oxidation. Titanite grains in C147 are in clear equilibrium with the mineral assemblage at
467 lawsonite eclogite facies conditions. This contrasts with the Monviso eclogites, where rutile is
468 the HP stable Ti-phase and occasionally titanite is only found as a retrograde phase (PG).

469 **3.2. TiO₂ polymorphism**

470 IODP and Chenaillet samples yielding what was recognised petrographically as rutile were
471 subjected to an inspection using Raman spectroscopy to confirm the polymorphism. Grains
472 from 4 IODP samples (1316, 1319, 1320, and 5450) and from 3 Chenaillet samples (Ch1, Ch3
473 and ChDi) were analysed. The Raman spectrum for each of these measurements was
474 compared with spectra of rutile, anatase and brookite from the RRUF database (Figs. 7-9 in
475 Supplementary Data: Electronic Appendix 1). Without exception, for both coarser grains and
476 finer lamellae, rutile was always the TiO₂ mineral identified.

477 A summary of P-T conditions associated with stable Ti-phase in each sample, can be
478 found in Figure 7. The larger T range in the IODP and Chenaillet samples is related to the
479 uncertainty in the conditions of Ti-phase growth in those rocks.

480 **3.3. Ti-phase mineral chemistry**

481 3.3.1. *Titanite*

482 *Major and minor elements.*

483 Titanite from the Atlantic IODP samples (1316, 1319, 1320) shows lower X_{Al} , between 0.026
484 and 0.045, and F ($F_{pfu} < 0.003$), with respect to those of Chenaillet titanite, that yield X_{Al}
485 between 0.066 and 0.077, and F_{pfu} between 0.003 and 0.007. On the other hand, titanite grains
486 retrieved from the Indian IODP (5450) sample exhibit higher values, with X_{Al} of about 0.106
487 and F_{pfu} of about 0.026. Titanite formed at higher pressure, from the blueschist and eclogite
488 metagabbros, are homogeneous with X_{Al} between 0.022 and 0.033 and low F concentrations
489 with $F_{pfu} < 0.005$ (Fig. 8).

490 In a X_{Al} vs F_{pfu} diagram, published data of titanite chemistry from mafic eclogite,
491 carbonate-rich eclogite, from UHP, HT calcisilicate, upper amphibolite calcisilicate, high-
492 temperature marbles, felsic and mafic amphibolite rocks are also reported. The upper
493 amphibolite calcisilicate rocks and the marbles present the highest X_{Al} and F_{pfu} . These are
494 followed by the felsic upper amphibolite units and UHP-upper amphibolite calcisilicate
495 titanite. Our data plot at slightly lower F_{pfu} than the literature dataset reported in Figure 8.
496 LT-HP metamorphic titanite show the lowest X_{Al} and can be separated from MT-LP
497 metamorphic titanite in this dataset based on X_{Al} but not F_{pfu} . It should be noted that titanite
498 found in the granulite oceanic samples show the lowest F values (Fig. 8). The lack of mineral
499 chemistry data from more mafic protolith titanite prevents a more robust comparison.

500 *High Field Strength Elements, V and Sn.*

501 Titanite grains from the IODP amphibole-bearing gabbros yield average Nb/Ta values around
502 16.8 ± 9 (Fig. 9a). This contrasts to the extreme, and variable Nb/Ta ratios in titanite grains
503 from the pre-Alpine Chenaillet ophiolite (Ch3; Nb/Ta 16.2 – 208). The Chenaillet ophiolite
504 seems to have been strongly affected by an ocean-floor, late-stage fluid circulation (*e.g.*
505 Nicollet et al., 2022). Some titanite found in thin veins show increased Nb/Ta values in
506 comparison to other titanite grains found in amphibole coronas (average Nb/Ta 128 vs 47,

507 respectively). However, even corona-related titanite in the Chenaillet yields higher Nb/Ta
508 values than titanite in IODP amphibole-bearing gabbros (Fig. 9a). Unlike Nb/Ta, Zr/Hf values
509 are less variable across all HT samples (Fig. 8a). For instance, Chenaillet titanite yields
510 average Zr/Hf of ca. 29 ± 11 , in good agreement with Zr/Hf from IODP titanite (28 ± 15).
511 Finally, blueschist (QE-1, QE-6, TR10, C16) titanite grains yield average Nb/Ta 17.1 ± 13 (6
512 – 67) and Zr/Hf 19.8 ± 9 (9 – 45). Overall, the majority of titanite grains have average
513 chondritic Nb/Ta but lower than chondritic Zr/Hf values (Fig. 9a).

514 Figure 9b depicts our titanite data in a V vs Sn diagram (Xiao et al., 2020). The
515 majority of titanite grains yields Sn content < 40 ppm, irrespective of metamorphic grade, and
516 only a few titanite crystals (4 grains) show higher Sn content (40-100 ppm). Vanadium
517 compositions of all titanite are more variable and can go as high as 2500 ppm (Fig. 9b).
518 Titanite grains from MT-LP conditions show a wide range of Sn compositions: i) IODP1320
519 sample yield the lowest Sn and V values (< 10 and < 500 ppm, respectively), while ii)
520 IODP1319 and Ch3 yield higher Sn content (10 – 100 ppm) but similar V chemistry. In the
521 Queyras blueschist, Sn is within uncertainty of granulite-facies titanite grains but have
522 variable and higher V contents (up to 1500-2500 ppm). This is clearly at odds from titanite
523 grains from Corsica (LT eclogite) that yield narrower V (< 500 ppm) and Sn (10-30 ppm)
524 compositions (Fig. 9b).

525 *Yttrium, Rare Earth Elements and Th/U ratios.* In the metagabbros, titanite grains
526 yield highly variable Y and REE compositions, from more enriched to depleted, from strongly
527 fractionated to unfractionated (Fig. 10).

528 Titanite can incorporate significant amounts of Y (Scibiorski et al., 2019). LP titanite
529 yields two groups based on Y (Supplementary Data: Electronic Appendix 3), one yielding a
530 narrow concentration range (sample 1320: 1 – 252 ppm; and from Chenaillet Massif Ch3: 5 –

531 222 ppm), while the other shows more elevated and variable Y contents (samples 1316 and
532 1319: 152 to 4095 ppm). HP titanite (QE-6, TR10, and C16) range between 70 and 500 ppm
533 Y content.

534 Titanite grains that form under LP conditions (Ch3 and IODP1319 amphibolites) show
535 a wide range of REE contents ($\Sigma\text{REE} = 10 - 3350$ ppm) and flat REE patterns ($\text{La}/\text{Yb}_\text{N} = 0.02$
536 $- 1.67$)(Fig. 10). Eu anomalies on these units are highly variable ($\text{Eu}/\text{Eu}^* = 0.22 - 29.5$).
537 Titanite grains from IODP1320 sample have a similar range of REE contents ($\Sigma\text{REE} = 12 -$
538 2006 ppm), but show different chondrite-normalised REE patterns (Fig. 10), with LREE
539 enriched patterns ($\text{La}/\text{Yb}_\text{N} = 11 - 113$), ($\text{La}/\text{Sm}_\text{N} = 5 - 93$) and clear positive Eu anomalies
540 ($\text{Eu}/\text{Eu}^* = 4.7 - 314$). Th/U ratios from all these LP rocks are commonly < 1.1 except for one
541 outlier. They show either small scatter within sample (Ch3 and 1319 $\text{Th}/\text{U} = 0.35 \pm 0.1$) or
542 some variation (sample 1320; $\text{Th}/\text{U} = 0.08 - 1.06$).

543 Titanite grains retrieved from HP metagabbros and metabasalts show low REE
544 contents ($\Sigma\text{REE} = 32 - 465$ ppm), of which ΣLREE represent 4 to 210 ppm. Blueschist rocks
545 from Queyras (QE-1, QE-6, TR10) show consistent depleted titanite LREE pattern ($\text{La}/\text{Sm}_\text{N} =$
546 $0.02 - 0.7$), similarly to eclogite rocks from Corsica ($\text{La}/\text{Sm}_\text{N} = 0.02 - 0.3$). Eu anomalies are
547 non-existent to only slightly positive ($\text{Eu}/\text{Eu}^* = 0.84 - 1.8$). Th/U ratios show small scatter
548 within the HP samples, commonly < 0.4 . These features are summarised in Table 3.

549 3.3.2. Rutile

550 Minor and trace element composition of rutile grains from two IODP gabbros, from eclogite
551 and as inclusions within titanite from blueschist rocks are reported in Figure 11. Chromium
552 and Nb contents are generally low (< 600 ppm), with only some rutile grains exhibiting more
553 elevated Nb (800 - 1500 ppm) and Cr (800 - 1400 ppm) (Fig. 11a; IODP gabbros). In this
554 way, almost all rutile plot in the *metamafic field* defined by Meinhold et al. (2008), and only a
555 few IODP grains plot within the *metafelsic field*. In the diagram Zr/Hf vs Nb/Ta (Fig. 11b),

556 the majority of the analyses display Zr/Hf (20 – 40) and Nb/Ta (12 – 30) values in the
557 *metamorphic field* defined by Pereira et al. (2019). A limited number of rutile grains display
558 lower ratios, typical of rutile precipitated from hydrothermal-related fluids (Fig. 11b), that
559 plot within the metamafic in Figure 11a.

560 A ternary diagram constructed using V+Cr, Sc and Nb as endmembers to explore the
561 variance in our dataset (as determined by a PCA analysis) is presented in Figure 11c. Rutile
562 grains from eclogite plot in a field closer to higher V and Cr, unlike rutile grains that formed
563 under LP conditions (IODP gabbros or inclusions in blueschist titanite; Fig. 11c). The latter
564 plots closer to Sc, while the former has variable Sc-Nb. In Figure 10d we present the sum of
565 the HFSE ratios against the Nb/V ratio. Rutile in eclogite display high $\sum\text{Zr/Hf}$, Nb/Ta (>25)
566 and low Nb/V (<0.3), while low pressure granulite to upper amphibolite rutile show higher
567 Nb/V (>0.3). Using previously defined fields shown in Figure 11b, rutile grains with low
568 $\sum\text{Zr/Hf}$, Nb/Ta (<25) are likely related to fluid precipitation processes (Pereira et al., 2019).

569 As shown in section 3.1, ilmenite in amphibole-bearing gabbros (IODP) often display
570 partial replacement by titanite and rutile. These are also the samples where rutile chemistry is
571 less homogeneous. In these composite grains, rutile is found such as the grain depicted in
572 Figure 12, either as very fine lamellae, frequently displaying 120°/60° intersections, or as
573 coarser grains, mostly frequently rimmed by titanite (example spot 4 and 9* in Fig 12a), but
574 also just by ilmenite (example spot 7 in Fig 12a).

575 A diagram of Cr against Zr values is shown in Figure 12b, where two distinct rutile
576 groups are highlighted: (i) grains that are low in Cr (< 600 ppm) and low in Zr (< 300 ppm),
577 and (ii) grains with Cr content above 600 ppm with higher Zr content (500-2500 ppm). The
578 most enriched analysis (7 in Fig. 11) corresponds to a rutile grain surrounded by ilmenite as
579 an inclusion (Fig 12a). The analyses with the lowest Zr, Cr, and V are fine-grained rutile,

580 commonly closer to the former edges of the ilmenite grain (Fig 12a, spots 10, 11, 12, and 1).
581 Yet, the relation between the enrichment in Cr-Nb-Zr and the rutile textural position within
582 the composite grain (lamellae or inclusion) is not perfectly systematic.

583 3.3.3. *Ilmenite*

584 The chemical composition of ilmenite was investigated using EPMA and LA-ICP-MS,
585 focusing on two samples from the IODP drill cores, 1320 and 5450, which contained the
586 largest ilmenite grains. These ilmenite grains yield variable Fe₂O₃ (4.7 and 17.6 wt%) and
587 MgO (12.6 and 4.7 wt%) compositions, similar MnO (0.22 and 0.25 wt%) and TiO₂ (3.3 and
588 4.9 wt%) and overall low Al₂O₃ (120 to 365 ppm). In sample 1320 (higher bulk Al and Mg),
589 ilmenite incorporates more Al (341 ppm) and Mg (4.1 wt%) than in sample 5450 (Al 134 ppm
590 and Mg 0.5 wt%). A small variation in WR MnO abundance ($\Delta \sim 0.03$ wt%) seems to reflect
591 highly on Mn incorporation into ilmenite (4 369 to 8 289 ppm; Supplementary data:
592 Electronic Appendix 2).

593 To investigate WR dependency of trace elements in ilmenite, we use Al against other
594 trace elements (Mn yields similar results; Fig 13). Vanadium and Zr inversely correlate with
595 Al, while Cr correlates positively. This seems to imply that at higher Al₂O₃ compositions in a
596 gabbro, ilmenite incorporates less V and Zr (Fig. 13). Conversely, more Cr will be
597 incorporated by ilmenite in a more enriched Al₂O₃ composition. Incorporation of Yb, Nb and
598 Ta by ilmenite does not correlate with Al (nor Mn). Particularly for one sample (IODP1320;
599 Fig. 13) there is either clustering (Yb and Ta) or scatter (Nb) for similar Al values in ilmenite
600 and for the same bulk Al₂O₃. This can indicate the presence of two distinct ilmenite
601 populations.

602 4. DISCUSSION

603 4.1 Ti-phase stability: from oceanisation to subduction

604 4.1.1 *Ti-minerals in ocean floor amphibole-bearing gabbros (LP-MHT)*

605 Our data show that ilmenite, titanite, and rutile are present at very low-pressure, and moderate
606 to high temperature (amphibolite to granulite) conditions in the oceanic crust. Their formation
607 history can be deduced from their petrographic textures and chemical signatures. Ilmenite can
608 be a product of magmatic and/or later metamorphic recrystallisation, as it commonly forms
609 the core of ilmenite-titanite composite grains. Fine rutile-titanite lamellae showing 120°/60°
610 angles indicate an exsolution of excess Ti during titanite overgrowth on ilmenite at lower
611 grade. Rutile found as coarser grains (10-40 µm; *e.g.* sample IODP1316) is more puzzling. At
612 times, coarse rutile is found with no ilmenite surrounded by a titanite corona (*e.g.* IODP1316),
613 while other rutile grains are found surrounded/intergrown by/with ilmenite and then by a
614 titanite corona (IODP1320; Fig. 3). The former is suggestive of an earlier origin of rutile in
615 the gabbro, either HT metamorphic or a late-stage magmatic crystallisation related to
616 fluctuations in the magma fO_2 (Toplis and Carroll, 1995) enhanced by fluid/melt percolation
617 (Tribuzio *et al.*, 2000). While texturally these could be envisaged as magmatic, this would
618 require magmas with elevated TiO_2 (7-10 wt%; Ryerson and Watson, 1987), such as in more
619 fractionated magmas in the oceanic crust (Whitaker *et al.*, 2007). The second is more difficult
620 to interpret, as rutile can be potentially coeval with ilmenite recrystallisation or interpreted as
621 an earlier phase. Rutile (lamellae, coarse grain) found in IODP samples, can also be
622 tentatively distinguished by their chemistry: one group exhibits slightly higher Nb-Zr-Cr
623 compositions, potentially formed at higher temperatures than the other group that yields lower
624 Nb-Zr-Cr, corresponding strictly to the fine lamellae formed during titanite-rutile replacement
625 of ilmenite (Fig.12). This replacement process generally takes place at amphibolite-facies

626 conditions (ca. low temperature ~600 °C). At such temperature, amphibole does not
627 incorporate Ti released by ilmenite breakdown, and thus rutile forms within titanite.

628 Titanite in the IODP and Chenaillet metagabbros is texturally associated with the
629 breakdown of brown to green amphibole, or occurs as replacement textures after ilmenite,
630 under amphibolite facies conditions. A rare third type of titanite exhibiting strong chemical
631 zoning (Fig. 4-H), formed at a later stage during fluid percolation associated with actinolite at
632 greenschist facies cooling conditions, is well documented in Ch1 – Chenaillet ophiolite
633 (Nicollet et al., 2022).

634 In summary, Ti-phases in oceanic gabbros evolve from magmatic ilmenite at HT
635 conditions to low temperature titanite by hydrothermal metamorphism during cooling, while
636 rutile potentially crystallise during early and late ocean floor metamorphism due to local or
637 global Ti excess.

638 4.1.2 *Ti-phase stabilities in the subducting slab*

639 As the oceanic crust undergoes subduction, new minerals stabilise to higher pressures and
640 temperatures. Here, metagabbros from the Queyras-Monviso and Corsica preserve mineral
641 assemblages equilibrated at blueschist and eclogite facies conditions. Overall, titanite appears
642 stable at LT-HP, except in eclogite samples from Monviso. Typically, titanite grains are found
643 in equilibrium with lawsonite in the Queyras blueschist and Corsica eclogite. They have small
644 grain sizes and are idioblastic in the metabasalt (Fig. 6), while often xenoblastic (Fig. 5) or in
645 small aggregates (Fig. 5) in the metagabbros. Occasionally in the blueschist from Queyras
646 (metagabbros), titanite shows porous texture, with frequent but very small rutile grains. We
647 interpret these as inclusions, as they only occur in the cores of larger titanite and never where
648 titanite grains are fine-grained, and thus better equilibrated. These inclusions are often
649 lamellar, with the preserved 120°/60° angles. These textures are very similar to those found in
650 the oceanic metagabbros. No ilmenite is found in these titanite grains. These textures are

651 suggestive of rutile being a relic phase (such as magmatic clinopyroxene in the same rocks)
652 preserved within the recrystallised titanite grains, as testimonies of the ocean-floor
653 metamorphic evolution.

654 Rutile is stable in Monviso eclogite and titanite is absent, except for retrograde titanite
655 rimming rutile in the Passo Gallarino eclogite. At similar P-T conditions, titanite is found
656 stable together with lawsonite and omphacite in Corsica samples and rutile is found stable
657 with garnet, omphacite and zoisite in Monviso. A study has argued that titanite reacts to form
658 rutile tied with lawsonite disappearance to form epidote (Frost *et al.*, 2001), which potentially
659 could explain the differences between these eclogite samples.

660 *Lawsonite and titanite stabilities.* The Corsican lawsonite eclogite is one of the few
661 occurrences reported in the literature where titanite is found in equilibrium at eclogite facies
662 (*e.g.* Vitale-Brovarone *et al.*, 2011; Sassi *et al.*, 2000; Ng *et al.*, 2016). Frost *et al.* (2001)
663 explains that the dehydration reaction of lawsonite producing zoisite and quartz leads to
664 titanite breakdown and rutile formation, as zoisite incorporates more Ca than lawsonite. In
665 detail, this reaction is sensitive to water activity (Clarke *et al.*, 2006), bulk composition (Davis
666 and Whitney, 2006, SiO₂-Al₂O₃) and CO₂ in fluids (Castelli and Rubatto, 2002). Previous
667 authors showed that this reaction yields a positive Clapeyron slope, taking place at higher
668 pressure with increased temperature, until about 500 to 600 °C, and then it becomes mainly
669 temperature-dependent (Clarke *et al.*, 2006; Whitney and Davis, 2006). Perhaps as a
670 consequence, there have been few reported occurrences of rutile stable with lawsonite under
671 blueschist or eclogite conditions (*e.g.* Bianco *et al.*, 2019 in the Elba Island – Italy, Maruyama
672 *et al.*, 1998 in the Franciscan Complex – USA, and Groppo and Castelli, 2010 in Monviso -
673 W. Alps). This implies that while the net-reactions might influence what components are
674 available to stabilise titanite or rutile, rutile can appear while lawsonite is yet stable, and

675 therefore the transition between HP titanite to rutile is an interplay with other factors such as
676 major oxides, H₂O, and/or CO₂.

677 4.1.3 *The effects of MORB whole rock composition on titanite and rutile stabilities*

678 It has been well established that for rocks with MORB composition, rutile is more stable at
679 high pressure conditions, while titanite and ilmenite at lower pressures, with ilmenite more
680 stable than titanite at higher temperatures (e.g. Liou et al., 1998). Yet, due to natural
681 exceptions to these phase relationships, in the following we explore other factors that play a
682 role in their stability, with a special focus on bulk rock chemistry. The stability of rutile, for
683 example, is mostly influenced by the bulk TiO₂ concentration, whereas titanite is more
684 sensitive to variations in overall bulk geochemistry (Frost et al., 2001; Angiboust and Harlov,
685 2017). In Ca-rich protoliths, the following reactions maximise titanite stability: rutile +
686 grossular + diopside ↔ pyrope + titanite (1) and rutile + quartz + calcite ↔ titanite + CO₂ (2)
687 (Spear, 1993). The presence of carbonate minerals favour titanite (Cave *et al.*, 2015). The
688 water activity at HP is also an important parameter, as H₂O will favour hydrous minerals and
689 thus titanite as indicated by the following reaction: lawsonite + titanite ↔ zoisite + rutile +
690 quartz + H₂O (3) (Frost *et al.*, 2001). Indeed, the Corsican eclogite is richer in H₂O than the
691 Monviso eclogite (Supplementary Data: Appendix 1), where one yields titanite and the other
692 rutile. Yet, as also seen in reaction (2), CO₂-rich fluids favour rutile instead of titanite. It has
693 also been shown that the F-Al coupled substitutions seems to stabilise titanite above its
694 commonly known stability conditions, more so in Ca-rich protoliths (marbles), at HT or HP
695 (Castelli and Rubatto, 2002; Tropper *et al.*, 2002; Fig. 8). While titanite tends to disappear
696 with increasing SiO₂ (Frost *et al.*, 2001), this is mostly linked to the concurring decrease of
697 CaO. Thus, it has long been established that the Ca/Al ratio is key in the stabilisation of
698 titanite (Frost *et al.*, 2001). Reflecting on the considerations above, the presence or absence of
699 Ti-phases are shown in a plot comparing Ti/Ca and Ca/Al bulk rock values, using data from

700 our LT-HP samples combined with Syros HP rocks, and average MORB published data (Fig.
701 14). Ti/Ca values are expected to influence rutile stability while Ca/Al should mostly
702 influence titanite stability as it competes with plagioclase, lawsonite or epidote. With
703 increasing Ca/Al, even at elevated Ti/Ca, we expect titanite to be stable, leading to a field
704 where both rutile and titanite should be stable (Ca/Al values above ~ 0.85). It should be noted
705 that the compositional range of this set of rocks is limited: bulk SiO₂ (46-51 wt%) and CaO
706 (9-14 wt%), and TiO₂ ranges from 0.25 to 3.17 wt% (Supplementary Data: Electronic
707 Appendix 1 – Table 5) and that no calcite was observed or reported.

708 Several conclusions can be drawn from the figure:

- 709 – Blueschist and eclogite rocks with stable titanite yield variable Ca/Al (commonly higher
710 than 0.75) while Ti/Ca is systematically below 0.15. On the contrary, blueschist and
711 eclogite with stable rutile exhibit Ti/Ca values above 0.15. These data seem to support the
712 major control of reaction (1) on rutile and titanite stability at LT-HP conditions. This is
713 consistent with data on blueschist and eclogite from Syros (Marschall 2005) for which
714 rutile-bearing HP units plot in the Ti/Ca > 0.15 field.
- 715 – Unaltered MORB values yield relatively low Ti/Ca values (0.05 - 0.15) and variable but
716 high Ca/Al that range from 0.8 to 1.2 (Fig. 14; Klein, 2004). These compositions at low
717 geothermal gradients favour titanite over rutile growth. However, a relative increase of Ti
718 and Fe in respect to the other major elements is expected for the altered oceanic crust
719 metamorphosed by either hydrothermal or seawater fluid interactions (Staudigel and Hart,
720 1983). Al-saponite type alteration of Fe-bearing minerals (e.g, oxides and
721 clinopyroxenes) and albitisation of primary plagioclase can also lead to considerable Ca-
722 loss (Alt *et al.*, 1992), and thus alteration of MORB rocks are likely to increase Ti/Ca and
723 variably decrease Ca/Al values.

724 4.2. Controls on Ti-bearing mineral chemistry

725 4.2.1. Ilmenite trace element compositions: bulk-rock effect

726 Our limited data show that ilmenite major and minor element compositions are
727 strongly dependent on bulk-rock composition. The binary plots of Al against multiple trace
728 element contents in ilmenite illustrate how Cr seems to be positively correlated with Al,
729 which in turn correlates with the Al_2O_3 content of the metagabbros (Fig. 13). Niobium and Ta
730 showed no significant correlation with Al, indicating the presence of two chemically distinct
731 ilmenite populations in each sample. This implies that other factors are more relevant in the
732 intake of Nb and Ta by ilmenite, such as competition with rutile (*e.g.*, sample 1320).

733 4.2.2. Titanite trace element compositions: the role of fluids and HREE-bearing minerals

734 *Titanite from LP settings.* Titanite from our dataset show Nb/Ta values that are within
735 uncertainty of the chondritic value (Fig. 9a; McDonough and Sun, 1995), except a few titanite
736 from the Chenaillet Massif amphibolite metagabbro that yield Nb/Ta > 100 (Fig. 9a). Titanite
737 Zr/Hf values, on the other hand, are persistently subchondritic. The high-field strength
738 elements (HFSE) pairs, Zr and Hf, and Nb and Ta have a degree of geochemical affinity that
739 prevents their fractionation under most geological conditions. However, some fractionation
740 between each pair in metamorphic titanite has previously been reported (Chen and Zheng,
741 2015). Cherniak (2015) has constrained the diffusivities of these elements in titanite and her
742 data suggest that unlike in rutile, diffusive fractionation should not be a likely mechanism to
743 explain variable titanite Nb/Ta ratios. Therefore, fractionation between these elements in
744 titanite is promoted by other processes such as growth in the presence of external fluids or
745 after the break down of minerals with variable compatibility with these elements. In this way,
746 we interpret these higher to extreme Nb/Ta values found in the Chenaillet Massif to be
747 derived from external fluids related with differentiation of the lower oceanic crust, during
748 cooling. Additionally, some of the Chenaillet Massif titanite grains (Ch3) can also yield

749 higher Sn (> 60 ppm; Fig. 9b). Elements like Sn and Mo are usually associated with titanite
750 growth from hydrothermal fluids (such as in Fu et al., 2016). This may support the existence
751 of externally derived late-stage fluids during titanite growth. These data are compatible with
752 the late-stage texture of titanite in these rocks, associated with brown to green amphibole
753 breakdown and alteration of former ilmenite grains.

754 *Y-REE behaviour.* Titanite mineral chemistry has been heavily addressed over the
755 years, as its chemistry seems to be a faithful recorder of petrogenetic processes (Bruand *et al.*,
756 2014, 2016; Scibiorski *et al.*, 2019). Titanite incorporates a significant amount of REE and Y,
757 of which Y and HREE are also incorporated by major mineral garnet. Scibiorski *et al.* (2019)
758 have demonstrated how titanite growing in competition with garnet, in garnet amphibolites
759 and gneisses (both felsic and intermediate), is significantly affected in key HREE. In their
760 study, they demonstrated that by using Y vs Dy/Yb titanite grown in garnet-absent rocks can
761 be distinguished from garnet-bearing rocks. We have plotted titanite data from our study in
762 the same diagrams (Fig. 15). Except for the Corsica eclogite, all our titanite-bearing samples
763 are garnet-free. Commonly Y substitutes with HREE and thus there is a linear correlation
764 between HREE and Y values (Fig. 15a). Titanite grains retrieved from Corsica eclogite-facies
765 rocks have a different Y/HREE slope than titanite from the other units (Fig. 15a). While in the
766 original study by Scibiorski *et al.* (2019), Dy/Yb faithfully discriminated titanite of garnet-
767 bearing from garnet-free domains, some of our titanite from garnet-absent units (e.g. Ch3)
768 yield high Dy/Yb (> 2.0 ; Fig. 15b). This suggests that i) there is competition with another
769 mineral for Yb, such as zircon and/or xenotime, ii) that titanite might incorporate more Dy
770 due to break down of a main phase that favours MREE such as zoisite (Frei *et al.*, 2004) or iii)
771 that at HP-LT conditions, Yb partition coefficient in titanite changes.

772 4.2.3. Rutile trace element compositions: Cr and High Field Strength Elements

773 *Cr-Nb discrimination diagram.* Rutile from our dataset (n=69) are derived from
774 metagabbros (mafic). We applied the Cr-Nb protolith discrimination diagram from Meinhold
775 *et al.* (2008) and while the majority of rutile plot in the metamafic field, there are 7 grains
776 classified as metafelsic-derived. These correspond to grains retrieved from the IODP
777 amphibole-bearing gabbros (samples 1316 and 1320). These results confirm that rutile
778 derived from high-grade conditions can be misidentified using this discrimination tool, such
779 as highlighted by Meinhold *et al.* (2008). This suggests that at high-grade conditions, Cr is not
780 always preferentially retained by rutile, as Cr is more diffusive than Nb at HT (Sasaki *et al.*,
781 1985; Marschall *et al.*, 2013) or as a new source of Nb becomes available (*e.g.* amphibole or
782 ilmenite breakdown). This seems to be the case, due to the incongruent replacement of
783 ilmenite in the ilmenite-titanite-rutile aggregates. It is not fully clear whether rutile and
784 ilmenite are initially stable together, as textural relationships seem to suggest, and thus
785 competing for Nb. No evidence of HT amphibole breakdown was identified in sample 1320,
786 and this can explain Cr being sequestered by either HT amphibole or by primary
787 clinopyroxene (such as in Tribuzio *et al.*, 2019).

788 *High Field Strength Elemental ratios.* Zr/Hf and Nb/Ta values from rutile plot within
789 known compositions of other metamorphic rutile, except for 4 grains that have $\sum\text{Nb/Ta}$, Zr/Hf
790 < 25 , which plot within the fluid-related rutile field of Pereira *et al.* (2019). These grains
791 come from sample IODP1319, a LP amphibolite, where rutile occurs as spherical inclusions
792 always surrounded by ilmenite. As discussed in the previous section, these rutile grains seem
793 to reflect a higher temperature growth, in respect with the fine rutile lamellae, as they occur as
794 inclusions in ilmenite. Yet, these grains yield the lowest values for several trace elements
795 within the IODP dataset (Cr < 11 , Zr < 204 , Nb < 450 ppm). This might imply that these

796 grains have strongly fractionated Nb/Ta and Zr/Hf values, typical of LT-LP rutile. The low
797 Zr/Hf values (7.22 - 13.66) are consistent with rutile formed at LT (Ewing *et al.*, 2014).

798 **4.3. Are titanite and rutile chemistries pressure and temperature sensitive?**

799 *4.3.1. Titanite*

800 We filtered our titanite chemical data using Principal Component Analysis. This allowed to
801 select certain trace elements that better explain the variance in the dataset, and plot them in a
802 Yb, V+Cr and Nb+Ta ternary diagram (Fig. 16). Titanite grains from our LP oceanic
803 amphibole-bearing gabbros plot closer to the Nb+Ta endmember, while titanite belonging to
804 the HP assemblage from HP metagabbros plot between the Yb and the V+Cr endmembers
805 (Fig. 16a). Similarly, La/Sm, V and Yb allow to distinguish LP from HP titanite, as HP
806 titanite are strongly depleted in LREE relatively to the MREE (Fig. 16c). To test these
807 systematics, we compiled titanite trace element data from igneous rocks, gneisses (ortho-
808 mafic and intermediate, and calc-rich), and HP-LT rocks from the literature (Fig. 16b, d).

809 *Nb+Ta vs Yb.* Igneous-related titanite (either more mafic or more felsic) is clearly
810 distinguishable from other titanite and plot closer to the Nb+Ta apex in the ternary diagram
811 (Fig. 16b). HT metamorphic titanite grains show similar behaviour, despite some outliers in
812 the garnet amphibolite that tend towards the Cr+V endmember. Blueschist-derived titanite
813 from the literature plot mostly in the centre of the diagram and tend to the Yb endmember (Nb
814 = 10 to 30 ppm; Yb = 15 to 40 ppm) and are easily distinguished from the remaining titanite
815 (igneous and HT). Yet, a few eclogite titanite grains from our study and other blueschist
816 titanite (New Caledonia) plot closer to the Nb+Ta endmember (Fig. 16b). This implies that
817 another mechanism is accountable for those titanite compositions. Eclogite from Corsica are
818 often garnet bearing (C147), and therefore those titanite were in competition for HREE,
819 explaining their lower Yb. The New Caledonia titanite form in garnet-free, metamafic and
820 metapelitic blueschist (Spandler *et al.*, 2003) and thus another explanation is required for their

821 Yb depletion when compared with the other blueschist-related titanite grains (8 to 30 ppm vs
822 20 to 50 ppm, respectively). There may be: i) competition for Yb with other minerals (such as
823 apatite and lawsonite; Spandler et al., 2003), scavenging it before titanite growth, or ii) a
824 relatively low Yb bulk composition (Spandler et al., 2003). In summary, although a ternary
825 diagram such as presented in Fig.16b discriminate a large part of our LT-HP titanite dataset
826 from other titanite, this diagram fails to discriminate all LT-HP reported in this figure and
827 therefore a few other competing factors can influence these trace-elements distribution into
828 titanite.

829 *V content in HP titanite.* Some titanite grains in our dataset show significant V
830 contents (> 500 ppm), markedly in blueschist titanite from the Queyras. Minerals like
831 pargasite, phengite, omphacite or epidote are major carriers of V. Indeed, titanite growth after
832 biotite or hornblende breakdown has been shown to be more enriched in V (Xiao et al., 2020).
833 Blueschist from Queyras are predominantly lawsonite and glaucophane bearing, minerals that
834 do not particularly favour V incorporation (Spandler et al., 2003). This might explain why V
835 is more incorporated by titanite in these samples. Conversely, in Corsica samples, titanite is
836 equilibrated with omphacite who disputes more with titanite for V (Hermann, 2002; Spandler
837 et al., 2003). This seems to suggest that in the absence of omphacite or phengite, HP titanite
838 grains can yield higher V contents, but the role of fO_2 should also be considered as V is redox
839 sensitive.

840 *REE contents, and the La/Sm vs Yb vs V ternary.* It must be noted that on average,
841 REE contents of igneous titanite are high, particularly for the LREE (Xiao et al., 2014;
842 Bruand et al., 2020; Scibiorski and Cawood, 2022) and therefore these can be used as a first
843 approach to distinguish igneous from metamorphic titanite. Titanite from our LP amphibolite
844 facies samples go from depleted to mildly enriched REE, with no significant REE
845 fractionation patterns. On the other hand, LT-HP blueschist and eclogite titanite grains show

846 strong depletion in REE and more particularly in LREE, which results in La/Sm_N
847 fractionation and in negative slopes of the LREE pattern (Fig. 10). Lanthanum is mostly taken
848 by lawsonite and epidote in mafic blueschist, and by apatite in more pelitic compositions
849 (Spandler *et al.*, 2003; Martin *et al.*, 2014). However, epidote does not intake Sm (Spandler *et*
850 *al.*, 2003). This implies that epidote can lead to variable La and Sm availability and thus
851 fractionation of these REE. We have thus used La/Sm_N against Yb (HREE) to investigate their
852 potential as pressure discriminants applied to titanite from our study (Fig. 16c). Alpine HP
853 units plot at lower La/Sm_N values and spread in between Yb and V endmembers. This is in
854 sharp contrast with titanite grains from igneous and lower amphibolite conditions that tend to
855 display no to small La/Sm_N fractionation and therefore plot close to the La/Sm_N apex (Fig.
856 16d). By applying a multiple-step discrimination approach, resorting to the two ternary
857 diagrams (Fig. 16b and d), different titanite types formed at variable settings and grades can
858 be distinguished.

859 4.3.2. Rutile

860 Previously, we showed that the variability in our rutile dataset can be best described using
861 V+Cr, Sc, and Nb endmembers in a ternary diagram (Fig. 11c). Based on this diagram, rutile
862 from eclogite can be distinguished from amphibolitic-granulitic rutile. Available rutile data
863 from the literature are also reported in two diagrams (Fig. 17). Despite the limited amount of
864 mafic rutile data available from the literature reporting Sc, data from the literature agree well
865 with the defined fields from this study. HP rutile from Syros and from a talc eclogite from
866 Monviso plot together with eclogite rutile from our study in the HP field, whilst rutile from a
867 mafic granulite plot together with LP rutile from our study in the newly defined LP fields of
868 both diagrams (Fig. 17). Some rutile grains from the talc eclogite reported by Angiboust *et al.*
869 (2014) are partially overlapping with some HT rutile in the 0.3-0.5 Nb/V range (Fig. 17b).
870 Yet, the overall reproducibility of this diagram seems to sustain.

871 While only mafic-derived rutile were used here to test trace-element pressure
872 dependency, there are however subtle compositional variations within gabbros and basalts
873 from the ocean floor that can lead to variable rutile compositions. Due to the lack of available
874 bulk trace element compositions for the set of rocks used in this study, we have, alternatively,
875 tested the influence of composition by forward modelling trace element re-distribution in
876 rutile using RStudio®. We have resourced to a compilation of trace element abundances from
877 MORB and ophiolite gabbros (including some in our collection), and selected the elements of
878 interest (V, Sc, Nb). We used parageneses and a range of mineral proportions like those found
879 in our samples for both eclogitic (HP) and granulitic (HT) conditions, and published
880 partitioning coefficients for all the mineral pairs considered. Details on the modelling
881 parameters, inputs and outputs (ternary diagrams) are presented in the Supplementary Data:
882 Electronic Appendix 5 and the script is available as Appendix 6. Even though generous
883 uncertainties are incorporated into the forward model, as a reflection of large uncertainties in
884 the mineral partitioning coefficients and phase proportions, rutile compositions of eclogite
885 and granulite are always clearly separated, showing that this dichotomy is not primarily a
886 function of bulk-rock geochemistry. The eclogite to granulite split occurs along the line
887 parallel to the Sc-Nb join, as seen with our natural rutile data (Fig. 11c), indicating significant
888 differences in the fractionation of V. Given that rutile-pyroxene partition coefficients differ
889 only within a factor of two or so (Holycross and Cottrell, 2022), it appears that V distribution
890 is controlled by major mineral assemblages which are clearly different between eclogite and
891 granulite. It suggests that variations towards Nb are mainly controlled by the fraction of rutile
892 in the rock, and the smaller the rutile fraction is, Nb in rutile increases.

893 *What can explain V and Sc pressure dependency in rutile?* Metamorphic rutile formed
894 from mafic protoliths is commonly enriched in V (> 1000 ppm; Miller *et al.*, 2007; Gao *et al.*,
895 2010; Hart *et al.*, 2016). In our dataset, HP rutile plot systematically closer to the Cr+V

896 endmember when compared to the HT rutile. Vanadium is generally compatible with rutile,
897 although at times it behaves as incompatible ($D_V^{\text{Rutile/melt}}$ 0.1 – 18, Klemme et al., 2005;
898 $D_V^{\text{Rutile/melt}}$ 47 – 124, Foley et al., 2000). It has been demonstrated that V partitioning to rutile
899 increases with increasing fO_2 (Holycross and Cottrell, 2020). This implies that V will be
900 easily incorporated by rutile under most oxidising geological conditions. At higher
901 temperature conditions, V might be preferentially incorporated into other phases (e.g.
902 ilmenite, clinopyroxene, amphibole; Nehring *et al.*, 2010), and, as a consequence, rutile will
903 contain less V and proportionally more Nb, for which it is the main carrier (Zack *et al.*, 2002).
904 This seems to be the case of HT rutile grains that are plotted in Figure 17. The modelling also
905 highlighted that V incorporation by rutile can also be controlled by bulk composition. Thus, V
906 in rutile is controlled both by the element concentration, fO_2 and the paragenesis. As a result,
907 if the compositional system is homogeneous (MORB), the critical factors are oxygen fugacity
908 and minerals in competition with rutile.

909 Although rutile is commonly poor in Sc, Sc can behave similarly to Sb. Sc is often
910 retained by amphiboles (e.g. glaucophane, pargasite), garnet or clinopyroxene in metabasites
911 (Hermann, 2002; Spandler *et al.*, 2003). Some Sc can be partitioned into rutile due to
912 amphibole or primary Ti-rich clinopyroxene breakdown. This could account for some variable
913 Sc found in rutile from the IODP and as inclusions in the Queyras titanite grains. As the
914 modelling highlighted, the interplay of partitioning coefficients between the different
915 clinopyroxenes (omphacite and augite) and hornblende exerts a strong control on Sc contents
916 in rutile.

917 **4.4. Applications and future directions using Ti-bearing minerals for LT-HP hunting**

918 As we probe the exposed and preserved metamorphic rock record through time, retrieving HP
919 metamorphic conditions before the Neoproterozoic has been challenging (e.g., Brown and
920 Johnson, 2018; Palin *et al.*, 2020). There seems to exist a preferential preservation of HT

921 orogenic cores in ancient mountain belts (e.g. *Hercynian Massif Central*). The scarcity or
922 absence of HP metamorphic rocks in ancient metamorphic belts can potentially be explained
923 by a low preservation potential and, therefore, detrital grains formed after weathering of those
924 units may be their only survivors. Rutile and titanite are easily found as detrital grains in
925 stream sediments and in sedimentary rocks (e.g. *Pereira et al., 2020; Zhou et al., 2020*) but P-
926 dependent proxies are required to discriminate HP from LP metamorphic rutile and titanite. In
927 this contribution we have presented data of mafic rocks that underwent metamorphism both in
928 the ocean floor, at low pressure, and during subduction at higher pressures. Our data suggest
929 that Nb/V and Sc-Nb-V+Cr in rutile and V-Yb-La/Sm in titanite are pressure dependent (Figs.
930 16-17). To apply these petro-geochemical tools in sediments, we propose a multi-step
931 discrimination approach, explained in the following.

932 For titanite, a first approach is to examine BSE images to select igneous from
933 metamorphic titanite based on BSE zoning (*Bruand et al., 2014* for magmatic; *Garber et al.,*
934 *2017* for metamorphic). Grains showing no zoning or metamorphic zoning can be selected
935 and their trace element data plotted on the ternary diagrams proposed here (Fig. 16-b,d). This
936 discrimination, however, leads to some HP titanite potentially being discriminated as LP.

937 For rutile, data should be plotted in diagrams that discriminate metamorphic derived
938 from other rutile types (*Agangi et al., 2019; Pereira et al., 2019*). A sub-selection of
939 metamorphic-only rutile can then be plotted using the Cr-Nb discriminating diagram
940 (*Meinhold et al., 2008* and others) to pool metamafic-derived rutile grains. This subset can
941 then be projected in the Sc-Nb-V+Cr and Nb/V vs $\sum\text{Zr/Hf}$, Nb/Ta diagrams (Fig. 17). Further
942 work is required to assess suitability for rutile grains derived from metapelitic protoliths.

943 5. Conclusions

944 In this study we have compared metamafic rocks with MORB affinities of modern and pre-
945 alpine oceanic crust (IODP and Chenaillet massif) to metagabbros and metabasalts that have
946 undergone HP during subduction to the blueschist and eclogite facies conditions (Queyras,
947 Monviso, Corsica).

948 Despite rutile being mostly recognised as a stable phase at eclogite and upper
949 amphibolite/granulite conditions due to Barrovian metamorphism, we found that amphibole-
950 bearing gabbros in the oceanic crust can stabilise rutile at HT and very LP (ca. 2 kbar) (Fig.
951 18). Pressure and temperature are among the most critical parameters affecting the stability of
952 these two minerals. Yet, we show that for rocks that experienced similar P-T conditions,
953 subtle variations in the Ti/Ca and Ca/Al values seem to affect the titanite to rutile formation,
954 even in rocks with similar bulk chemistry and grain sizes. These ratios, which may shift
955 locally during mineral replacement reactions, can be the main controlling factor when both
956 rutile and titanite are found stable at either LP or HP, especially in only partially equilibrated
957 metagabbros.

958 Trace-element based discrimination tools are complex, and the chemistry of mineral
959 phases are dependent on several factors including the bulk rock composition, local equilibria,
960 partitioning coefficients between minerals and rocks and co-existing mineral phases during
961 Ti-phase growth. Our results highlight that titanite trace element compositions in different
962 MORB affinity metamafic rocks show strong variability. Titanite from our study shows that
963 blueschist and eclogite yield a marked La depletion, and thus low La/Sm_N values in
964 comparison to amphibolite grade titanite. La/Sm_N or $Nb+Ta$ together with Yb and V can be
965 used to discriminate HP titanite. Rutile grains from the metagabbros studied here display
966 distinctive chemistries that are pressure indicative; rutile from HT-LP yield V/Nb values
967 above 0.5 while V/Nb in rutile from eclogite facies are below 0.3. In the same way, V+Cr, Sc,

968 and Nb compositions can be used, following the workflow described in the discussion, to
969 distinguish HP from LP metamafic rutile.

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978 Data Availability Statement

979 The data underlying this article are available in the article and in its online supplementary
980 material.

981 Supplementary Materials are available in the electronic version online. Appendix 1 comprises
982 a full method description, sampling details, amphibole mineral chemistry detailed description
983 and phase diagram modelling details. Appendix 2 includes silicate mineral chemistry and
984 amphibole thermometry. Appendix 3 includes the full titanite, rutile and ilmenite chemistry
985 dataset.

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Grade	LP-HT (IODP)					LP-HT (W. ALPS)			HP-LT			HP-MT					
	1316	1318	1320	1319	5450	CH1	CH3	CHDI	QE01	QE06	TR10	C16	Cv16	C147	PG	LS	VE29
Locality	LEG 209		305	118	Chenaillet Massif			Queyras			Corsica			Passo Gallarino	Lago Superiore	Verne Monviso	
T (°C)	740 - 840			-	700-900			330 - 400		380 - 470	500 - 550			450; 525	550 - 600	500-550	
P (kbar)	< 2				< 2			11 - 15		12 - 18	21 - 24		12; 25-29		19; 25 - 27	21-27	
Ref.	This study				This study; G			E		A	F			C; B		C; B	D

A. *Debret et al., 2016*; B – *Angiboust et al., 2012*; C. *Schwartz et al., 2000*; D. *Groppo and Castelli, 2010*; E. *Agard et al., 2001*; F. *Vitale-Brovarone et al. 2011*; G. *Mevél et al., 1978*.

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1292 Table 2. Paragenetic chart of the studied samples, organised by metamorphic grade, from LP-H/MT to HP-LT
 1293 and HP-MT. The peak assemblage represents the highest metamorphic grade each rock attained.

Samples/ Minerals	LP – H/MT								HP – L/MT								
	13 16	13 18	13 19	13 20	54 50	CH1	CH3	CH DI	QE-01	QE-06	TR 10	C 16	C v 16	C 147	PG	LS	VE 29
Augite	m	m	m		m	m				m							
Plagioclase	m	m	m	m	m												
Hornblende ^a	x	x		x	x	x	x										
Heden- Diopside			x														
Garnet													x	x		x	x
Omphacite													x	x	x	x	x
Hornblende ^b	x	x	x	x	x	x	x		x	x							
Actinolite						ret	ret	x	x								
Glaucofanec ^c									x	x	x	x	x	x			
Epidote					x				x	x	x	x		x	x	x	x
Lawsonite									x	x	x		x	x			
Albite		ret	ret	ret		ret	ret	x	ret	ret	ret	ret					
Quartz	x									x	x			x	x	x	x
white mica										x	x				x	x	x
Ilmenite		m/x	m/x	m/x	m/x	x	x	x			x				x	x	x
Rutile	x i		x i	x i	x i	x i	x i	x i	i	i	i				x	x	x
Titanite	x	x	x ov	x ov	x	x	x	x	x	x	x	x	x	x	ret		
Chlorite	ret				ret	ret					x		ret				
Carbonate	ret		ret		ret					ret	ret						

Note that some minerals indicated on the list are one possible end-member. a. another equivalent brown amphibole; b. another equivalent green amphibole; c. another equivalent blue Na, Na-Ca amphibole. In the table, m – magmatic, x – metamorphic and i – inclusion, ov – overgrowth, ret – late retrograde phase

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	Σ REE	Σ LREE	La/SmN	Eu/Eu*	Th/U
Amphibolite	10 - 3358	2.6 - 2400	0.02 - 1.2	0.2 - 29	0.01 - 0.48
Granulite*	13 - 2000	11 - 1678	5 - 93	4.7 - 314	0.08 - 1.06
Blueschist	32 - 370	3.8 - 59	0.02 - 0.7	0.9 - 3.2	0.04 - 0.2
Eclogite	168 - 465	20 - 208	0.02 - 0.3	0.84 - 1.8	0.004 - 2.3

REE sums are in ppm. Data was normalized to Sun and McDonough (1995) Chondrite values. Amphibolite corresponds to a combination of Ch3 and IODP1319, granulite to IODP1320, blueschist to a combination of QE1, QE6, TR10, and eclogite to C147, C16v and C16 data. *titanite in the granulite sample is most probably formed at lower temperatures, but the sample rock forming minerals did not retrogress to those conditions.

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1319 Figure captions

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1321 *Figure 1 Experimental vs pseudosection modelling diagrams for a MORB composition and focusing on Ti-*
1322 *phases. A. P-T diagram highlighting the fields of known stability of rutile, titanite and ilmenite for MORB*
1323 *compositions (Liou et al., 1998). In white a field of no experimental data available for mafic compositions. Grey*
1324 *lines delimit different metamorphic facies. Two dotted lines separate the fields of low (< 440 °C/GPa) from*
1325 *moderate and higher geothermal gradient (> 770 °C/GPa) (Palin et al., 2020). Boxes correspond to*
1326 *approximate peak P-T conditions from samples collected for this study, highlighting cooling of amphibole-*
1327 *bearing gabbros from the oceanic crust, and burial history of blueschist and eclogite rocks from the Alpine*
1328 *chain. B. MORB phase diagram modelling from Green et al. (2016) showing titanite-out reaction (in green),*
1329 *solid blue line the same as in a) and a dotted white line at 600 °C, below which there are no data constraints on*
1330 *Ti-phase stability for mafic compositions (taken from inset a).*

1331 *Figure 2. Sampling locations in the Western Alps and Corsica (A, B) and from IODP drilling sites (C). Western*
1332 *Alps maps were modified after Angiboust et al., (2012), Debret et al. (2016), Locatelli et al., (2019), and Vitale-*
1333 *Brovarone et al., (2011), and IODP drilling site information and map were modified from data and maps*
1334 *available at iodp.org.*

1335 *Figure 3. Microphotographs of samples from IODP using PPL (A, C, G, H), reflected light (D- E) and BSE*
1336 *imaging (B, F). A-B, occurrence of sub-spherical rutile grains rimmed by titanite enclosed by altered*
1337 *plagioclase and brown to green amphibole; C, occurrence of large ilmenite grains rimmed by titanite; E-F*
1338 *ilmenite-rutile-titanite grains, with titanite replacing former ilmenite usually found at the cores; F is a zoom in*
1339 *of E, illustrating the occurrence of rutile rimmed by ilmenite and only then by titanite; G-H, occurrence of*
1340 *idioblastic titanite associated with altered plagioclase and brownish-green amphibole.*

1341 *Figure 4. Microphotographs of samples from the Chenaillet Massif using PPL (A-D, G), reflected light (E, F)*
1342 *and BSE imaging (H). A and B, general textural and mineralogical relationships; C, of the late-stage titanite*
1343 *and actinolite shear zone in ChI; D-F. detail of ilmenite-rutile-titanite occurrences in ChDi; G-H ilmenite-*
1344 *titanite growth in the shear zone, and strong BSE zoning in these titanite grains.*

1345 *Figure 5. Microphotographs of blueschist samples from the W. Alps in the Queyras using PPL (A, B, C),*
1346 *reflected light (D) and BSE imaging (E, F). A- plagioclase domains with lawsonite, albite and epidote, and relic*
1347 *magmatic clinopyroxene replaced by glaucophane and epidote and titanite. B- overgrowth of actinolite on*

1348 glaucophane in plagioclase sites, with titanite. C-F, titanite grains with very fine-grained lamellae rutile
1349 inclusions in all the different investigated samples.

1350 Figure 6. Microphotographs of eclogite samples from the W. Alps in the Passo Gallarino-Monviso-Lago
1351 Superiore (A-D) and Corsica (E-H) using PPL (B, C, E-G), reflected light (A) and BSE imaging (D). A – in
1352 Passo Gallarino a thin rim of titanite is found around rutile; B- omphacite, garnet, and phengite associated to
1353 rutile; C – Foliated eclogite with omphacite, garnet, and rutile ribbons; D – a detail of a rutile ribbon from C
1354 with ilmenite/Fe-Ti oxide exsolution; E- blueschist eclogite, with albite, glaucophane crossed-cut by a vein
1355 showing omphacite, garnet, lawsonite and titanite; and G-H- eclogitic metabasalt with garnet, omphacite,
1356 lawsonite, glaucophane and titanite.

1357 Figure 7. P-T diagram illustrating the conditions of either rutile or titanite stability found in the samples studied
1358 here. P-T constraints are from references found in table 1.

1359 Figure 8. X_{Al} vs F_{apfu} in titanite binary diagram. Different symbols and colours represent average titanite values
1360 from different samples. Green symbols are data from this study. Diamonds are for titanite from IODP samples,
1361 squares for the obducted Chenaillet metagabbro, circles for blueschist and triangles for eclogite titanite grains
1362 from this study. All other coloured symbols correspond to titanite data from other metamorphic grades and bulk-
1363 rock compositions that are plotted for comparison. Literature data from Franz and Spear, 1985, Markl and
1364 Piazzolo, 1999, Ye et al., 2002, Castelli and Rubatto, 2002, Garber et al., 2017.

1365 Figure 9. Trace element diagrams for titanite, A) Nb/Ta vs Zr/Hf. Star refers to Chondrite compositions (Sun
1366 and McDonough, 1989); B) Sn vs V binary diagram. Diamonds correspond to IODP, squares to the Chenaillet
1367 Massif amphibolite metagabbro, circles to blueschist and triangles to eclogite rock titanite.

1368 Figure 10. REE and Th/U values of titanite from HT and HP conditions.

1369 Figure 71. Discrimination diagrams for rutile. A) The Cr-Nb discrimination modified from Meinhold et al.
1370 (2008). B) The Zr/Hf vs Nb/Ta modified from Pereira et al. (2019); C) based on PCA, the three main vectors V,
1371 Cr, Sc and Nb used in a ternary diagram; D) Binary diagram of the sum of HFSE ratios vs Nb/V. HT IODP are
1372 presented with two colours, the brighter red corresponding to sample 1320 and darker red to 1319 and 1316.
1373 Some labelled datapoints correspond to LA-ICP-MS acquired data used in Figure 11.

1374 Figure 82. Rutile occurrence in amphibole-bearing gabbros on the oceanic crust (example from IODP1320); a)
1375 Reflected light microphotograph in grayscale of an ilmenite grain being replaced by titanite and rutile. Rutile

1376 occurs as exsolution lamellae and coarse grains; b) elemental concentrations of Zr and Cr in rutile from
1377 different textural domains labelled in part a). The symbol * correspond to a laser spot measurement.

1378 Figure 93. Binary trace element diagrams for ilmenite grains from samples IODP 1320 and 5450.

1379 Figure 10. Composition in Ti/Ca and Ca/Al of several metagabbros, blueschist and eclogite rocks from different
1380 settings (IODP, Syros, Western Alps, Corsica) grouped by titanite, rutile or titanite-rutile stable (symbols).
1381 Range and average of unaltered MORB composition is given by the red circle and the red star respectively
1382 (Klein, 2004). Altered vectors correspond to altered MORB compositions from Zhang and Smith-Duque (2014).
1383 Corsica data are from Vitale-Brovarone et al. (2011) and Syros samples from Marschall (2005). Data used here
1384 can be found in the Supplementary Data: Electronic Appendix 1.

1385 Figure 11. Trace element diagrams for titanite modified from Scibiorski et al. (2019). A. Σ HREE vs Y; B. Y vs
1386 Dy/Yb; In grey, labels adapted from Scibiorski et al. (2019) and in pink based on data from this study. Pink-
1387 coloured circles correspond to titanite grains in a garnet-bearing assemblage.

1388 Figure 12. Ternary diagrams for titanite compositions from multiple geodynamic environments. Compiled
1389 literature data are from a. Bruand et al., 2014, b. Garber et al., 2017, c. Holder and Hacker, 2019, d. Marsh and
1390 Smye, 2017, e. selection of Scibiorski et al., 2019, f. Spandler et al., 2003, g. Ng et al. 2016, h. El Korh et al.,
1391 2009.

1392 Figure 13. Rutile discrimination diagrams for metamafic, gabbro in composition, rocks .a) is a V+Cr-Sc-Nb
1393 ternary and b) a V/Nb vs sum of HFSE ratios. Grey bar is highlighting some overlap between HP and LP
1394 signatures and arrow indicate the field towards HP rutile. a. relic in bs correspond to HT-LP rutile grains
1395 found as inclusions in blueschist titanite. Published data are from metamafic rocks b. and c. from blueschist and
1396 eclogite from Syros (Hart et al., 2016), talc and lawsonite eclogite from Monviso (Angiboust et al., 2014) and d.
1397 LP granulite (Cruciani et al., 2020).

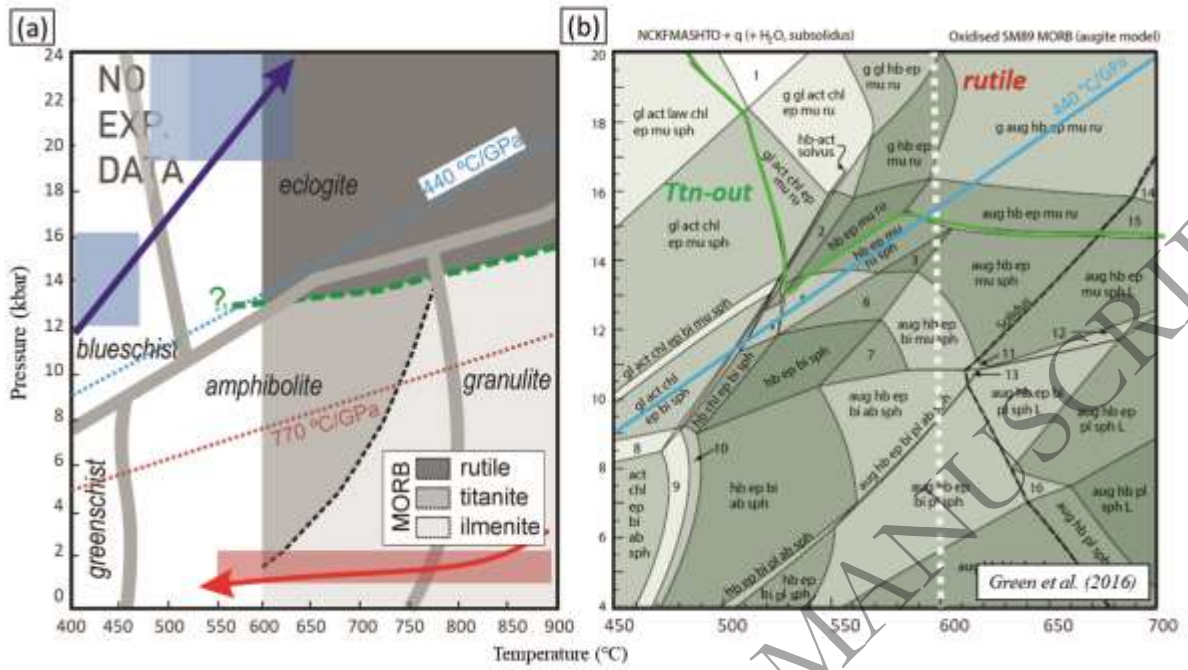
1398 Figure 14. Ti-mineral phases in metagabbros and metabasalts associated to divergent and convergent settings
1399 and some of their main trace element systematics.

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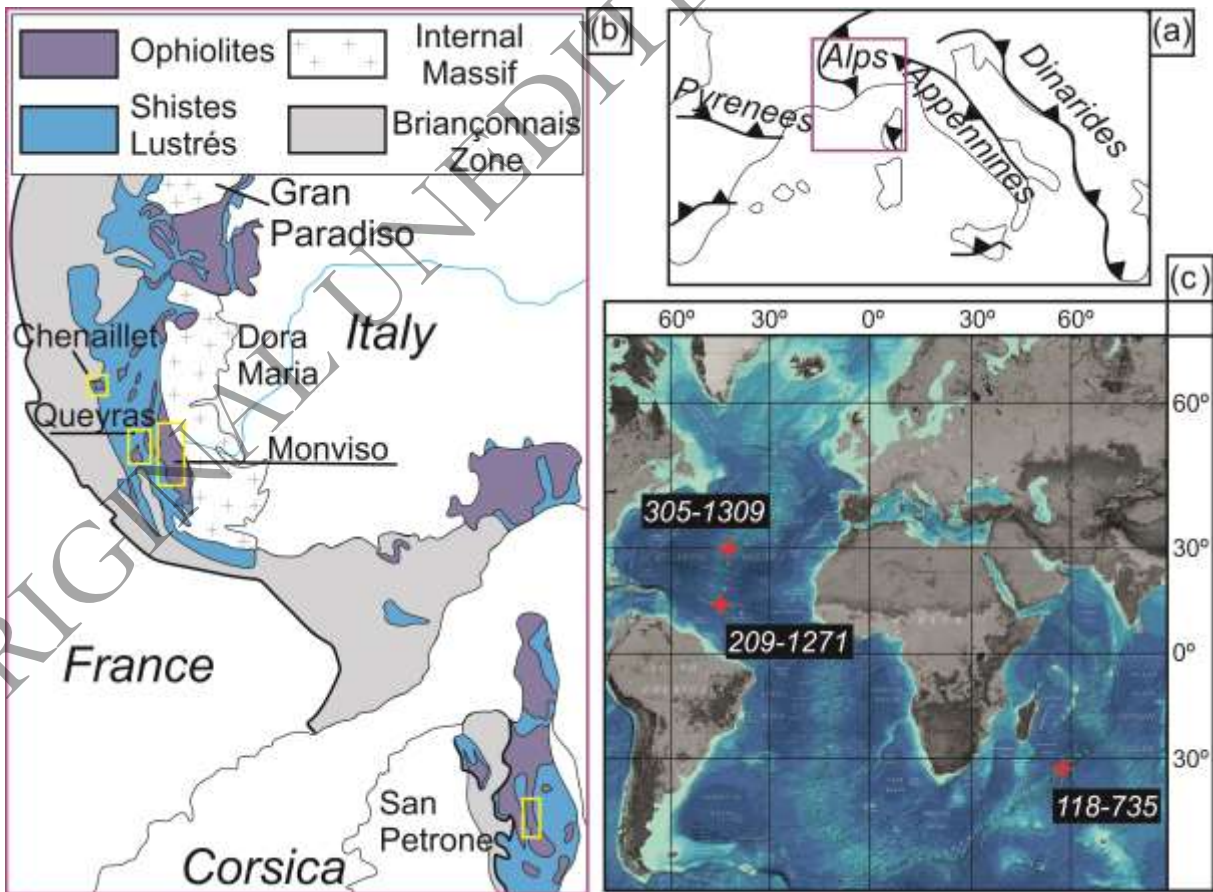
1403 Fig. 1.



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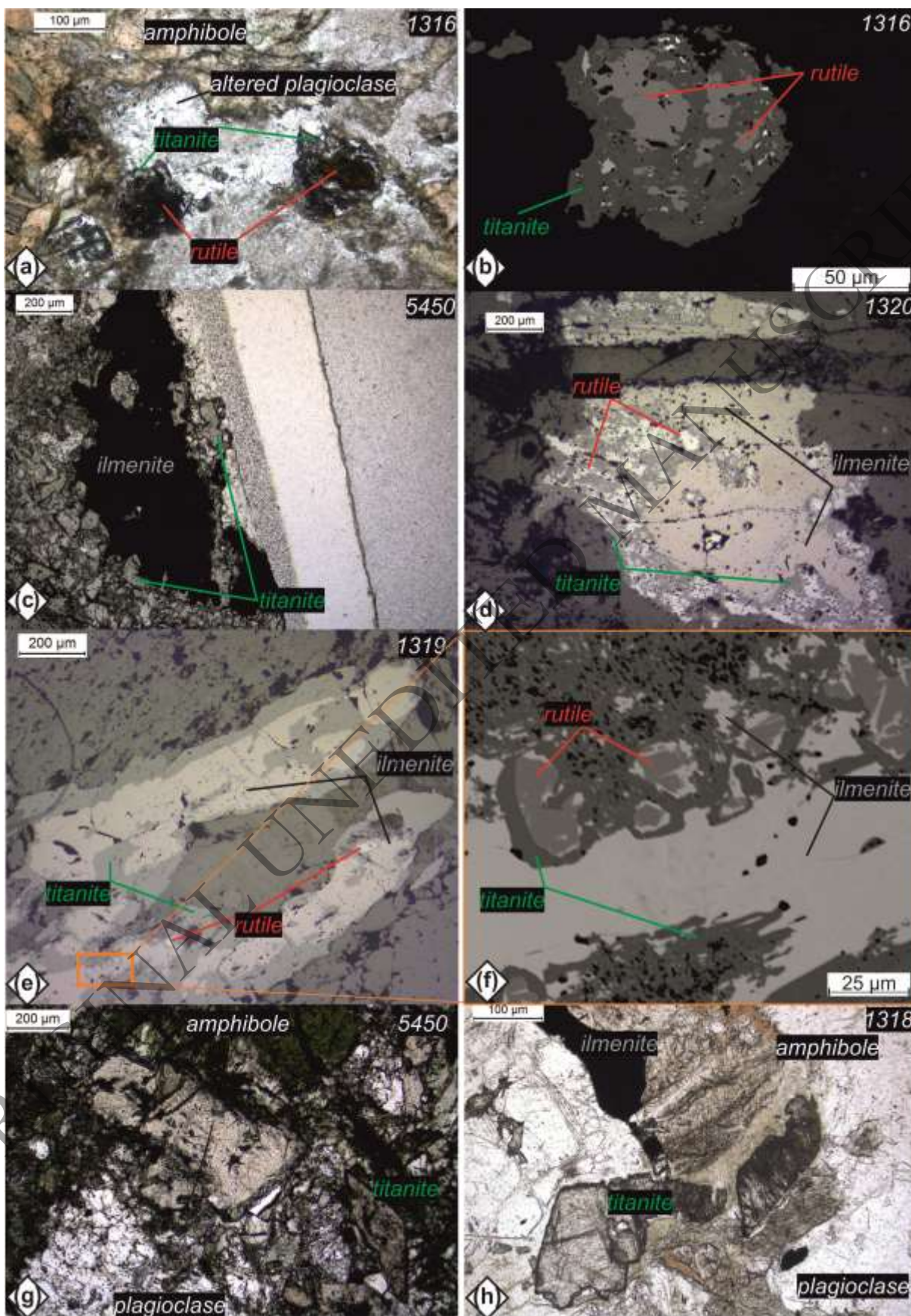
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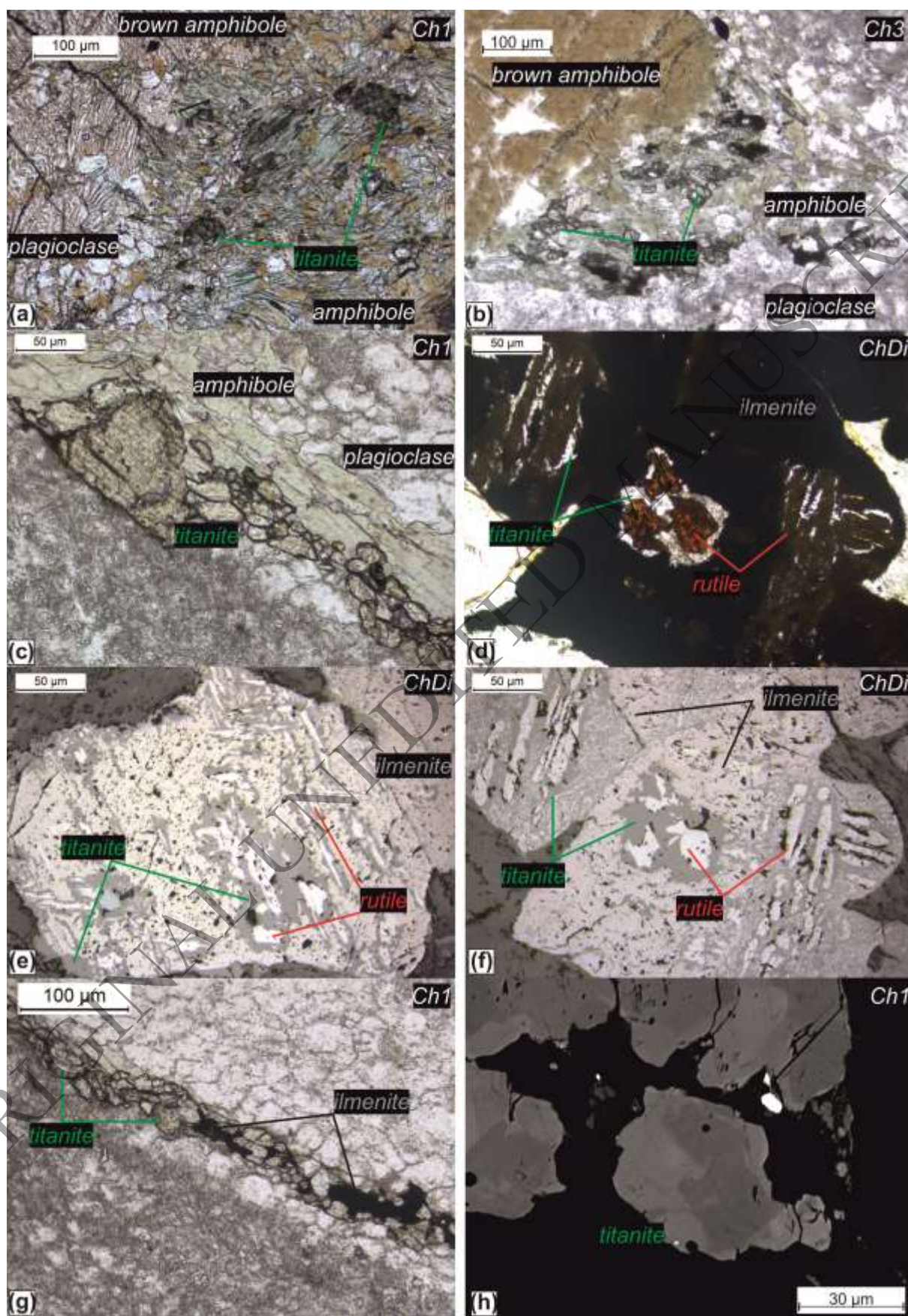
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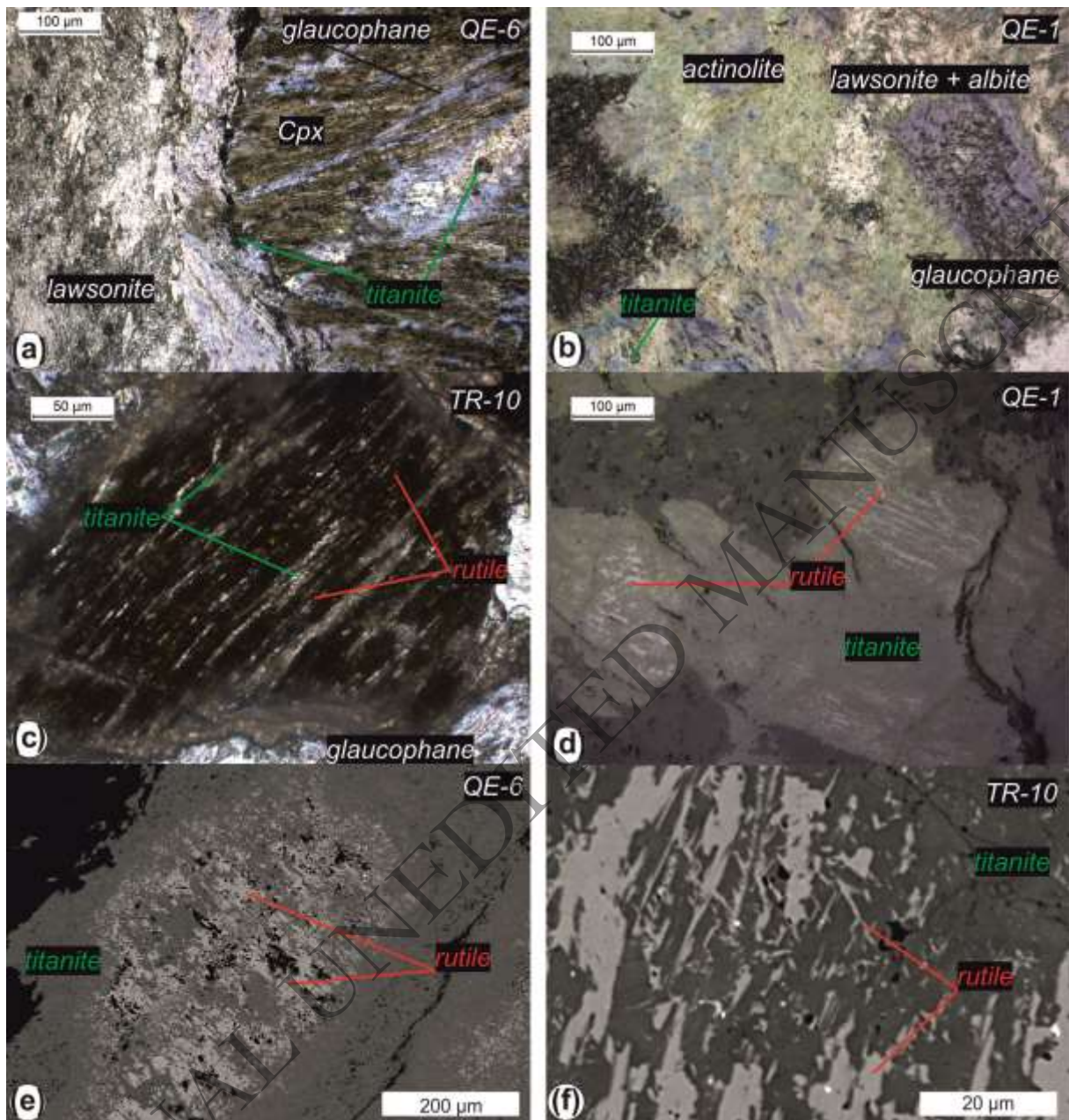


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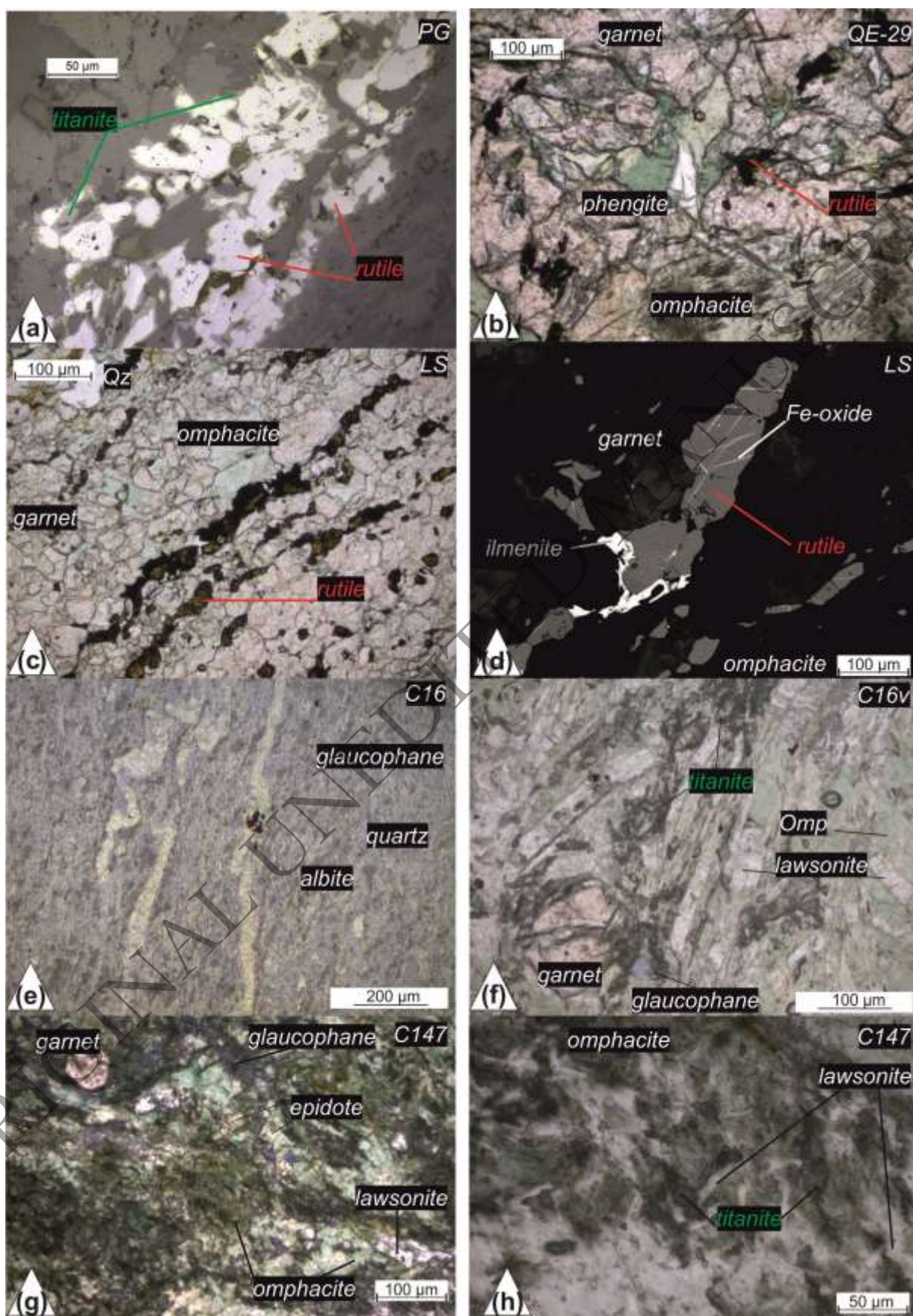
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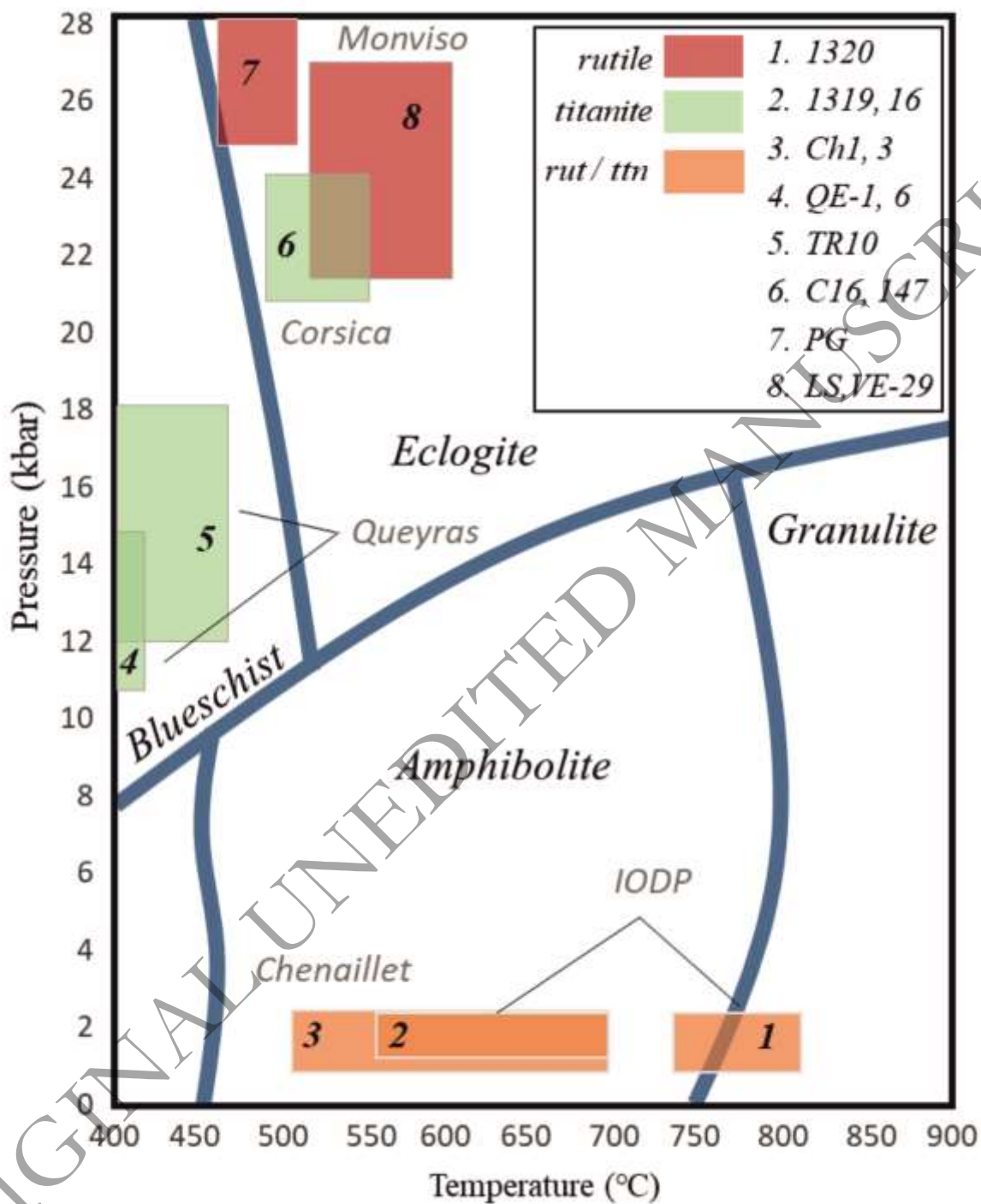






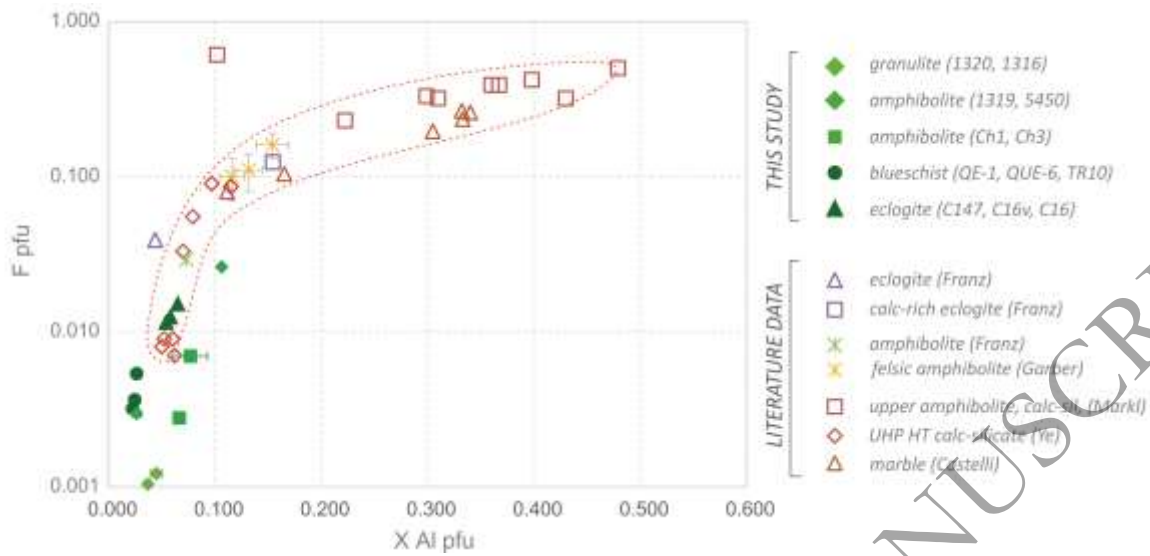
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1434 **Fig. 8.**

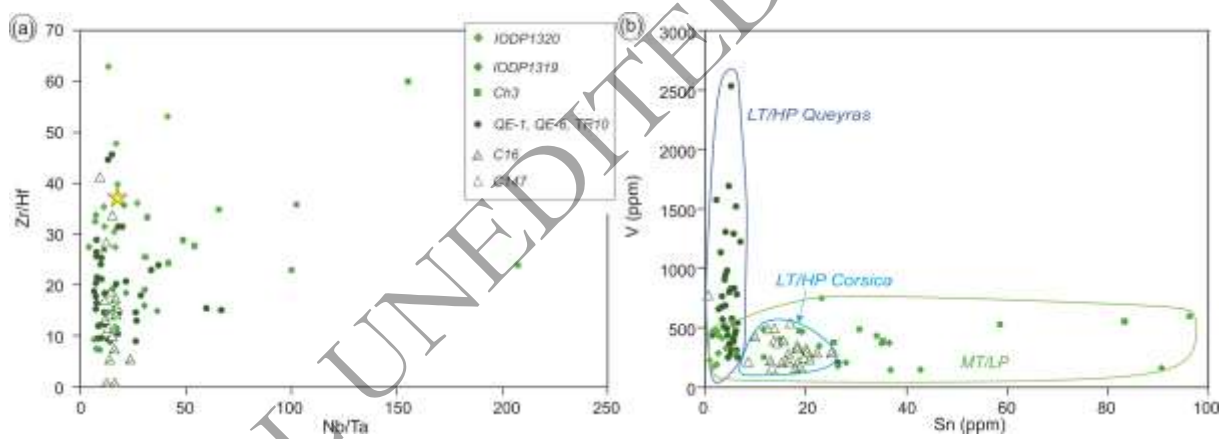


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1438 **Fig. 9.**



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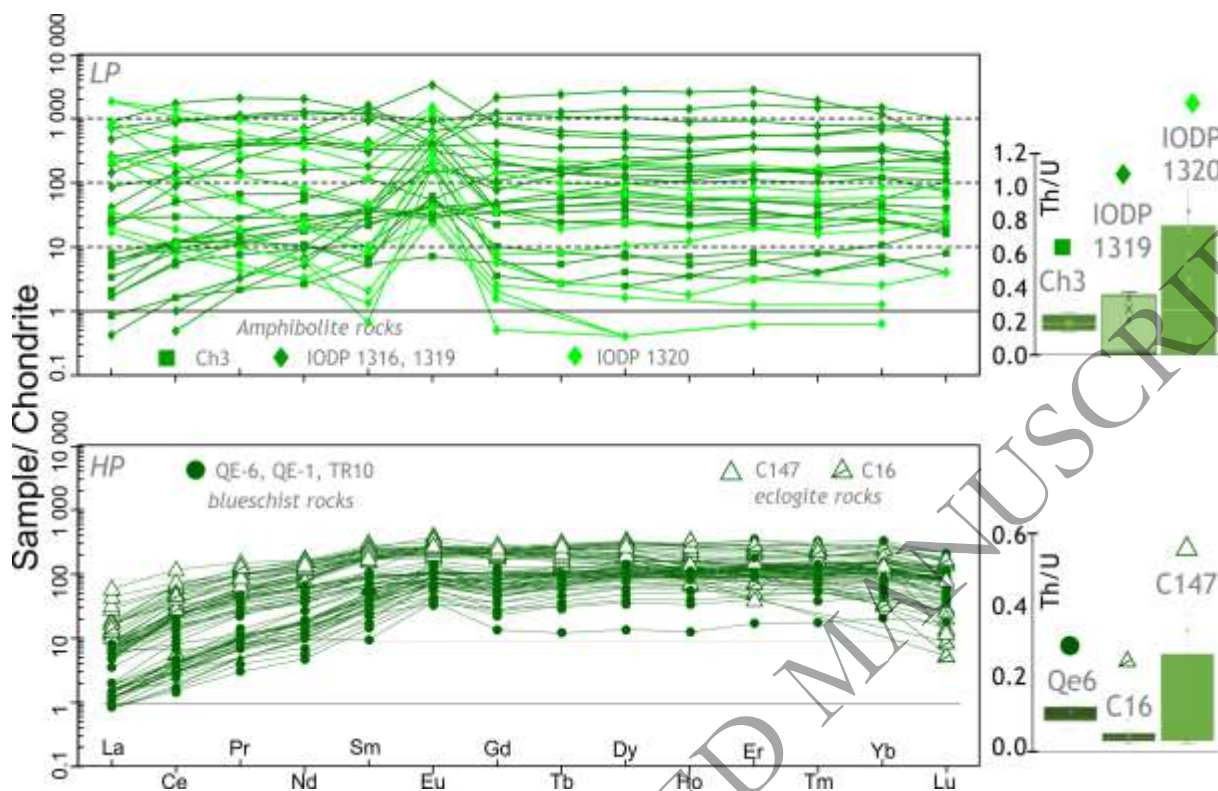
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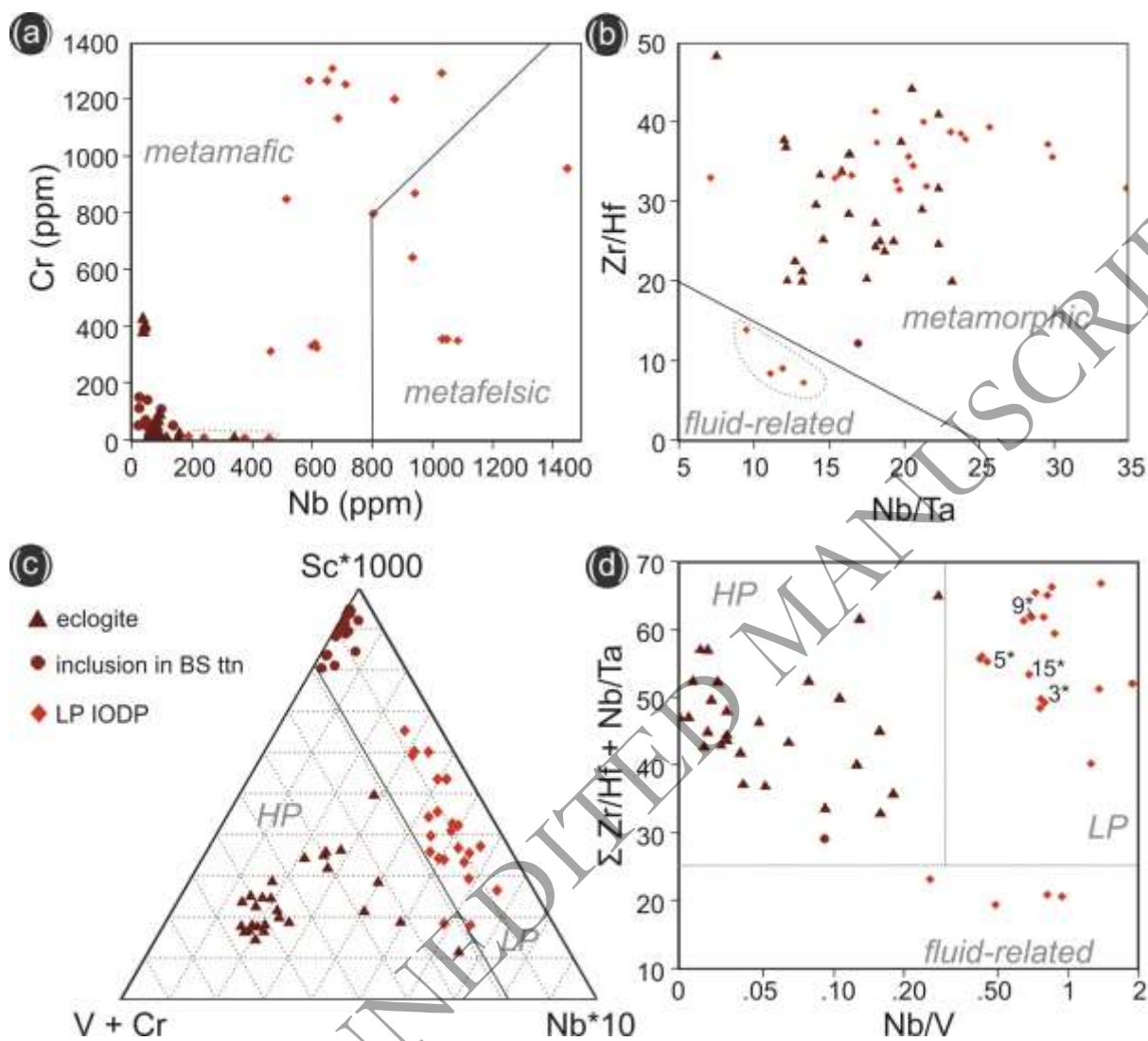
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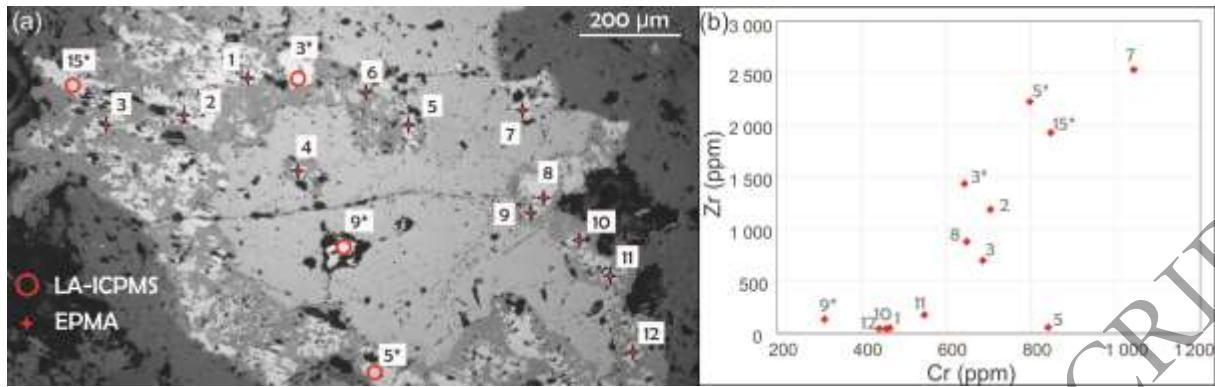


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1480 **Fig. 12.**



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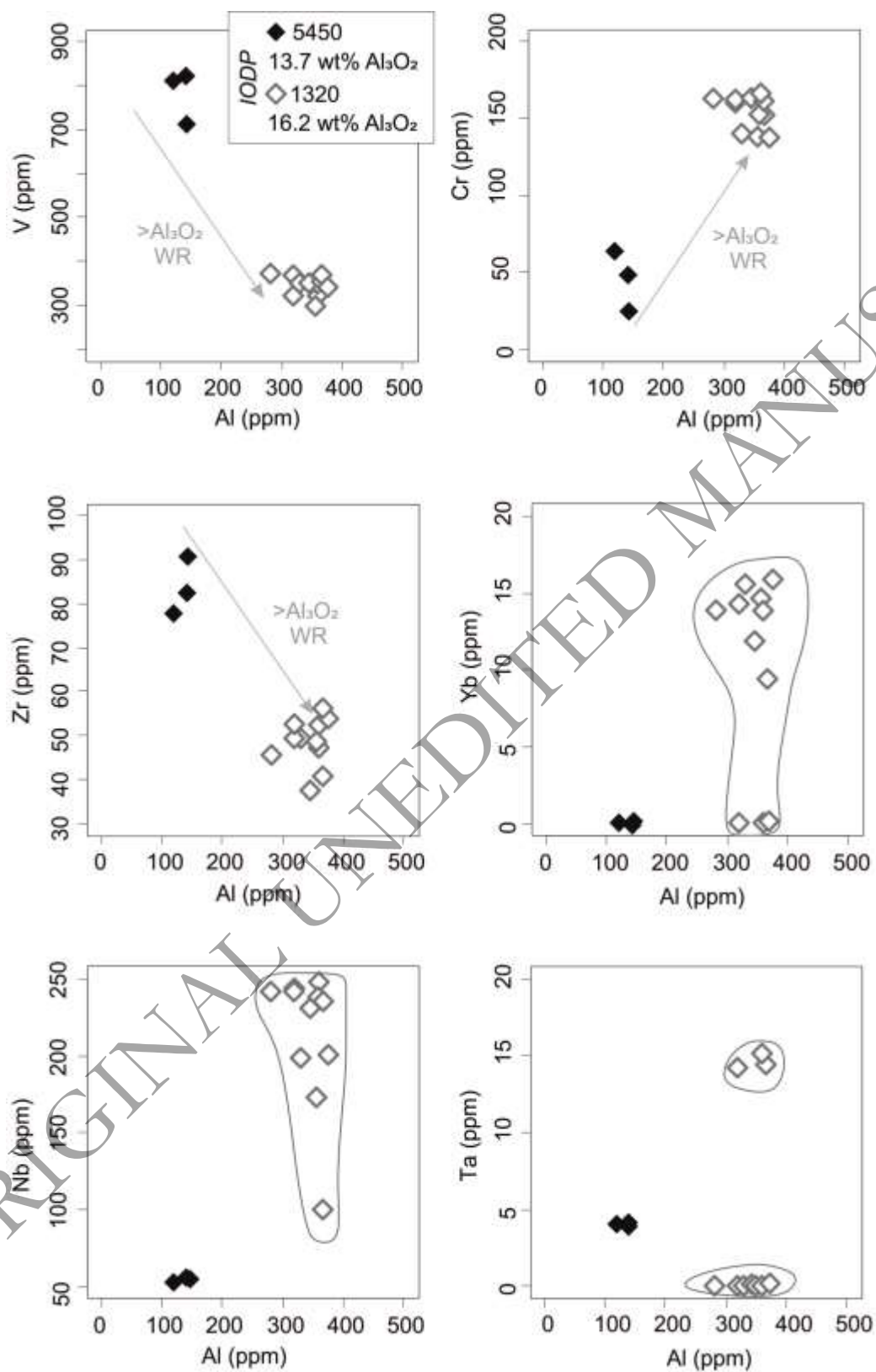
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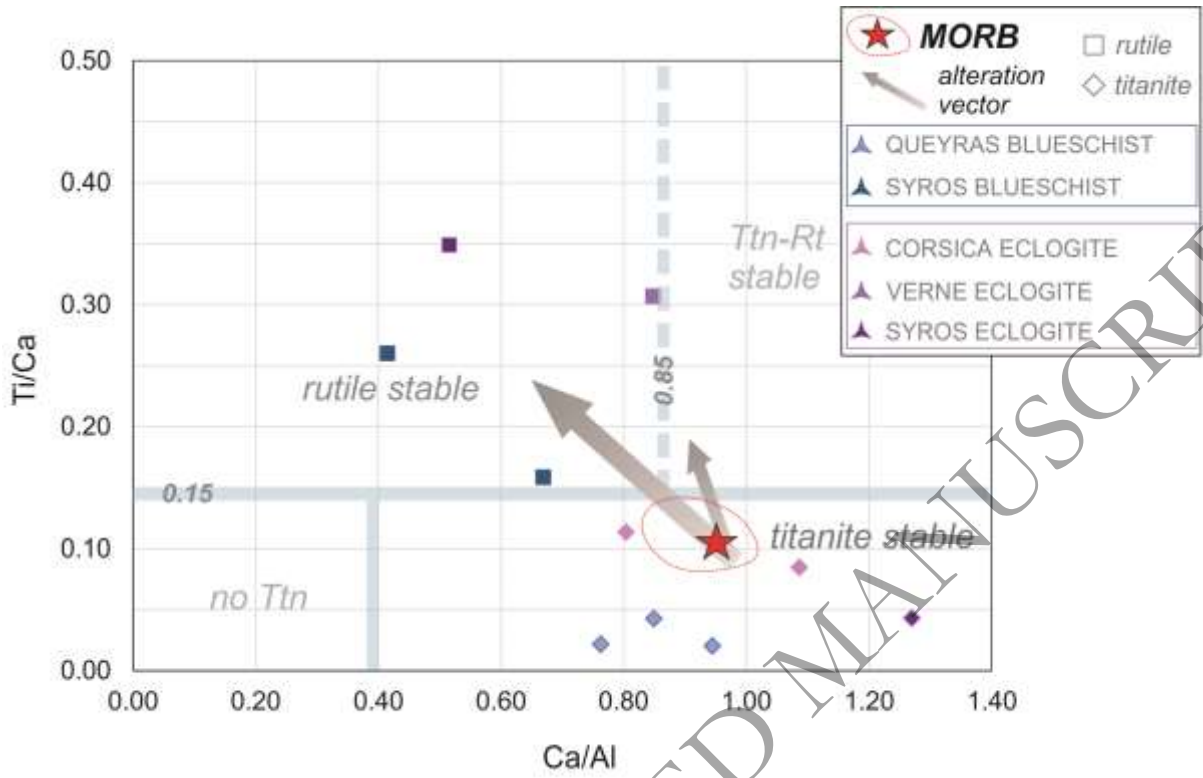
1504 Fig. 13.

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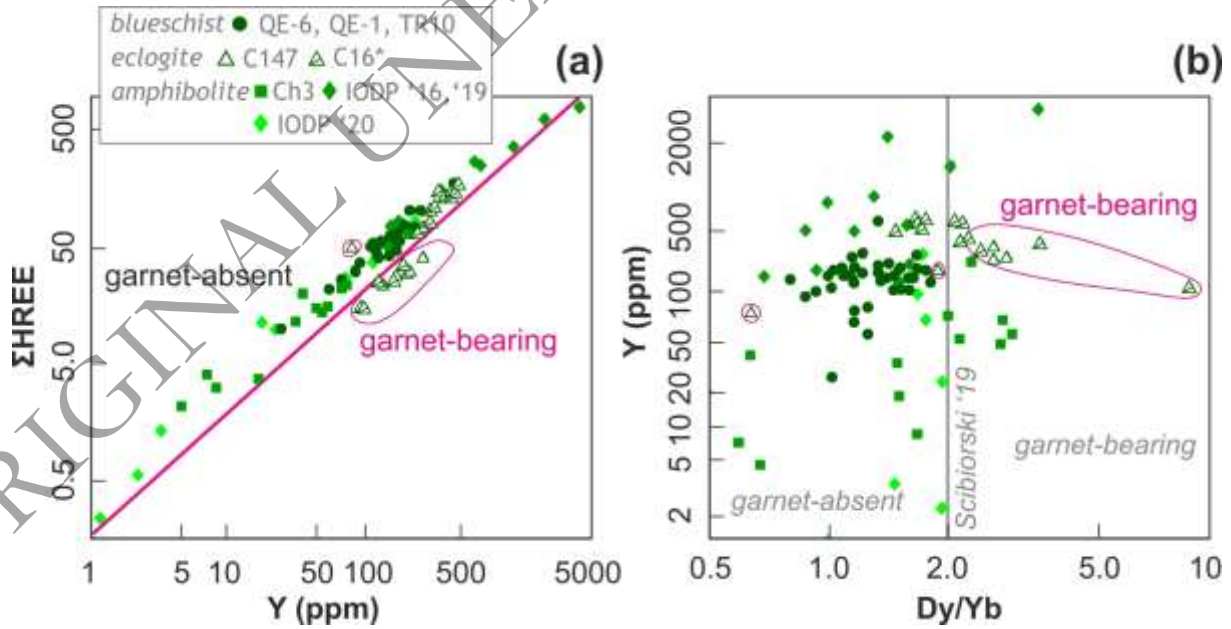
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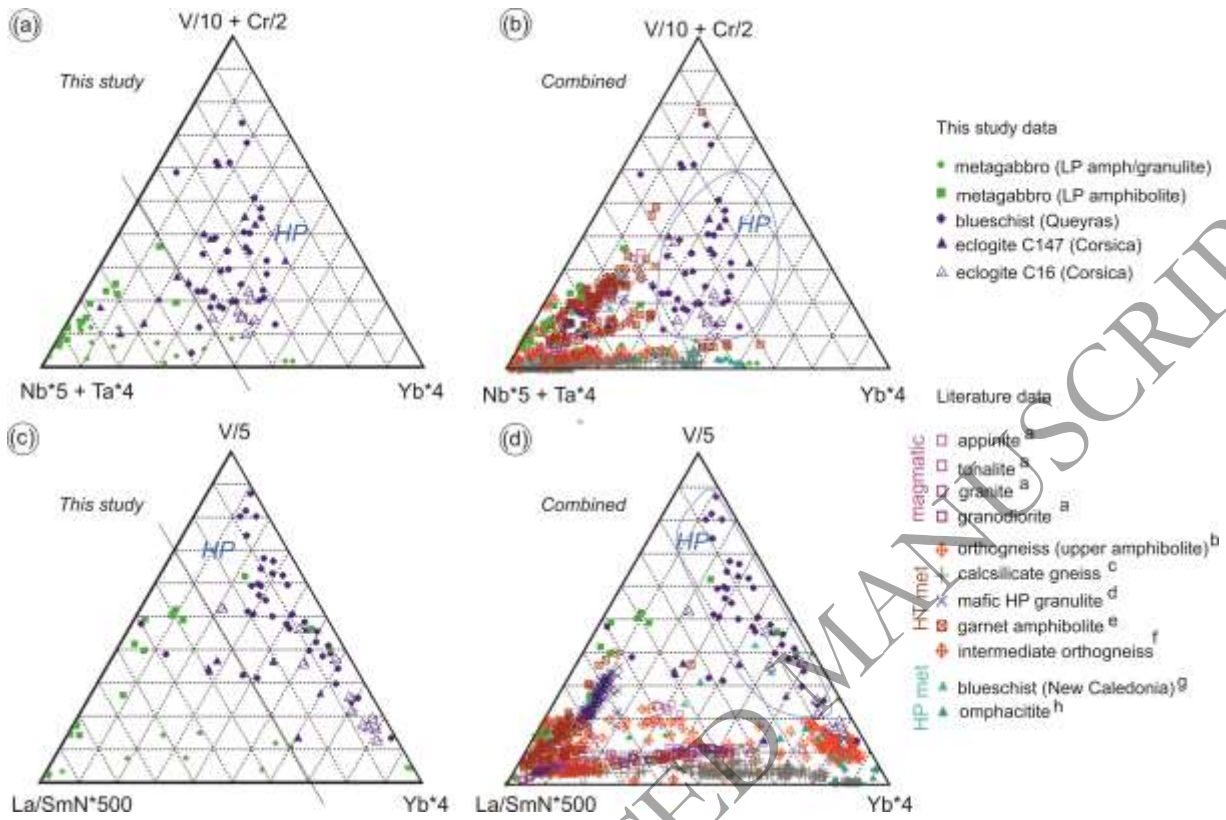
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1517 **Fig. 16.**

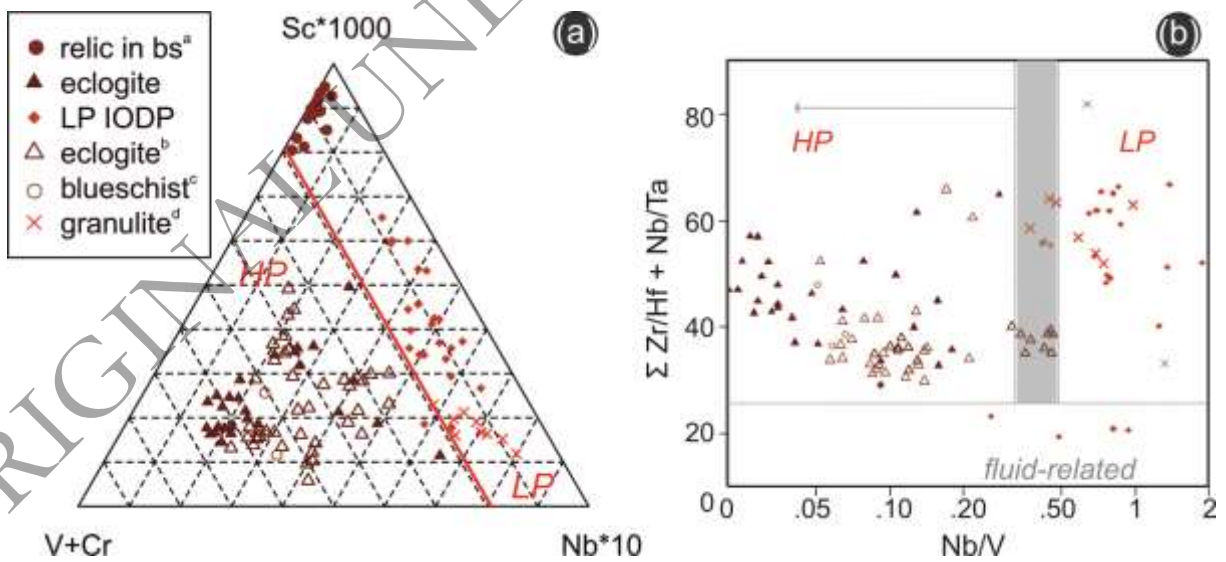


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1521 **Fig. 17.**



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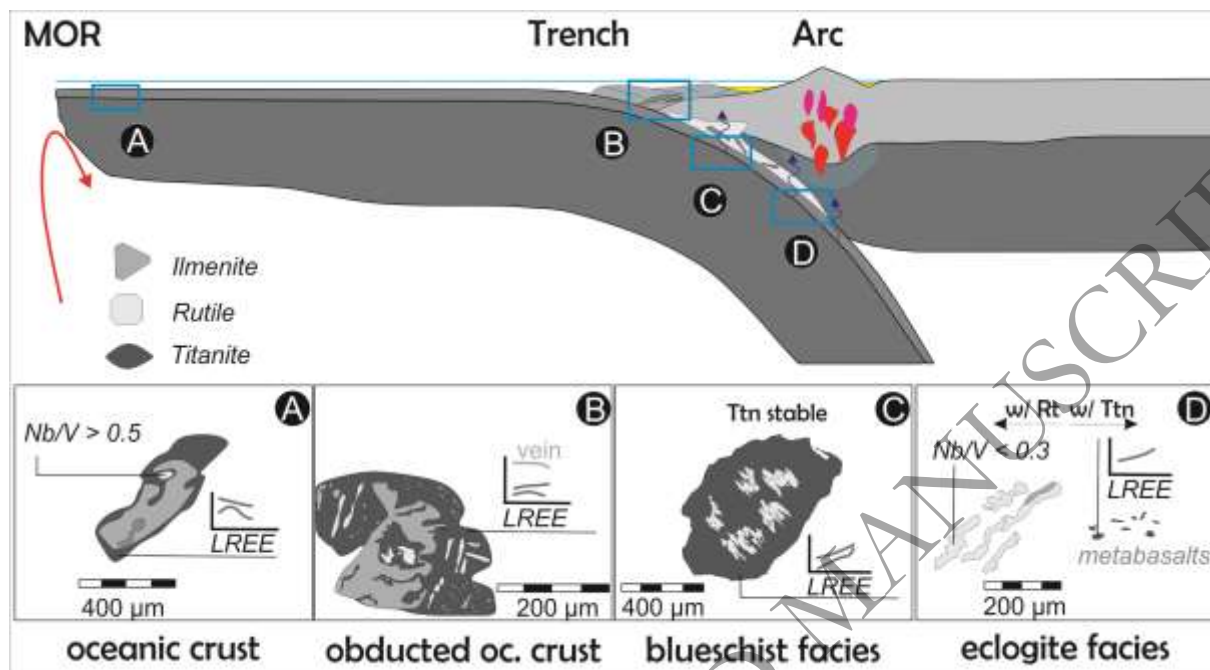
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1526 **Fig. 18.**

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