

Probing the hidden magmatic evolution of El Misti volcano (Peru) with the Pb isotope composition of fumaroles

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Ivan Vlastélic, Fredy Apaza, Pablo Masias, Marco Rivera, Jean-Luc Piro, et al.. Probing the hidden magmatic evolution of El Misti volcano (Peru) with the Pb isotope composition of fumaroles. Bulletin of Volcanology, 2022, 84 (2), pp.17. 10.1007/s00445-021-01521-9. hal-03538390

HAL Id: hal-03538390 https://uca.hal.science/hal-03538390

Submitted on 21 Jan 2022

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28 Abstract

29

This work proposes a new method to probe the hidden magmatic evolution of quiescent Andean volcanoes from the Pb isotope composition of gases. The method is based on an Assimilation-Fractional Crystallization-Degassing model linking the Pb isotope composition of gases with the SiO₂ content of their magmatic source. The model is applied to El Misti volcano that threatens Arequipa, the second most densely populated city of Peru.

- 35 Gas condensates and Pb-rich solid deposits (PbS, PbCl₂, PbSO₄) collected in 2018 in the bottom 36 of El Misti crater at 260 - 150°C fumaroles vents were used to reconstruct the mean 37 composition of degassing magmas (60.8-61.8 wt.% SiO₂). These compositions are slightly more 38 evolved than the lavas from the last AD 1440–1470 eruption, suggesting either the secular 39 differentiation of the main magma reservoir, or the contribution of more evolved magmas to 40 volcanic gases. On the other hand, the slight but significant difference between the 41 instantaneous composition recorded in gas condensates and the time-integrated composition 42 recorded in solid deposits points to the degassing of less evolved magmas over the last 43 decades. This trend is ascribed to a recent recharge of El Misti reservoir with hot mafic 44 magmas, in agreement with the evolution of fumarolic deposit mineralogy in the last half a century. The Pb isotope composition of gas appears to be a promising tool for probing the 45 46 hidden magmatic evolution of quiescent volcanoes where Assimilation-Fractional 47 Crystallization operates.
- 48

49 Keywords: El Misti volcano, Assimilation-Fractional Crystallization, magmatic gases,
50 fumaroles, sublimates, Pb isotopes

51

52 Acknowledgments:

53 We thank two anonymous reviewers for their constructive comments and William W 54 Chadwick for handling the manuscript. IV acknowledges OVI personnel for their welcome in 55 Arequipa. This research was financed by the Institut de Recherche pour le Développement, 56 the French Government Laboratory of Excellence initiative n° ANR-10-LABX-0006, the Region 57 Auvergne and the European Regional Development Fund. This is Laboratory of Excellence 58 ClerVolc contribution number XXX.

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- 60

61 Introduction

62

63 El Misti stratovolcano threatens the city of Arequipa, the second most densely populated city 64 of Peru (>1.1 million inhabitants, INEI, 2018) that is continuously growing at its foot (Fig. 1). 65 Over the last 50 000 years the volcano has shown a cyclic activity, with Plinian eruptions every 2000 to 4000 years and smaller to moderate explosive eruptions every 500 to 1500 years, on 66 67 average (Thouret et al., 2001). The lavas produced since 0.12Ma are andesitic to dacitic in 68 composition, with rare rhyolites and no marked evolution through time (Rivera et al., 2017). 69 This relative homogeneity contrasts with the mineral evidence for regular recharge and 70 mixing, which led Ruprecht and Wörner (2007) to suggest that magma recharge events are 71 more frequent in comparison to the eruption frequencies, and that El Misti reservoir evolves 72 most of the time cryptically. Since the last major eruption (A.D. 1440–1470), El Misti crater 73 has showed a persistent but fluctuating degassing activity (Thouret et al., 2001; Moussallam 74 et al., 2017). Over the last decades, the temperature of gases has varied between 125°C and 75 310°C (Hantke and Parodi 1966; Birnie and Hall, 1974; Thouret et al. 2001, and observations 76 from Observatorio Vulcanológico del Ingemmet (OVI) and Instituto Geofísico del Perú (IGP)). 77 The gases sampled in 2015 showed a dominant magmatic signature with little evidence for 78 interaction with rocks and hydrothermal fluids (Moussallam et al, 2017), which contrasts with 79 the dominance in 1967 of low-temperature minerals formed by gas-rock interactions (Birnie 80 and Hall, 1974). Moussallam et al. (2017) acknowledged that the current composition of the 81 magmatic source is unknown, and may be similar to either the rhyolitic or andesitic magmas 82 that interacted during the c. 2030 - 2000 years B.P. eruption (Thouret et al., 2001; Cobeñas et 83 al., 2012; Harpel et al., 2011). In the absence of direct measurement, the estimation of the 84 current state and evolution of the El Misti magma reservoir is challenging but important.

85 This study aims to reconstruct the composition of the El Misti magmatic reservoir from the Pb 86 isotopic composition of gas emitted in the summit crater. Lead isotopes are well suited for 87 this purpose because (1) lead is a volatile metal enriched in El Misti gases and fumarolic 88 deposits (Birnie and Hall, 1974), (2) the heavy Pb isotopes (²⁰⁸Pb, ²⁰⁷Pb, ²⁰⁶Pb and ²⁰⁴Pb) do not 89 fractionate measurably during magma degassing, due to their very small relative mass 90 differences (Vlastelic et al., 2013), so that the Pb isotopic composition of gas reliably records that of the degassing magma source, (3) the Pb isotopic composition of magma is tightly 91 92 related to the degree of magma differentiation through the Assimilation - Fractional

Crystallization (AFC) process that dominates the compositional evolution of El Misti magmas
(Mamani et al., 2008; Kiebala, 2008; Rivera et al., 2017). Lead isotopes are especially sensitive
to AFC as mantle melts and crustal basement display highly contrasted Pb contents and
isotopic compositions (Tilton and Barreiro, 1980; Barreiro and Clark, 1984). We have applied
this approach to gas condensates and fumaroles deposits sampled in the bottom of El Misti
crater in October 2018.

99

100 Geological setting

101

102 El Misti volcano (5822 m above sea level) is located to the north of the Andean Central 103 Volcanic Zone (CVZ) at 16°16' S and 71°22' W (Fig. 1a). It sits on the western flank of the 104 Western Cordillera, about 240 km east of the Peru-Chile trench. A detailed description of El 105 Misti's evolution is given in Thouret et al. (2001) and Rivera et al. (2017). Briefly, the composite 106 edifice comprises an older main stratovolcano (Misti 1, ca. 833–112 ka), on which grew three younger strato-cones (Misti 2, 3 and 4). Large volumes (4 - 6.5 km³) of dacitic pyroclastic 107 108 density currents (PDC) were emplaced between c. 50 000 and 40 000 and again between c. 13 109 700 and 11 300 years B.P., the last of which resulted in the formation of today's summit crater. 110 The most recent Plinian eruption (c. 2030 – 2000 years B.P.) produced 1.4 km³ of pumice-fall deposits that spread \geq 25 km from the vent and 0.01 km³ of PDC deposits (Harpel et al., 2011). 111 112 The mingling texture of rhyolitic and less evolved and esitic components in the pumices were 113 ascribed to the injection of an andesitic magma from 7-12 km depth into a shallower rhyolitic 114 reservoir (Tepley et al., 2013). This major event is most likely responsible for the formation of 115 today's 900 m-wide summit crater. The last major eruption (AD 1440–1470, VEI 2) produced 116 andesitic ash-fallout and led to the formation of a lava dome inside the present-day inner 117 crater. Since then, the El Misti crater has shown a persistent degassing activity, with periods 118 of increased fumarolic emissions and possible phreatic events in 1667, 1784-1787 and 1949 119 (Bullard, 1962; Chavéz Chavéz, 1992; Thouret et al., 2001). Less than 12 km separate El Misti 120 crater from the nearest Arequipa suburbs (Fig. 1b).

121

122 Samples and methods

124 We carried out field work in the El Misti crater in October 2018, at the end of the dry season. 125 After a survey of the lava dome, gas condensates and sublimates were collected at the hottest 126 fumarole vent (Fig. 1c) where a temperature of 260°C was measured using a Fluke 62 Max 127 infrared thermometer. Gas condensates were collected in a savillex teflon elbow emplaced in 128 the vent. Gas sublimates were subsequently sampled 30 cm underground at the lower surface 129 of a lava block that obstructed the vent. Sub-surface sublimates were also collected at another 130 vent with a temperature of 150°C (Fig. 1c). Samples were chemically treated and analysed at 131 Laboratoire Magmas et Volcans (LMV) in France. The sublimates were investigated by 132 Scanning Electron Microscopy (SEM, JSM 5910LV) operating in Backscattered Electron (BSE) 133 imaging mode. Sublimate mineralogy was determined from qualitative analyses of major 134 elements using Energy Dispersive Spectrometry (EDS). All chemical treatment of the samples 135 was done in LMV's clean room, within hoods equipped with HEPA filters, and using savillex 136 teflon vials and highly pure acids. One millilitre of the gas condensate was gently evaporated 137 at 60°C and re-dissolved in 1 ml of 0.5M HNO₃ - 0.05M HF. Representative chips of the 138 sublimates (0.2 to 100 mg in size) were fully dissolved in 6N HCl at 80°C. The condensate and 139 sublimate solutions were aliquoted for the analysis of trace element abundance and Pb 140 isotope composition.

141 Prior to ICPMS analysis, the sublimate aliquots were evaporated and re-dissolved in 0.5M HNO₃ - 0.05M HF, to reach dilution factors of 10⁵ (solution/sublimate mass ratio). All volumes 142 were precisely weighted with a precision better than 10⁻⁴g. The trace element abundances 143 144 were determined by quadrupole ICPMS (Agilent 7500), in plasma robust mode (1550 W), using 145 the reaction cell (He mode) to reduce interferences on masses ranging from 45 (Sc) to 77 (Se). 146 The signal was calibrated externally (every 4 samples) with a synthetic standard (Inorganic 147 Venture). The acid media was also measured every 4 samples to correct for background signal. 148 Repeated analysis of the analysed solutions yielded 2σ errors of < 5% for Ti, V, Cr, Mn, Co, Ni, 149 Cu, Rb, Sr, Zr, Nb, Sn, Ba, Pb and < 8 % for As, Cd, In, Sb, Tl and Bi.

Lead was separated using Eichrom Sr specific resin following the method of Deniel and Pin (2001). Procedural blanks were < 0.03 ng for Pb, which is insignificant compared to the amounts of Pb extracted from samples. Lead isotope compositions were measured on a Thermo Fisher Scientific Neptune Plus MC-ICPMS. Purified Pb was dissolved in 0.05M HNO₃ in order to obtain 150 ppb Pb solutions. The sample solutions were introduced in the mass spectrometer using a 100 μ L/min nebulizer coupled to a quartz cyclonic spray chamber, 156 yielding a total Pb signal between 18 and 20 volts. The Pb instrumental mass fractionation was 157 corrected during measurement by doping the Pb solutions with the Thallium (TI) NBS997 158 standard, applying a Pb/Tl ratio of 5. The NBS981 Pb standard was measured between every 159 two samples to correct Pb ratios for instrumental drift and the slight difference in mass 160 fractionation between Pb and Tl. Data were normalized to the NBS981 standard using the 161 values of Todt et al. (1996). Repeated analysis of the NBS981 standard every two samples yielded intra-session 2σ reproducibility better than 100, 120 and 140 ppm for 206 Pb/ 204 Pb, 162 ²⁰⁷Pb/²⁰⁴Pb and ²⁰⁸Pb/²⁰⁴Pb, respectively. 163

- 164
- 165 Results

166

The chemical and isotopic composition of the 2018 El Misti condensates and sublimates are 167 168 reported in Table 1. Images of the sublimates are shown in Fig. 2, and trace element contents 169 are plotted in Fig. 3. The 260°C gas condensate contains 177 ppb Pb, 37-38 ppb of Tl and Zn, 170 and 9-12 ppb of Bi, Cu, and Sb, with the latter concentrations being at the low range of volcanic 171 gases (Zelenski et al., 2021). The condensate also contains significant amounts of refractory 172 elements (244 ppb Ba, 45-46 ppb of Sr and Ti). The 260°C gas sublimates occur as a foam of 173 anhydrous alum (NaAl(SO₄)₂), several millimetres in thickness, and coated in many places with 174 galena (PbS) (Fig. 2). The five representative chips analysed for bulk composition contain 2 -175 10 wt. % Pb, 0.4 – 1 wt. % Ti, 4100 - 9500 ppm Tl, 5400 - 9400 ppm Zn, 2200 - 4600 ppm Bi, 176 2100 - 3600 ppm Cd, and 600 - 3200 ppm Sn. The 150°C gas sublimates occur as thin patches 177 of anglesite (PbSO₄) and cotunnite (PbCl₂), with minor anhydrite (CaSO₄), coating the lower 178 surface of a lava block (Fig. 2). The light yellow cotunnite and shiny black anglesite are barely 179 visible without the use of optical aids. The bulk analysis of four separates yielded 16.8 - 44.5 180 wt.% Pb, 142 - 1349 ppm Tl, 399 - 1689 ppm Zn, 81 - 114 ppm Bi, and 539 - 2257 ppm Cd. The 206 Pb/ 204 Pb, 207 Pb/ 204 Pb and 208 Pb/ 204 Pb ratios of the 260°C gas condensate are 17.724 ± 0.003, 181 182 15.592 ± 0.004 , and 38.628 ± 0.005 , respectively. The 260°C and 150°C gas sublimates display 183 significantly less radiogenic, but very homogeneous Pb isotope composition, with average 184 ²⁰⁶Pb/²⁰⁴Pb, ²⁰⁷Pb/²⁰⁴Pb and ²⁰⁸Pb/²⁰⁴Pb ratios of 17.676 ± 0.002, 15.584 ± 0.004, and 38.598 185 \pm 0.014, respectively. These ratios are within the ranges of previously analysed El Misti rocks 186 (Mamani et al., 2008).

187 The composition of the gas condensates and sublimates collected in 2018 show marked 188 differences when compared to those collected in July 1967 by Birnie and Hall (1974) (Fig. 3) 189 while the temperature of the fumaroles was only 100-125°C: (1) The Zn concentration of the 190 2018 gas condensate (37 ppb) is an order of magnitude lower than in the 1967 condensates 191 (240-740 ppb Zn); (2) The 2018 sulfates are dominantly Na-Al sulfate instead of Ca sulfates in 192 1967 (anhydrite and gypsum) in agreement with the hotter temperature of the 2018 gases. 193 This difference in sulfate mineralogy is recorded by Sr, which is much lower in 2018 sulfates 194 (Fig. 3); (3) Pb-rich phases such as galena, cotunnite, and anglesite are common in 2018, 195 whereas they were not detected in 1967 samples. Consequently, the 2018 Na-Al sulfates are 196 enriched in Pb and Tl compared with the 1967 anhydrite and gypsum with less than 2400 ppm 197 TI and 8300 ppm Pb.

198

Origin of the Pb isotopic composition of fumaroles samples

200

The Pb isotopic composition of fumarolic samples potentially records several processes, such entrainment of rock particulates, mixing between magmatic and hydrothermal gases, and variations of the magmatic source composition. The contribution of rock particulates to the Pb content of gas samples is assessed using the enrichment factor of Pb (Eq. 1):

205

206 EF_{Pb} = (Pb/X)_{gas samples}/(Pb/X)_{lava}

207

208 where X is a refractory rock-derived element (Symonds et al., 1987). Using titanium as rock-209 derived element, and the average composition of the lavas from "Misti 4" (12.3 ppm Pb and 5050 ppm Ti; Rivera et al., 2017), EF_{Pb} is 1.6 x 10^3 in the condensate, 0.8 – 9.4 x 10^3 in the 210 211 260°C sublimates, and 0.8 – 4.4 x 10⁵ in the 150°C sublimates. The lead fraction unsupported 212 by silicate input (1 - 1/EF_{Pb}) is 99.94 % in the condensate, >99.87 % in the 260°C sublimates 213 and >99.999 % in the 150°C sublimates (Fig. 4). Repeating this calculation with other rock-214 derived elements, such as Zr, Nb or Y, and "Misti 4" end-member rock compositions, yields in 215 all cases more than 99.7 % Pb unsupported by silicate input, confirming that the contribution 216 of rock particulates is extremely small. This tiny contribution does not account for the small 217 isotopic variations in gas condensates and sublimates, which do not correlate with EF_{Pb} 218 (Fig. 4). The dry composition (H_2O/SO_2 molar ratio of 32 ± 3) of El Misti gases also only requires a minimal contribution of the hydrothermal system (Moussallam et al., 2017). The contribution of low temperature hydrothermal vapours does not seem to affect the Pb isotope signature of the sublimates, which does not vary with vent temperature. The remaining and most likely possibility is that Pb isotope ratios of the fumarolic samples reflect the composition of magmatic gases.

224

225 Assimilation-Fractional Crystallization-Degassing model

226

227 The Pb isotopes of magmatic gases can be linked to the degree of crystallization of their 228 magmatic source through an Assimilation-Fractional Crystallization-Degassing (AFCD) model. This model takes advantage of the properties of heavy Pb isotopes which do not fractionate 229 230 during magma degassing, allowing us to reconstruct the Pb isotope composition of magmas 231 from gas composition. Then, we infer the degree of magma crystallization from the Pb isotope 232 composition of magmas using an Assimilation-Fractional Crystallization model. The two steps 233 and related assumptions of the AFCD model are described below, and solutions for one and 234 two magma reservoirs are provided.

235

236 Degassing

237

Lead is a moderately volatile element in magmatic systems, with a coefficient of emanation (i.e., fraction of Pb outgassed) in the range of 1-2% (Gauthier et al., 2000). Lead volatilizes dominantly as chloride and condenses mostly below 500°C (Symonds et al., 1987). The Pb isotope compositions of gas is related to that of the degassing magma by the coefficient of isotopic fractionation α (Eq. 2) :

243

244 $(^{206}Pb/^{204}Pb)_{gas} = \alpha_{64} (^{206}Pb/^{204}Pb)_{magma}$

245

in the case of the ²⁰⁶Pb/²⁰⁴Pb ratio. The theoretical kinetic isotope fractionation (α_{kin} = (m_{204}/m_{206})^{0.5} = 0.995) represents an extreme value of α_{64} , which is closer to unity in natural systems due to back condensation reactions at the gas - magma interface (Richter, 2004). Practically, isotopic fractionation of Pb during magma degassing is smaller than the precision of Pb isotope measurements (Vlastelic et al., 2013), and is not significant compared to the

251	variations occurring in lavas. In such situations, $lpha$ must be very close to unity for all Pb isotope
252	ratios, and the Pb isotope compositions of the fumarole samples must reliably reflect that of
253	the degassing magma. In the case of simultaneous degassing of andesite and rhyolite magma
254	reservoirs, the composition of the gas mixture is obtained by mass balance (Eq. 3):
255	
256	$(^{206}Pb/^{204}Pb)_{gas} = \Phi_{andesite} \times (^{206}Pb/^{204}Pb)_{andesite} + \Phi_{rhyolite} \times (^{206}Pb/^{204}Pb)_{rhyolite}$
257	
258	with Φ_{andesite} + Φ_{rhyolite} = 1
259	
260	where Φ is the fraction of the total Pb gas flux, which depends on the size and degassing rate
261	of each reservoir.
262	
263	Assimilation-Fractional Crystallization
264	
265	Mantle derived magmas ascending through continental arcs tend to melt and assimilate
266	crustal rocks with low melting temperature. Because wall-rock assimilation consumes heat,
267	while crystallization releases heat, the two processes are potentially coupled. However, the
268	ratio of assimilation to crystallization rate (referred to as r) largely varies in magmatic systems,
269	generally between 0.2 and 2.7. These variations depend on several parameters including the
270	initial temperature and melting point of the country rock, and the thermal potential of the
271	parental magma (Thompson et al., 2002). For instance, r is typically in the range of 1.0-1.5 for
272	an initial magma temperature of 1350-1150°C, compared to 0.6-0.2 for a temperature of
273	1150-900°C. The efficiency of assimilation is predicted to decrease (1) during differentiation
274	due to the decreasing liquidus temperature, and (2) with decreasing depth as the temperature
275	of host rocks decreases following the geothermal gradient. Both trends suggest that

assimilation takes places most pronouncedly during earlier, deeper magmatic stages.

A major outcome of AFC is that the isotopic variations of continental arc lavas often correlate with the degree of differentiation (De Paolo, 1981). The isotopic record of AFC is particularly clear along the CVZ where Pb isotope variations are controlled by the assimilation of Pb-rich rocks from the locally thick continental basement (e.g., Tilton and Barreiro, 1980; Wörner et al., 1992; Mamani et al. 2008; Kiebala, 2008; Sainlot et al., 2020a). At El Misti, the Assimilation-Fractional Crystallization process translates in a decrease of ²⁰⁶Pb/²⁰⁴Pb, ²⁰⁷Pb/²⁰⁴Pb and

²⁰⁸Pb/²⁰⁴Pb with increasing SiO₂, due to assimilation of the proterozoic Charcani gneiss 283 (Arequipa Massif) of the Western Cordillera basement with unradiogenic Pb ratios 284 (16.792<²⁰⁶Pb/²⁰⁴Pb< 17.088) (Tilton and Barreiro, 1980; Barreiro and Clark, 1984; Mamani et 285 al., 2008; Kiebala, 2008). Tilton and Barreiro (1980) pointed out that the granulitic gneiss must 286 287 have a lower crustal origin, but is probably assimilated by El Misti lavas at shallower depth. 288 We hypothesize that the frequent passages of magma in the lower crust favour the 289 disaggregation of crustal rocks into small xenoliths (Thompson et al., 2002). Entrained 290 xenoliths could partially dissolve during magma ascent in the upper crust due to superheating 291 above the liquidus (Brearley and Scarfe, 1986). How and where El Misti magmas are 292 contaminated remains uncertain, but contamination must be related to magma 293 crystallization, which dominantly occurs between 7 and 12 km depth (Tepley et al., 2013).

294 The AFC relation might have substantially varied during the growth of El Misti volcano due to 295 slight variations in magma path and thermal history, in particular following major collapses, 296 and heterogeneity of the basement. For this reason, we use the AFC relation defined over the 297 most recent Misti 4 stage (<11 ky) (Fig. 5), assuming that the composition of mantle derived 298 magmas and assimilated rocks, as well as the r value, do not vary at short time scales. This 299 relation is defined based on four samples analysed for Pb isotopes by Mamani et al. (2008), 300 including 6 ka old scoria from pyroclastic density current deposits (MIS-00-20), two pumices from fall deposits (quebrada San Lazaro) and one pumice (MIS-02-10a) from the 2030-2000 301 302 years B.P. eruption. The SiO₂ contents of these samples, between 58.7 and 63.2 wt.%, 303 encompass the range of variations of El Misti rocks with the exception of rhyolites. The 304 samples show a Pb isotope range (17.3965<²⁰⁶Pb/²⁰⁴Pb<17.8442) that brackets the 305 composition of the fumarole samples (17.6740<²⁰⁶Pb/²⁰⁴Pb<17.7243), and are thus 306 appropriate for calibrating the AFC relationship. In addition, these samples do not contain 307 large minerals, with potential Pb isotope disequilibria relative to the matrix, so it can 308 reasonably be assumed that their bulk composition reflects that of the degassing melt. The 309 black ash fall sample (MIS-02-06), with a proposed age of 11 ky (Kiebala, 2008) is not included 310 because (1) this sample pertains to Misti 3 (Rivera et al., 2017), and (2) the juvenile origin of 311 the ashes is uncertain. For the purpose of modelling, we preferentially use the ²⁰⁶Pb/²⁰⁴Pb and ²⁰⁸Pb/²⁰⁶Pb ratios, which show well-defined AFC trajectories owing to large variations in lavas 312 and precise measurements. The ²⁰⁷Pb/²⁰⁴Pb and ²⁰⁸Pb/²⁰⁴Pb ratios, with smaller variations and 313

larger errors are not used. The relationships between SiO₂ content and Pb isotope ratios in
 lavas from Misti 4 stage are best fitted (R²> 0.988) by the second-degree polynomial functions:
 316

317 SiO₂ (wt%) magma = -26.015 x (206 Pb/ 204 Pb) magma² + 906.56 x (206 Pb/ 204 Pb) magma -7834.7 (Eq. 4)

318

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319 SiO<sub>2</sub> (wt%) magma = -3163.95 x (<sup>208</sup>Pb/<sup>206</sup>Pb) magma<sup>2</sup> + 13949.16 x (<sup>208</sup>Pb/<sup>206</sup>Pb) magma
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- 320 -15311.5 (Eq. 5)
- 321

322 These relations defined over a SiO₂ range of 58.7-63.2 wt.% probably do not hold for rhyolites 323 for two reasons: (1) if the final crystallisation of rhyolites occurs mostly within the volcanic 324 edifice (~3 km depth), as suggested by Tepley et al. (2013), it is expected that rhyolites 325 assimilate older Misti rocks and not the Charcani gneiss (Arequipa Massif). In keeping with this 326 idea, Barreiro and Clark (1984) noted that the most contaminated El Misti rhyolite has a 327 ²⁰⁶Pb/²⁰⁴Pb ratio of 17.38, barely lower than the dacitic pumice sample MIS-02-10a (17.3965). 328 (2) Crustal assimilation markedly decreases during late stage crystallisation (Thompson et al., 329 2002). Thus, the Pb isotope composition of gases is probably less sensitive to late stage 330 crystallisation of magmas at shallow depth, which may constitute the main limitation of the 331 model.

332

333 Reconstructed mean composition of degassing magmas

334

335 The model provides, as a general solution, the mean composition of degassing magmas. The 336 mean composition does not constrain the composition and distribution of heterogeneities (i.e. 337 multiple magma reservoirs) within the system. The mean composition of degassing magmas 338 can be reconstructed by combining degassing (Eg. 2) and AFC equations (Eqs. 4 and 5). From 339 Eqs. 2 and 4, we estimate the SiO₂ content of degassing magmas at 60.81 wt.% for the gas 340 condensate and at 61.50 wt% for the average of the gas sublimates. Using Eq. 5 based on the 341 ²⁰⁸Pb/²⁰⁶Pb ratio instead of Eq. 4 gives slightly higher SiO₂ concentrations of 61.19 wt.% and 342 61.81 wt%, respectively. In both cases, the gas sublimates record a more evolved composition 343 (+0.62 to +0.69 wt.% higher SiO₂) compared to the gas condensate. The 100 ppm analytical 344 uncertainty on the Pb isotope ratios of the fumaroles samples (Table 1) translates into a 345 relative uncertainty of less than 0.03 wt.% on the estimated SiO₂ values. The absolute error on the SiO₂ estimates still remains controlled by the precision of the SiO₂ concentration data
used in Eq. 4-5, which does not exceed 0.3 wt.%. In Fig. 6, the estimated SiO₂ concentrations
of the degassing magmas are compared to the long-term evolution of the SiO₂ content of El
Misti lavas. The results of the model can be directly used to trace the evolution through time
of a single large magma reservoir (Fig. 6a).

351

352 Degassing of two reservoirs of known composition

353

354 In the model of Tepley et al. (2013), an additional smaller rhyolitic magma reservoir transiently 355 forms in the shallow crust (ca. 3km) every 2000 to 4000 years. It is not known whether such a 356 rhyolite reservoir has existed since the last 2030-2000 years B.P. Plinian eruption. If it exists, 357 one can envision that both deep and esite and shallow rhyolite magmas contribute to the gases 358 escaping from El Misti crater. In this case, specific model solutions can be obtained considering 359 that the Pb flux from each reservoir (Φ) might also contribute to the Pb isotope composition 360 of the gas mixture at the surface according to Eq. 3, in addition to the contrasted compositions 361 of the magma sources. In an end-member situation, the compositions of the magma reservoirs 362 are constant and the composition of the gas is entirely controlled by varying Φ (Fig. 6b). The 363 model then allows the relative contributions of two reservoirs of known composition to be 364 quantified. The ²⁰⁶Pb/²⁰⁴Pb ratio of the andesite reservoir can be assumed to be the same as 365 the least evolved scoria MIS-00-20 (17.8442). In the absence of published Pb data on rhyolites, 366 the Pb composition of the rhyolite reservoir can be assumed to be the same as the dacitic pumice MIS-02-10a (²⁰⁶Pb/²⁰⁴Pb =17.3965) because, as argued above, late-stage 367 368 differentiation of magmas within the edifice probably does not markedly modify Pb isotope composition. With these compositions, solving Eq. 3 yields $\Phi_{andesite}$ =73% and $\Phi_{rhyolite}$ =27% for 369 the gas condensate, and $\Phi_{andesite}$ =62% and $\Phi_{rhyolite}$ =38% for the gas sublimates. In this model, 370 371 the contribution of gaseous Pb from the rhyolite reservoir is about 10% higher in gas 372 sublimates than in gas condensates.

373

374 Hidden evolution of the El Misti magma reservoir

375

376 Evolution since the last eruption (AD 1440–1470)

378 The modelled composition (60.81- 61.81 wt.% SiO₂) of magmas currently degassing at El Misti 379 volcano is similar to the average composition of El Misti magmas (61.0 ± 3 wt.% SiO₂, n= 116, 380 Rivera et al., 2017). The model results preclude highly evolved magmas, such as those sampled 381 during the last Plinian eruption with >63 wt.% SiO₂ (Thouret et al., 2001; Tepley et al., 2013), 382 being the dominant source of the 2018 gases. Yet, the modelled 2018 composition is slightly 383 more evolved than that of the most recent lavas, in particular those forming the conduit plug 384 with a SiO₂ content generally below 58.9 wt.% (Rivera et al., 2017) (Fig. 6). There are two 385 possible interpretations of this shift depending on the interpretation chosen for the structure 386 of the El Misti plumbing system. In the hypothesis of a single, large magma reservoir 387 undergoing cycles of recharge and differentiation (Ruprecht and Wörner, 2007), the SiO₂ 388 increase inferred here from Pb isotopes indicates that magma differentiation has been on 389 average more important than magma recharge over the last few centuries (Fig. 6a). On the 390 other hand, in the two-reservoir model of Tepley et al. (2013), one can assume that the 391 andesite and rhyolite magmas preserve more or less constant compositions at the secular 392 scale (Fig. 6b). The evolved composition of currently degassing magmas (relative to the lavas 393 from the last eruption) can be explained by the contribution of gases from the shallow rhyolitic 394 magmas to the main gas flux.

395

396 Current evolution

397

398 The small Pb isotope difference between gas sublimates and gas condensate cannot reflect 399 the simultaneous input of different gases at the same vent. Instead, this difference can be 400 related to the fact that the condensate provides an instantaneous composition (at the time of 401 sampling in October 2018) whereas the solid deposit records a composition averaged over the 402 time of formation of the deposit. This time must be less than 48 years because the Na-Al 403 sulfates and the Pb-rich minerals were not already formed in July 1967 (Birnie and Hall, 1974). 404 The difference between the instantaneous (60.81- 61.19 wt.% SiO₂) and time-integrated 405 (61.50- 61.81 wt.% SiO₂) SiO₂ estimates thus records the degassing of less evolved magma 406 through time. A straightforward implication from the single magma reservoir hypothesis is 407 that the magmatic reservoir is currently undergoing recharge by hot mafic magmas (Fig. 6a). 408 In the two-reservoir model of Tepley et al. (2013), the current evolution of gas composition 409 could instead reflect the decreasing contribution of gases from the shallow rhyolitic reservoir,410 possibility due to its progressive solidification (Fig. 6b).

411 The recharge hypothesis is independently supported by the mineralogy of the gas sublimates. 412 The Pb minerals (PbS, PbSO₄ and PbCl₂) occurring in the El Misti 2018 fumaroles deposits are 413 typically formed by the condensation of high-temperature magmatic gas (Symonds et al., 414 1987; Taran et al., 2000; Africano et al., 2002). These gases can be similar to those sampled in 415 2015, with an equilibrium temperature of 532 °C and fO₂ equivalent to Δ QFM = +2.8 416 (Moussallam et al., 2017). In these conditions, galena forms first and is replaced by cotunnite 417 and anglesite as the temperature decreases and fO₂ increases. At El Misti, the occurrence of 418 PbSO₄ and PbCl₂ below 300°C requires little or no mixing of magmatic gas with air (Africano et 419 al., 2002). The occurrence of galena at the 260°C vent and cotunnite-anglesite at the 150°C 420 vent is thus consistent with closed-system condensation of the same high-temperature 421 parental gas that probably feeds the whole fumarolic field. The 2018 mineral assemblage 422 contrasts with the 1947 low temperature mineral suite (sulfur, anhydrite, gypsum and 423 ralstonite) that results from extensive gas rock interaction (Birnie and Hall, 1974). Part of this 424 difference might reflect sampling bias because Birnie and Hall (1974) did not sample 425 underground deposits whereas we did not sample sulfur and other minerals from low 426 temperature vents. However, the absence of high-temperature minerals in all five samples 427 collected in 1967 suggests overall colder degassing conditions in agreement with the lower 428 measured temperature (100-125°C). Fumarole deposits grow very slowly, on the order of 429 µm/day (Johnson and Burnett, 1993; Sainlot et al., 2020b) and average out short-term 430 fluctuations of degassing conditions. The difference in mineral assemblage between 1967 and 431 2018 is a robust observation that indicates a recent increase in gas temperature. The 432 important increase of gases temperature (>100°C) cannot reflect a comparable increase of 433 magma temperature but might reflect both a shallower and slightly hotter magma source, and overall conditions severely limiting gas scrubbing, H₂SO₄ condensation and gas-rock 434 435 interaction. In summary, both sublimate mineralogy and Pb isotope composition of gas 436 support the idea of a recent hidden recharge of the shallow reservoir with hotter mafic 437 magmas. However, the inferred magma recharge and related temperature increase have not 438 been detected since the onset of thermal monitoring from space of the El Misti lava dome in 439 2000 (Moussallam et al., 2017), suggesting the recharge occurred before 2000. The recharge

perhaps could have been related to the last high fumarolic activity observed from Arequipa in
1984–1985 (Thouret et al., 2001).

442

443 Implications for hazards and volcanic monitoring

444

445 The activity of El Misti volcano is monitored from the height of the gas plume, edifice 446 deformation, thermal emission, volcano-tectonic earthquakes, and the composition of 447 geothermal waters. Collectively, the monitored parameters point to rather stable activity of 448 the volcano at the decadal time scale. Moussallam et al. (2017) emphasized the need to also 449 monitor major species of gas (H₂O, H₂, SO₂, H₂S), which allow for an estimation of gas-melt 450 equilibrium temperature and oxygen fugacity. Here we suggest that the Pb isotopic 451 composition of fumarolic minerals is less sensitive to short term change of degassing 452 conditions and represents perhaps the most direct probe of recent magma composition. 453 Monitoring the Pb isotope composition of El Misti gases would provide a window on the 454 hidden cycles of magma recharge and differentiation, and would allow us to detect any 455 departure from the steady-state regime that could potentially lead to an eruption.

456

457 **Conclusions**

458

We developed a new method to probe the hidden magmatic evolution of quiescent Andean volcanoes from the Pb isotope composition of gas emissions. The method is based on the properties of Pb isotopes that vary systematically during Assimilation - Fractional Crystallization (AFC) of magmas but do not fractionate during degassing. The method is applied to El Misti volcano that threatens the nearby city of Arequipa, the second most densely populated city of Peru.

Using this approach, we show that magmas currently degassing at El Misti volcano are on average more evolved than the centuries old lavas forming the conduit plug. This difference may reflect the secular differentiation of a single, large magma reservoir, or, alternatively, the contribution of a more evolved magma reservoir (possibly rhyolitic) to volcanic emanations. Conversely, the slight difference between the instantaneous composition recorded in gas condensate and the time-integrated composition recorded in solid deposits points to the degassing of less evolved magmas over the last decades. This trend indicates that El Misti

magma reservoir may be currently undergoing recharge by hotter more mafic magmas,
consistent with the lack of Pb-rich sublimates and the predominance of low temperature
minerals in fumarole deposits collected half a century ago. This work suggests that the Pb
isotope composition of gas emissions is a promising tool for probing the hidden magmatic
evolution of quiescent volcanoes where AFC is the dominant process.

- 477 Figure captions
- 478

Fig. 1 Location maps. (a) Map of South America showing the location of El Misti Volcano in the Central Volcanic Zone (CVZ). Northern (NVZ) and southern (SVZ) volcanic zones are indicated. (b) Satellite map of the El Misti area showing the location of Arequipa city. White line shows the approximate boundary of Arequipa suburbs in 2021. Source: Google, Images©2021 TerraMetrics. (3) Photo of El Misti lava plug inside the crater taken in October 2018. The locations and temperatures of the sampled fumaroles are indicated by arrows. The 120 m wide lava dome is located at the bottom of the 200 m deep inner crater.

486

Fig. 2 Images of the sublimates. (a-c) 150°C sublimates: (a) Binocular loupe photo of a grain
covered with PbCl₂ and PbSO₄. (b) Scanning Electron Microscopy (SEM) image of the same
grain. (c) SEM image of a grain with CaSO₄. (d-f) 260°C sublimates: (d) Macroscopic photo of
NaAl(SO₄)₂ deposit. (e) and (f): SEM images showing PbS coating NaAl(SO₄)₂ foam. All SEM
images were acquired at LMV using a JEOL JSM 5910LV microscope operating in Backscattered
Electron (BSE) imaging mode.

493

494 Fig. 3 Trace element concentrations of El Misti fumarolic minerals and gas condensates. Data
495 are shown for samples collected in 2018 (this work) and 1967 (Birnie and Hall, 1974). Trace
496 elements are sorted from left to right according to their decreasing abundances in the 260°C
497 condensate. Average concentrations are plotted for each sample type.

498

Fig. 4 ²⁰⁶Pb/²⁰⁴Pb plotted against the Pb fraction unsupported by silicate input. The silicate unsupported Pb fraction (i.e. the fraction of Pb originating from the gas) is equal to 1 - 1/EF_{Pb}, where EF_{Pb} is the Pb enrichment factor (see text). Here, EF_{Pb} is calculated relative to titanium, assuming that all titanium in gas samples originates from rocks. The absence of correlation in this diagram argues against a control of rock particulates on the small Pb isotope variations.

Fig. 5 Plot of SiO₂ versus ²⁰⁶Pb/²⁰⁴Pb (a) and ²⁰⁸Pb/²⁰⁶Pb (b) for El Misti lavas younger than 11
ky (stage 4). This relation reflects the assimilation of unradiogenic Pb from the proterozoic
Charcani gneiss (Arequipa Massif) during fractional crystallization of El Misti magmas. Samples
include 6 ka old scoria bombs (MIS-00-20), two pumices from fall (quebrada San Lazaro) and

509 one dacitic pumice (MIS-02-10a) from the 2000 years B.P. eruption. Data are from Mamani et 510 al. (2008) (Pb isotopes) and Rivera et al. (2017) (SiO₂ content). Errors bars are for analytical 511 uncertainties. Best fit curves are shown. The single Pb data (²⁰⁶Pb/²⁰⁴Pb of 17.38) reported for 512 El Misti rhyolites (Barreiro and Clark, 1984) suggests that rhyolites are not more contaminated 513 than dacites (vertical arrows), consistent with the final differentiation of rhyolites within El 514 Misti edifice (Tepley et al., 2013).

515

516 Fig. 6 SiO₂ content of El Misti lavas and tephra plotted versus time. Data are from Thouret et 517 al. (2001), Rivera et al. (2017) and Tepley et al. (2013). The star and shaded rectangle show 518 the instantaneous and time-integrated SiO₂ contents of degassing magmas inferred from 2018 519 gas condensates and sublimates, respectively. (A) Single reservoir model (Ruprecht and 520 Wörner 2007). In this model, arrows indicate the inferred evolution of the magma reservoir 521 through time. Arrow 1 links the composition of the last erupted lavas (the conduit plug) with 522 the inferred composition of the currently degassing magma, pointing to a secular 523 differentiation trend. Arrow 2 links the time-integrated and instantaneous modelled SiO₂ 524 contents. This trend suggests that the El Misti magma reservoir is currently undergoing 525 recharge by hot mafic magmas, in agreement with the evolution of fumaroles minerals during 526 the last half a century. (B) Two reservoir model (Tepley et al., 2013). In this model, emanations 527 from El Misti crater result from mixing gases escaping from the andesite and rhyolite 528 reservoirs (horizontal dashed arrows). In an end-member case, both reservoirs have more or 529 less constant compositions through time (vertical arrows), and the composition of gases 530 reflects varying contributions of the two magma reservoirs. In this model, the contribution of 531 the rhyolite reservoir to the Pb gas flux is estimated at 38% over the last decades, and 27% 532 currently (see text).

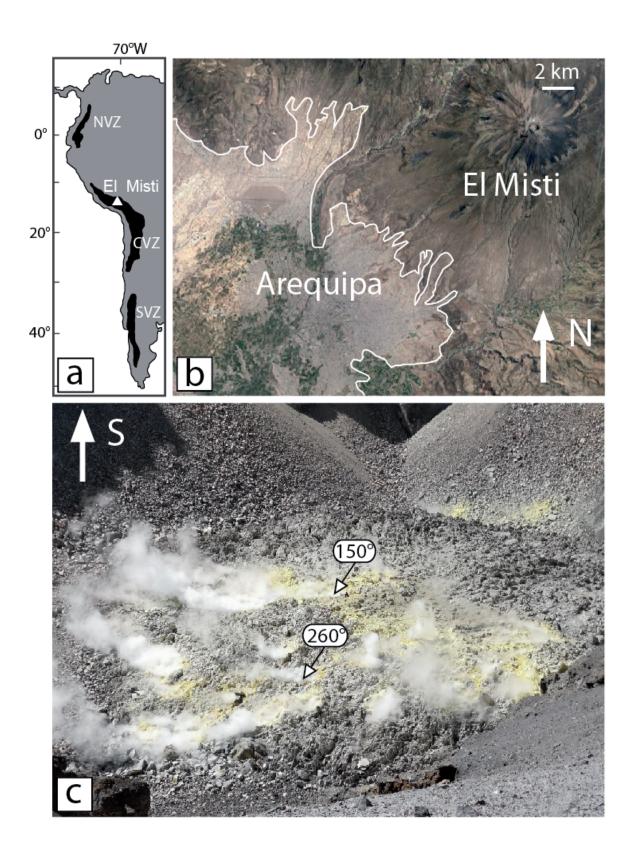
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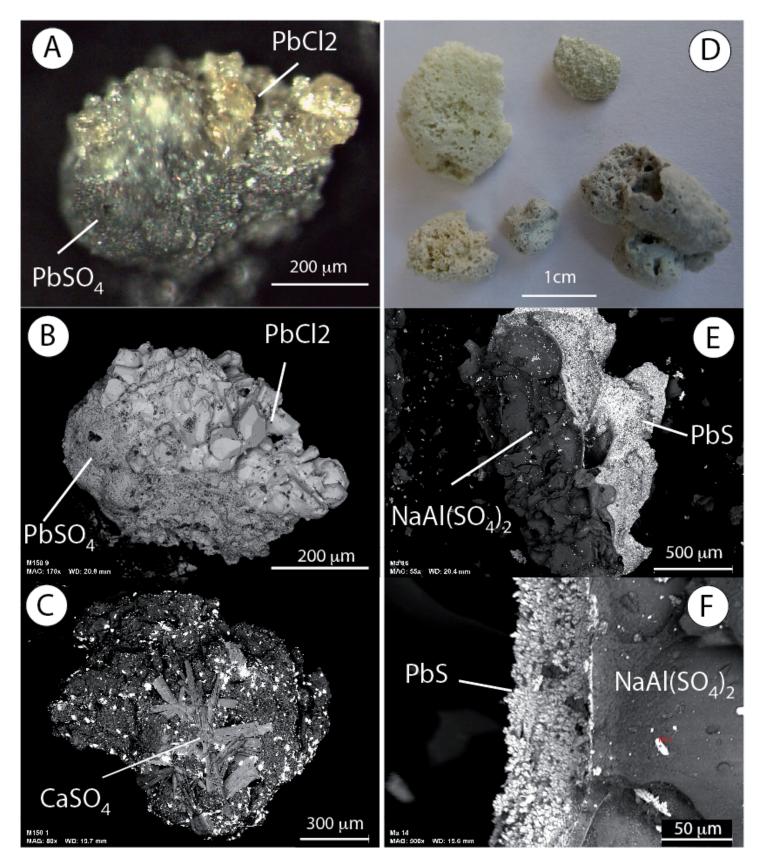
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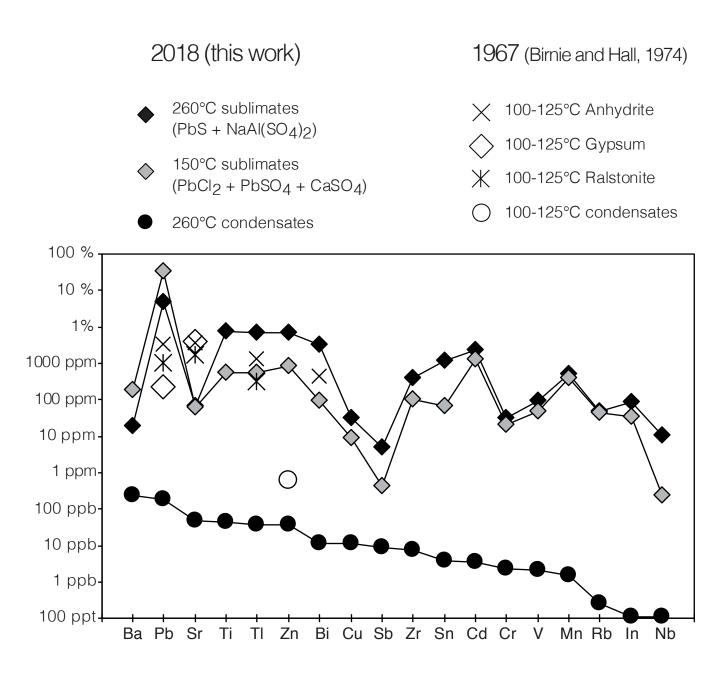
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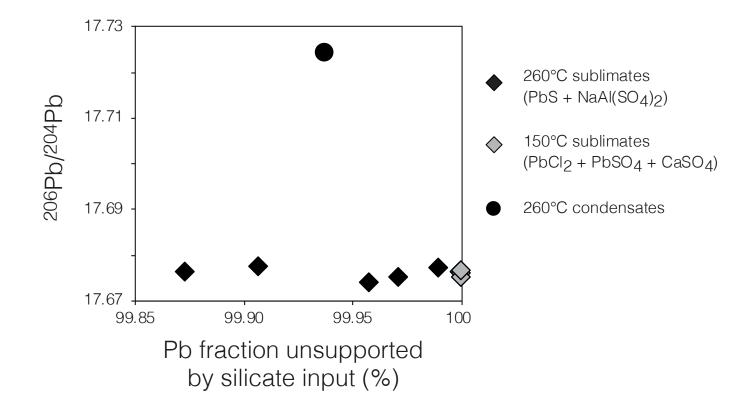


150°C sublimates

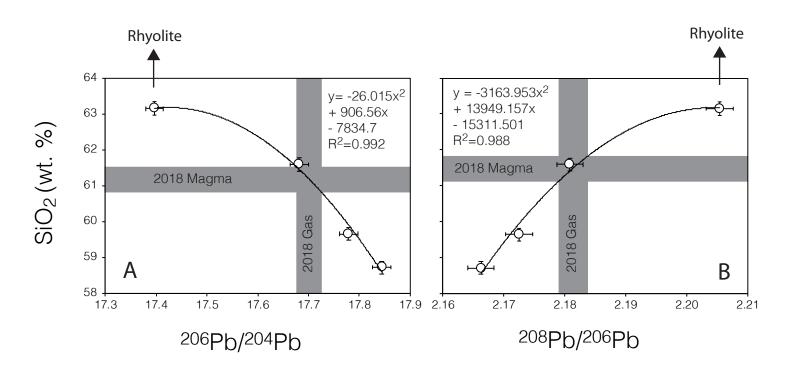
260°C sublimates







AFC Misti 4



Single reservoir



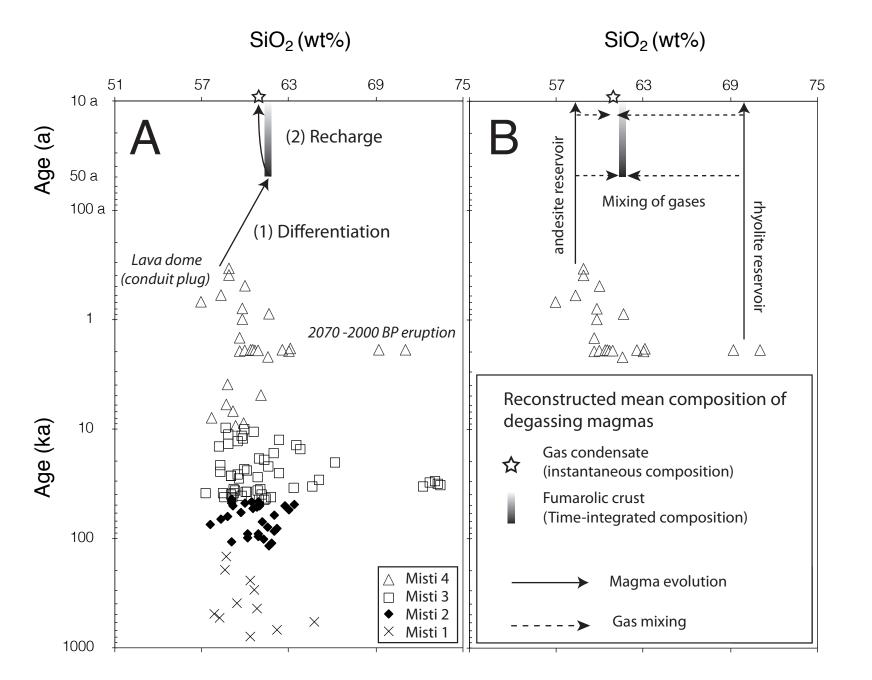


Table 1: Chemical and isotopic composition of 2018 El Misti condensates and sublimates

Sample type Temperature Mineralogy	Gas condensate 260°C	Gas sublimates 260°C PbS + NaAl(SO4)2					Gas sublimates 150℃ PbCl2 + PbSO4 + CaSO4			
		#a :	#b #d	c #	d #	e	#1 #	3 #	4 #	5
	ppb	wt%	wt%	wt%	wt%	wt%	wt%	wt%	wt%	wt%
Pb	177	3.80	4.90	2.01	2.82	9.86	16.8	36.3	32.4	44.5
	ppb	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm
Ti	45.0	6593	5734	10338	10649	4228	822	410	522	405
V	2.04	96	90	105	104	64	73	49	50	24
Cr	2.27	33	30	39	38	22	40	15	19	7
Mn	1.52	589	559	454	465	578	344	646	500	112
Cu	11.4	51	44	26	25	18	5	12	10	10
Zn	37.0	5393	6094	6842	7008	9363	1689	645	589	399
Rb	0.25	63	66	42	42	36	38	43	78	12
Sr	45.7	36	37	10	9	234	10	108	51	84
Zr	7.65	233	195	605	601	305	134	121	96	46
Nb	0.11	19	17	6	6	4	0	0	0	0
Cd	3.35	2141	2277	2158	2158	3306	2257	1304	1016	539
In	0.11	45	55	70	74	173	77	29	23	13
Sn	3.87	896	814	637	606	3160	138	51	51	35
Sb	8.74	9	9	2	2	3	0.2	0.6	0.4	0.6
Ва	244	34	33	6	6	18	39	26	411	260
TI	38.3	5283	9479	7731	7868	4097	1349	234	142	514
Bi	11.7	2314	2210	3182	3441	4617	81	115	82	113
206Pb/204Pb	17.7253	17.6743	17.6751	17.6764	17.6776	17.6778	17.6764	17.6762	17.6753	17.6765
duplicate	17.7234 r	17.6738 d				17.6770 d				
207Pb/204Pb	15.5935	15.5843	15.5851	15.5853	15.5866	15.5866	15.5824	15.5820	15.5818	15.5823
duplicate	15.5910 r	15.5838 d				15.5858 d				
208Pb/204Pb	38.6298	38.5999	38.6033	38.6048	38.6073	38.6076	38.5929	38.5906	38.5898	38.5923
duplicate	38.6264 r	38.5985 d				38.6049 d				

Pb isotope ratios are relative to the NBS 981 values of Todt el al. (1996)

r: duplicate run

d: duplicate chemistry